

23 **ABSTRACT**

24 U-Pb geochronology was conducted by Laser Ablation-Inductively Coupled Plasma Mass
25 Spectrometry (LA-ICPMS) on 7,175 detrital zircon grains from twenty-nine samples from the
26 Coconino Sandstone, Moenkopi Formation, and Chinle Formation. These samples were
27 recovered from ~520 m of drill core that was acquired during the Colorado Plateau Coring
28 Project (CPCP), located in Petrified Forest National Park (Arizona).

29 A sample from the lower Permian Coconino Sandstone yields a broad distribution of
30 Proterozoic and Paleozoic ages that are consistent with derivation from the Appalachian and
31 Ouachita orogens, with little input from local basement or Ancestral Rocky Mountain sources.
32 Four samples from the Holbrook Member of the Moenkopi Formation yield a different set of
33 Precambrian and Paleozoic age groups, indicating derivation from the Ouachita orogen, the
34 East Mexico Arc, and the Permo-Triassic arc built along the Cordilleran margin.

35 Twenty-three samples from the Chinle Formation contain variable proportions of Proterozoic
36 and Paleozoic zircon grains, but are dominated by Late Triassic grains. LA-ICPMS ages of these
37 grains belong to five main groups that correspond to the Mesa Redondo Member, Blue Mesa
38 Member and lower part of the Sonsela Member, upper part of the Sonsela Member, middle
39 part of the Petrified Forest Member, and upper part of the Petrified Forest Member. The ages
40 of pre-Triassic grains also correspond to these chronostratigraphic units, and are interpreted to
41 reflect varying contributions from the Appalachian orogen to the east, Ouachita orogen to the
42 southeast, Precambrian basement exposed in the Ancestral Mogollon Highlands to the south,
43 East Mexico arc, and Permian-Triassic arc built along the southern Cordilleran margin. Triassic
44 grains in each chronostratigraphic unit also have distinct U and Th concentrations, which are
45 interpreted to reflect temporal changes in the chemistry of arc magmatism.

46 Comparison of our LA-ICPMS ages with available CA-TIMS ages and new magnetostratigraphic
47 data provides new insights into the depositional history of the Chinle Formation, as well as
48 methods utilized to determine depositional ages of fluvial strata. For parts of the Chinle
49 Formation that are dominated by fine-grained clastic strata (e.g. mudstone and siltstone), such
50 as the Blue Mesa Member and Petrified Forest Member, all three chronometers agree (to
51 within ~1 m.y.), and robust depositional chronologies have been determined. In contrast, for
52 stratigraphic intervals dominated by coarse-grained clastic strata (e.g., sandstone), such as
53 most of the Sonsela Member, the three chronologic records disagree due to recycling of older
54 zircon grains and variable dilution of syn-depositional-age grains. This results in LA-ICPMS ages
55 that significantly pre-date deposition, and CA-TIMS ages that range between the other two
56 chronometers. These complications challenge attempts to establish a well-defined
57 chronostratigraphic age model for the Chinle Formation

58 **1. INTRODUCTION**

59 Triassic strata of the Colorado Plateau and environs provide rich and geographically extensive
60 records of environmental and biotic change during a critical period of Earth history, as well as
61 the transition from passive- to convergent-margin tectonism along the North American
62 Cordillera (e.g., Parker and Martz, 2011; Olsen et al., 2011). As demonstrated by Riggs et al.
63 (1996, 2003, 2012, 2013, 2016), Dickinson and Gehrels (2008), Irmis et al. (2011), Ramezani et
64 al. (2011, 2014), Atchley et al. (2013), Nordt et al. (2015), Kent et al. (2018, 2019), Olsen et al.
65 (2018, 2019), Marsh et al. (2019), and Rasmussen et al. (2020), Chinle Formation strata have
66 the potential to record the timing of these changes in great detail given their several-hundred-
67 meter thickness, abundance of near-depositional-age zircon grains, and recoverable
68 paleomagnetic reversal stratigraphy.

69 In an effort to further develop this record, ~520 m of continuous core was collected from
70 Triassic and underlying Permian strata at Petrified Forest National Park (PEFO), which is located
71 on the southern Colorado Plateau of northern Arizona (Fig. 1; (35.085933° N, 109.795500° W,
72 WGS84 datum). The objectives and primary findings of this project have been described by
73 Olsen et al. (2018, 2019), Kent et al. (2018, 2019), and Rasmussen et al. (2020), and numerous
74 related studies are currently in progress. This contribution to the project reports U-Pb
75 geochronologic analyses of detrital zircon grains that were extracted from twenty-nine samples
76 from this core (CPCP-PFNP13-1A). Analyses were conducted by laser ablation-inductively
77 coupled mass spectrometry (LA-ICPMS), with between 36 and 490 grains analyzed per sample
78 (total of 7,175 analyses). Grains were chosen for analysis by random selection in an effort to
79 provide unbiased information about provenance. Fortunately, a significant number of near-
80 depositional-age grains were recovered from many samples in the Chinle Formation, which
81 provides opportunities to also determine robust maximum depositional ages. This report
82 explores variations in both provenance and maximum depositional age of strata intersected in
83 the CPCP-PFNP13-1A core, and the implications for Permian-Triassic environmental and biotic
84 transformations and the tectonic evolution of southwestern North America.

85 **2. STRATA ENCOUNTERED IN THE PETRIFIED FOREST NATIONAL PARK DRILL CORE**

86 The lowest stratigraphic horizon encountered consists of quartz arenite belonging to the
87 Coconino Sandstone (Fig. 2). This unit belongs to regionally extensive erg deposits of early
88 Permian (Leonardian) age (Blakey et al., 1988; Lawton et al., 2015; Dickinson, 2018).

89 Overlying strata of the Coconino Sandstone are tabular, thin to thick-bedded, reddish
90 mudstone, siltstone, and sandstone layers of the Lower-Middle Triassic Moenkopi Formation. In
91 the PEFO region, the Moenkopi Formation consists of thin-bedded reddish siltstone with
92 interlayered sandstone and mudstone. Lower, finer-grained strata are assigned to the Wupatki
93 Member and Moqui Member, and upper sandstone-rich horizons dominate the Holbrook
94 Member. The base is a regional unconformity, the TR-1 unconformity of Pipiringos and
95 O'Sullivan (1978), along which strata of the lower Permian Toroweap Formation and Kaibab

96 Formation have been removed. Strata of the Moenkopi Formation are interpreted to have
97 accumulated on a northwest-sloping coastal plain, with thinner fluvial strata to the southeast
98 and thicker marginal marine strata to the northwest (Dickinson, 2018). The Moenkopi
99 Formation basin was bounded by residual uplifts of the Ancestral Rocky Mountains to the
100 northeast and highlands of the Ouachita orogen to the southeast. Highlands developed within
101 early phases of the Cordilleran magmatic arc may have existed to the southwest.

102 Strata of the Moenkopi Formation are overlain unconformably [Tr-3 unconformity of Pippingos
103 and O'Sullivan (1978)] by the Chinle Formation (Fig. 2). The transition is marked in most areas
104 by the Shinarump Conglomerate, which consists of cobbles of chert, quartzite, limestone and
105 subordinate felsic volcanic rocks. Riggs et al. (2012) have determined U-Pb ages of 232-224 Ma
106 on volcanic cobbles in the Shinarump Conglomerate. The Shinarump Conglomerate is
107 interpreted to correlate with finer-grained strata of the Mesa Redondo Member (Irmis et al.,
108 2011; Martz et al., 2012, 2017; Riggs et al., 2016). Strata of the Shinarump Conglomerate and
109 Mesa Redondo Member are interpreted to have accumulated in paleovalleys that were carved
110 into underlying strata. Strikingly variegated, strongly pedogenically modified, red, purple, and
111 yellow strata in the core are assigned to the Mesa Redondo Member given the lack of
112 conglomerate. Strata of the Mesa Redondo Member in outcrop have yielded U-Pb (zircon) ages
113 of ~227.6 Ma (Atchley et al., 2013) and ~225.2 Ma (Ramezani et al., 2011).

114 Gradationally overlying the Mesa Redondo Member are strata of the Blue Mesa Member,
115 which consist of purplish to gray and red bentonitic mudstone with sandstone beds that are
116 generally 0.5 m in thickness (Woody, 2006). Blue Mesa Member mudstones are pervasively
117 pedogenically modified in the core. These strata are interpreted to have accumulated primarily
118 as overbank deposits within a mixed-load meandering river system (Martz and Parker, 2010).
119 Previously reported U-Pb (ID-TIMS or CA-TIMS) ages from outcrop of the Blue Mesa Member
120 range from ~223 Ma to ~218 Ma (Heckert et al., 2009; Ramezani et al., 2011; Irmis et al., 2011;
121 Atchley et al., 2013; Rasmussen et al., 2020).

122 Strata of the Blue Mesa Member are overlain by sandstone-rich and conglomerate-bearing
123 strata of the Sonsela Member. Lucas (1993) and Heckert and Lucas (2002) refer to the base of
124 the Sonsela Member as a regionally significant unconformity, although this interpretation has
125 been questioned by Woody (2006) and Martz and Parker (2010) given that conglomeratic
126 sandstone of the Sonsela is interbedded with mudstone of the Blue Mesa Member. Martz and
127 Parker (2010) suggest that the transition from the Blue Mesa Member to the Sonsela Member
128 marks a change in depositional regime (from mainly overbank deposits to bedload-dominated
129 channel deposits) but does not mark a significant hiatus in deposition.

130 The Sonsela Member consists predominantly of sandstone with lesser mudstone and local
131 conglomerate. Sandstone beds are variable in thickness, have significant lateral extent, and
132 exhibit cut-and-fill structure (Woody, 2006). Conglomerate (with abundant volcanic clasts) is
133 common within the sandstone beds. Five units have been recognized, a lower sandstone
134 interval (Camp Butte beds), a lower-middle unit with abundant mudstone (Lot's Wife beds), a

135 middle sandstone and conglomerate unit (Jasper Forest/Rainbow Forest bed), a middle-upper
136 unit with pedogenic carbonate and abundant mudstone (Jim Camp Wash beds), and an upper
137 sandstone unit (Martha's Butte beds) (Martz and Parker, 2010). The five units are gradational,
138 with the main variation being the abundance of mudstone in two of the middle units. A reddish
139 siliceous horizon of uncertain regional extent has been recognized within the middle of the
140 upper mudstone-rich unit in the CPCP-PFNP13-1A core. Similar horizons within other exposures
141 of the Sonsela Member are marked by a significant die-off of the conifers that characterize
142 Petrified Forest National Park (Creber and Ash, 1990), a turn-over of the vertebrate fauna
143 (Parker and Martz, 2009, 2011), and perhaps a significant change in flora and paleoclimate
144 (Reichgelt et al., 2013; Nordt et al., 2015; Baranyi et al., 2017). U-Pb (CA-TIMS/zircon) ages from
145 the Sonsela Member range from ~220 to ~214 Ma (Ramezani et al., 2011; Marsh et al., 2019;
146 Rasmussen et al., 2020) from below the siliceous horizon and from ~214 to ~213 Ma (Ramezani
147 et al., 2011; Nordt et al., 2015; Kent et al., 2018; Rasmussen et al., 2020) from above.

148 Overlying the conglomeratic sandstones of the Sonsela Member is a purplish mudstone that
149 marks the base of the Petrified Forest Member (Fig. 2). This member consists of red and purple
150 mudstone with abundant paleosols and pedogenic carbonate nodules, with local conglomeratic
151 sandstone beds that formed in bedload-dominated streams. Near the top of the unit is the
152 Black Forest bed, which consists of limestone-pebble conglomerate and reworked andesitic tuff
153 (Ash, 1992). Zircon grains from the Black Forest bed have yielded U-Pb (ID-TIMS or CA-TIMS)
154 ages of ~213 Ma to ~210 Ma (Riggs et al., 2003; Heckert et al., 2009; Ramezani et al., 2011; Kent
155 et al., 2018; Rasmussen et al., 2020).

156 **3. SAMPLED HORIZONS**

157 We analyzed detrital zircon grains from twenty-nine samples collected from the Permian and
158 Triassic strata described above. Samples include one from the Coconino Sandstone, five from
159 the Moenkopi Formation (one that may be from the Wupatki Member and four from the
160 Holbrook Member), and twenty-three from the Chinle Formation (one from the Mesa Redondo
161 Member, three from the Blue Mesa Member, twelve from the Sonsela Member, and seven
162 from the Petrified Forest Member). Approximate stratigraphic positions of the samples are
163 shown on Figure 2, lithic characteristics are described in DR Table 1, and images of the sampled
164 material (both core and thin sections) are presented in Appendix 1. Each sample consisted of 20
165 cm (for sandstone) to 30 cm (for mudstone-siltstone) of $\frac{1}{4}$ sections of the core.

166 **4. ANALYTICAL AND INTERPRETIVE METHODS**

167 Zircon mineral separation was performed at the Arizona LaserChron Center
168 (www.laserchron.org) using methods modified from those outlined by Gehrels (2000), Gehrels
169 et al. (2008), and Gehrels and Pecha (2014) because of the small size of all samples and the
170 abundance of clay minerals in many samples. The process included using a hand-crusher to
171 break the samples apart, a gold pan for initial density separation, and an ultrasonic disruptor
172 (Hoke et al., 2014) to separate zircon crystals from clay mineral grains. Magnetic separation was

173 performed with a Frantz Isodynamic separator, followed by density separation using methylene
174 iodide.

175 Zircon grains greater than 60 μm in size were enclosed in 1-inch epoxy mounts along with
176 fragments of zircon standards SL (primary) and FC-1 and R33 (secondary). Mounts were
177 polished approximately 5-10 μm deep to expose the internal structure of the grains but retain
178 as much material as possible for subsequent CA-TIMS analysis. Imaging was performed with a
179 backscatter electron detector system (BSE) using a Hitachi S3400 scanning electron microscope
180 (SEM) to ensure analysis of zircon and to avoid inclusions and fractures. Mounts were cleaned
181 with 1% HCl and 1% HNO_3 prior to isotopic analysis.

182 U-Pb isotopic analyses were conducted by LA-ICPMS using a Teledyne/Photon Machines
183 Analyte G2 laser connected to a Thermo Element2 mass spectrometer. Analyses utilized a 20
184 μm diameter laser beam fired at 7 hz for 15 seconds, resulting in 10-12 μm deep pits. Details of
185 the analytical methods are reported in DR Table 2.

186 U-Pb ages are calculated with an in-house data-reduction routine (E2agecalc) following
187 methods of Pullen et al. (2018). Analyses of zircon grains from our samples are reported in DR
188 Table 3, with results filtered for discordance (using cutoffs of 80% and 105% concordance),
189 precision (10%), and common Pb (>600 cps counts of 204). Following the recommendations of
190 Horstwood et al. (2016), uncertainties for individual analyses include only internal (random or
191 measurement) uncertainty contributions, whereas uncertainties of pooled ages contain both
192 internal and external (systematic) contributions.

193 Detrital age distributions are displayed and analyzed with normalized probability density plots,
194 which are based on the individual ages and measured uncertainties from each sample.
195 Provenance interpretations are based on the main clusters of ages, with less emphasis on ages
196 that do not belong to clusters given the possibility that they are unreliable due to Pb loss,
197 inheritance, analysis of inclusions, high common Pb, or unusual Pb/U fractionation due to
198 ablation along fractures (Gehrels, 2014).

199 Analysis of provenance is conducted by comparison with age distributions from five likely
200 source regions for Permian-Triassic strata of the Colorado Plateau, which include the
201 Appalachian orogen, the Ouachita orogen, local basement rocks of southwestern Laurentia, the
202 East Mexico arc, and the Permian-Triassic magmatic arc developed along the Cordilleran margin
203 of southwestern North America (Fig. 1; Dickinson, 2018). The age distributions for these regions
204 include data from: (1) upper Paleozoic strata of the Appalachian foreland basin (Thomas et al.,
205 2017) and Illinois and Forest City basins (Kissock et al., 2018), (2) upper Paleozoic strata of the
206 Delaware (Xie et al., 2018), Fort Worth (Absalem et al., 2018), and Marathon (Thomas et al.,
207 2019) basins, (3) lower Paleozoic strata of the Grand Canyon (Gehrels et al., 2011) and
208 Cordilleran passive margin strata in southern California and northern Sonora (Gehrels and
209 Pecha, 2014), (4) Permian and Triassic strata of the Barranca and El Antimonio Formations of
210 Sonora (Gonzalez-Leon et al., 2009; Gehrels and Pecha, 2014), Jura-Cretaceous strata of the

211 Great Valley (DeGraaff-Surpless et al., 2002; Surpless et al., 2006; Wright and Wyld, 2007),
212 Permian-Triassic igneous rocks in California (Chen and Moore, 1982; Miller et al., 1995; Tobisch
213 et al., 2000; Barth and Wooden, 2006, 2011, 2013; Saleeby and Dunne, 2015), and (5) Mesozoic
214 strata that accumulated adjacent to the East Mexico arc (Ortega-Flores et al., 2014). Age
215 distributions for these five regions are presented in Figure 3.

216 Comparisons of age distributions are quantified using two different statistical measures that
217 examine the degree to which age distributions contain similar proportions of similar age
218 groups. Metrics used in this study include the Kolmogorov-Smirnov D (KS-D) values and Kuiper-
219 V values. The statistical basis as well as strengths and limitations of each of these metrics are
220 summarized by Saylor and Sundell (2016) and Vermeesch (2018a). Results from these
221 comparisons are presented in DR Table 4. The interpretations offered below are based on KS-D
222 values, although Kuiper-V values yield similar results. For both metrics, smaller values indicate a
223 higher degree of similarity of age distributions. Comparisons are also presented visually through
224 the use of multidimensional scaling (MDS) diagrams (Vermeesch, 2013; Saylor et al., 2017;
225 Wissink et al., 2018), which provide a 2-dimensional representation of the differences between
226 multiple age distributions. MDS analyses are based on KS-D values of the age distributions.

227 Maximum depositional ages (MDAs) are estimated from the youngest distinct cluster of ages in
228 each sample (e.g., Dickinson and Gehrels, 2009; Gehrels, 2014). The age of this cluster is
229 estimated using five different methods, each of which has strengths and limitations.

230 Complications with these methods arise from (1) the need to make unconstrained decisions
231 about which analyses to include or exclude from consideration, (2) the evidence that dates in
232 some clusters have been compromised by Pb loss, resulting in dates that post-date deposition,
233 (3) the evidence that some clusters also contain slightly older recycled grains that pre-date
234 deposition, and (4) issues of statistical robustness for some methods (Vermeesch, 2018b).

235 Following are short descriptions of the five methods:

- 236 • Age of the youngest peak on a probability density plot (PDP). This method is advantageous
237 because no decisions are made about which analyses are included/excluded, but it has the
238 disadvantage that no uncertainty is reported for the peak age.
- 239 • Weighted Mean age and uncertainty of the youngest cluster. This method calculates the
240 average age of a cluster by weighting each analysis according to the inverse-square of its
241 uncertainty. The reported uncertainty relates to the mean age (e.g., standard error of the
242 mean), not the age distribution of constituent analyses (e.g., standard deviation). An
243 advantage of this method is that it also yields a Mean Square of the Weighted Deviates
244 (MSWD), which is an indication of the degree to which the ages belong to a single
245 population (values of ~ 1 or less indicate a single population). A disadvantage of this method
246 is that the investigator must decide which ages are included in the calculation, which leads
247 to the possibility of subjective bias. In this study, clusters include the main set of continuous
248 ages, with boundaries selected at the youngest and oldest gap in ages. This calculation is
249 available from the Weighted Mean function in Isoplot (Ludwig, 2008).

- 250 • Tuffzirc age and uncertainty of the youngest cluster. This method uses the age extractor
251 function in Isoplot (Ludwig, 2008), which identifies the largest cluster of ages that overlap to
252 an acceptable degree (probability-of-fit > 0.05), reports the median value as the most likely
253 age, and uses the range of included ages to calculate an asymmetric uncertainty. The
254 reported uncertainty refers to the median value (not the range of constituent analyses).
255 Excluded ages are interpreted to pre-date the selected cluster (if older), or to be
256 compromised by Pb loss (if younger). This method is advantageous in that no subjective
257 decisions are made about including/excluding ages.
- 258 • Maximum Likelihood age and uncertainty. This method uses a maximum likelihood analysis
259 to determine the gaussian distribution that best fits the youngest cluster. The reported
260 uncertainty refers to the most likely value (not the range of constituent analyses). This
261 method is advantageous in that no subjective decisions are made about including/excluding
262 ages. It is available from the Unmix function of Isoplot (Ludwig, 2008).

263

264 Finally, we also use the minimum age model of Galbraith and Laslett (1993) and Vermeesch
265 (2020). This method assumes that a set of dates is a mixture of a discrete young component
266 and a continuous older component. It uses the method of maximum likelihood to determine
267 the age and uncertainty of the younger component. Calculations were conducted using IsoplotR
268 (Vermeesch, 2018b), which returns the minimum age and also a central age that is similar to
269 the weighted mean described above.

270

271 The results of these calculations are presented in DR Table 6. Shown separately are estimates
272 from the first four methods noted above, and the average of these four estimates, as well as
273 the minimum age (and uncertainty) which we interpret as the maximum depositional age.

274

275 DR Table 6 also reports the age and uncertainty of the youngest analysis from each sample. This
276 youngest age does not provide a reliable maximum depositional age given that the youngest
277 age of a distribution will always be younger than the true age due to analytical uncertainty
278 (Gehrels, 2014). For example, as described by Coutts et al. (2019), consider the analytical data
279 from a population of zircon grains that have exactly the same true age. Because of analytical
280 uncertainty, the measured ages of half of the analyses will be younger than the true age, and
281 half will be older, and the youngest age will be significantly younger than the mean (true) age.
282 Ironically, the more grains analyzed, the greater the inaccuracy of this youngest age
283 (Vermeesch, 2020)!

284 In addition to this statistical bias, the youngest single age will be even farther from the mean
285 (true) age if it has been compromised by Pb loss (e.g., Andersen et al., 2019). We report these
286 youngest ages because they provide important information about the possibility that analyses
287 included in the youngest cluster have also experienced Pb loss. DR Table 6 accordingly reports
288 this youngest age (and uncertainty), as well as information about its U concentration, the
289 average U concentration of the youngest cluster of ages, and whether the youngest age belongs
290 to the youngest cluster or is an outlier (based on Tuffzirc analysis). U concentration is important

291 because Pb loss is commonly correlated with the degree of radiation damage, which is a
292 function of U concentration (and age).

293 A second test of the likelihood that analyses belonging to the youngest cluster have
294 experienced Pb loss is provided by a plot of U concentration versus age for analyses belonging
295 to the youngest cluster. Such plots are shown for every sample in DR Table 3, and whether a
296 correlation exists is indicated in DR Table 6.

297 The average precision of individual analyses reported herein is 2.3% (2σ) for $^{206}\text{Pb}^*/^{238}\text{U}$ dates
298 and 2.6% for $^{206}\text{Pb}^*/^{207}\text{Pb}^*$ dates. For pooled ages, calculated as described above, the average
299 precision is 0.52% (2σ) including only internal uncertainties and 0.98% (2σ) including both
300 internal and external sources of uncertainty. The accuracy of our analyses can be estimated
301 from the age of the secondary standards that were analyzed with each set of unknowns. As
302 reported in DR Table 7 and shown on Figure 4, sets of $^{206}\text{Pb}^*/^{238}\text{U}$ dates for FC-1 are offset
303 between +0.25% and -0.45% from the reported $^{206}\text{Pb}^*/^{238}\text{U}$ date of 1099.9 Ma (Paces and
304 Miller, 1993), with an average offset for all 1,065 analyses of +0.03%. For R33, offsets range
305 from +0.85% to -0.95% from the assumed age of 419.3 Ma (Black et al., 2004), with an average
306 offset for all 291 ages of -0.23%. MSWD values for the sets of FC-1 and R33 ages are 0.95 and
307 0.92 (respectively) – this demonstrates that reported uncertainties for individual analyses are
308 accurate, and that MSWD values for sets of unknown ages are reliable indicators of the
309 existence of multiple age components.

310 Interpretation of our ages relative to the Geologic Time Scale is based on the August 2018
311 version of the International Chronostratigraphic Chart (Cohen et al., 2013).

312 U-Pb geochronology by LA-ICPMS also provides U concentrations and U/Th values for each
313 analysis, which can be used as a geochemical fingerprint of detrital zircon grains (e.g., Gehrels
314 et al., 2006, 2008; Riggs et al., 2012, 2016). This information is accordingly reported for each
315 analysis in DR Table 3, and for each set of analyses in DR Table 6.

316 **5. U-Pb GEOCHRONOLOGIC RESULTS**

317 Results of our U-Pb geochronologic analyses are described below, keyed to the age
318 distributions for individual samples that are shown on Figures 5, 6, and 7. Figure 8 presents age
319 distributions for combined sets of samples. Age distributions from all of the samples are
320 compared statistically in DR Table 4 using the five metrics described above, and MDS plots are
321 shown in Figure 9.

322 We note that Rasmussen et al. (2020) have reported a subset of the LA-ICPMS ages presented
323 herein. The ages reported in their study are for the grains selected for CA-TIMS analysis, which
324 in most cases are among the youngest grains in each of our samples (as documented in
325 Appendix 2). This strategy was followed assuming that these grains represent the youngest age
326 components in each sample, and accordingly provide the most useful maximum depositional
327 ages. The individual dates reported in the two studies are identical, but, given the selection

328 process noted above, the pooled ages reported by Rasmussen et al. (2020) are consistently
329 younger than the pooled ages reported herein. A comparison of the results of the two studies is
330 summarized in Appendix 2. The discussions below are based on the full set of ages from each
331 sample.

332 Sample numbers are registered to the CPCP core (CPCP-PFNP13-1A) by the number of the core
333 run and segment (e.g., our sample number 383-2 is from CPCP-PFNP13-1A-383Y-2, which
334 specifies that the material is from run 383, segment 2). The part of each segment that was
335 collected for geochronologic analysis is specified in DR Table 1.

336 **5.1 Coconino Sandstone**

337 Our sample from quartz arenite of the lower Permian (Leonardian) Coconino Sandstone
338 (sample 390-1) yielded 285 acceptable ages (DR Table 3; Figure 4). Most grains belong to two
339 broad age groups of ~2.0-1.0 Ga and ~640-295 Ma. Individual age peaks are at 2712, 1898,
340 1746, 1646, 1497, 1432, 1347, 1162, 1038, 667, 612, 590, 552, 476, 430, 419, 391, 374, 355,
341 341, and 300 Ma.

342 **5.2 Moenkopi Formation**

343 Five samples from the Lower-Middle Triassic Moenkopi Formation have been analyzed (Fig. 2).
344 The lowest sample (383-2) is assigned to the Wupatki Member based on the red-brown
345 laminated mudstone to fine-grained sandstone lithology (Fig. 2; Table DR 1). The age
346 distribution from this sample is very similar to that found in underlying upper Paleozoic strata,
347 with two dominant age groups from ~2.2 Ga to 1.0 Ga and from ~680 Ma to 250 Ma (Fig. 5).
348 Although the preferred interpretation for this sample is that it belongs to the lowest part of the
349 Moenkopi Formation, an alternative is that the sample is late Paleozoic in age, and perhaps
350 correlative with fine-grained clastic strata (e.g., the Toroweap Formation) that regionally overlie
351 the Coconino Sandstone. In an effort to provide a comparison with underlying and overlying
352 strata, the results from this sample are shown on Figures 5 and 6. Additional studies of the
353 sampled horizon are needed to resolve whether this sample belongs to the Moenkopi
354 Formation or underlying upper Paleozoic strata.

355 The upper four samples (349-3, 335-1, 327-2, and 319-2) are all from sandstone, siltstone, and
356 mudstone of the Holbrook Member. These samples yield generally similar age distributions
357 (average KS-D values of 0.19; DR Table 4), with significant proportions of ~1.42 Ga, 650-510
358 Ma, 290-270 Ma, and 250-235 Ma ages (Fig. 6). With ages from all four Moenkopi Formation
359 samples combined, PDP peak ages are 1420, 594, 543, 285, and 250 Ma (Fig. 8).

360 **5.3 Chinle Formation**

361 Twenty-three samples from the Mesa Redondo Member, Blue Mesa Member, Sonsela Member,
362 and Petrified Forest Member of the Chinle Formation have been analyzed (Fig. 2). Results from
363 each member are described separately below.

364 **5.4 Mesa Redondo Member**

365 One sample of sandstone from the Mesa Redondo Member (305-2) yields dominant age groups
366 of ~2.0-1.6 Ga, 1.44 Ga, 1.1-1.0 Ga, 750-500 Ma, and 450-300 Ma, and 290-220 Ma (Fig. 7), with
367 PDP peak ages of 1443, 1036, 618, 412, 323, 248, and 223 Ma. As reported in DR Table 4 and
368 shown on Figure 9B and 9C, the >240 Ma ages in this sample resemble ages in the underlying
369 Moenkopi Formation and Coconino Sandstone.

370 **5.5 Blue Mesa Member**

371 Three samples (297-2, 287-2, 261-1) of siltstone and mudstone from the Blue Mesa Member
372 yield similar results, with nearly identical <240 Ma ages and small but varying proportions of
373 ~1.64 Ga, 1.44 Ga, 1.1-1.0 Ga, 650-500 Ma, and 440-240 Ma ages (Figures 7 and 8). Both <240
374 Ma ages (Fig. 9A) and >240 Ma ages (Fig. 9C) differ from those in underlying strata of the Mesa
375 Redondo Member. Between 56% and 89% of the grains analyzed from these samples yield ages
376 between 232 and 210 Ma, with PDP peak ages of 221-220 Ma (Fig. 7; DR Table 6). With all three
377 samples combined, 62% of the ages are <240 Ma, and PDP peak ages are 1630, 1440, and 220
378 Ma (Fig. 8).

379 **5.6 Sonsela Member**

380 Twelve samples (243-3 to 158-2) from the Sonsela Member yield two different sets of age
381 distributions (Figures 7, 8, and 9; DR Table 3). The lower six samples (243-3 to 196-3), all
382 consisting of sandstone and subordinate siltstone (DR Table 1), yield small numbers of
383 Precambrian grains that are mostly ~1.65 and 1.44 Ga, with few ~1.1-1.0 Ga grains. These
384 samples yield between 53% and 79% ages <240 Ma, with most ages between 234 and 208 Ma,
385 and PDP peak ages of 221-218 Ma (Fig. 7). With ages from all six samples combined, 68% of the
386 grains are <240 Ma, and PDP peak ages are 1650, 1445, 1084, and 219 Ma (Fig. 8). Comparison
387 of age distributions (Figures 7 and 8), KS-D values (DR Table 4), and MDS patterns (Fig. 9)
388 suggests that the <240 Ma ages in lower Sonsela Member strata are similar to <240 Ma ages in
389 underlying Blue Mesa strata, whereas >240 Ma ages in the two sets of samples are less similar
390 due to the variability of ages from the three Blue Mesa Member samples. Ages that are >240
391 Ma in these strata have even less similarity to ages from the Mesa Redondo Member,
392 Moenkopi Formation, and Coconino Sandstone (Fig. 9; DR Table 4).

393 The upper six samples from the Sonsela Member (195-2 to 158-2) consist mainly of sandstone
394 and subordinate siltstone (DR Table 1). All six samples yield a subordinate but consistent
395 proportion of Precambrian ages that are mostly ~1.43 and 1.1-1.0 Ga, with few 1.65 Ga grains
396 (Fig. 7). Grains with ages of <240 Ma comprise between 39% and 77% of the grains analyzed.
397 These ages are somewhat younger than in lower Sonsela Member samples, with PDP peak ages
398 of 217-214 Ma. With all six samples combined, 50% of the grains are <240 Ma, and PDP peak
399 ages are 1643, 1434, 1082, 256, and 215 Ma (Fig. 8).

400 Statistical analysis (MDS patterns in Figure 9 and KS-D values in DR Table 4) shows that the <240
401 Ma ages in upper and lower Sonsela Member strata are significantly different, whereas >240
402 Ma ages are less distinct. Exceptions to this are >240 Ma ages in sample 243-3 (lower Sonsela
403 Member), which resemble equivalent ages in strata of the upper Sonsela Member (Fig. 9C), and
404 <240 Ma ages in sample 196-3, which share characteristics with strata of both the upper and
405 lower Sonsela Member (Fig. 9A). Ages from strata of the upper Sonsela Member show even less
406 overlap with ages from strata of the Blue Mesa Member and underlying units (Fig. 9 and DR
407 Table 4).

408 **5.7 Petrified Forest Member**

409 Seven samples (131-2 to 52-2) from the Petrified Forest Member were collected mainly from
410 claystone, mudstone, siltstone, and fine-grained sandstone, with only the lowest sample (131-
411 2) consisting of coarse-grained sandstone. The upper six fine-grained samples yield between
412 17% and 72% <240 Ma ages that are significantly younger than in underlying strata, with PDP
413 peak ages between 212 and 209 Ma. Ages that are >240 Ma in most of these samples differ
414 from equivalent ages in strata of the Blue Mesa Member and Sonsela Member, but overlap to
415 varying degrees with ages in strata of the Mesa Redondo Member, Moenkopi Formation, and
416 Coconino Sandstone (Fig. 9C; DR Table 4). With the six samples combined, 35% of the grains are
417 <240 Ma, and PDP peak ages are 1636, 1430, 1032, 629, 379, 287, and 209 Ma (Fig. 8). The
418 lowest sample (131-2), consisting of coarse-grained sandstone, differs from the other Petrified
419 Forest Member samples, with an age peak of 221 Ma, and a greater proportion (68%) of >240
420 Ma ages (Fig. 7). The <240 Ma ages are very similar to equivalent ages in strata of the lower
421 Sonsela Member (Fig. 9A; KS-D=0.12), whereas >240 Ma ages are slightly more similar to ages in
422 the upper Sonsela Member (KS-D=0.17) than in the lower Sonsela Member (KS-D=0.22) (Fig.
423 9C).

424 **5.8 Summary of Chinle results**

425 The patterns of LA-ICPMS ages described above suggest that the studied part of the Chinle
426 Formation comprises four different units, each of which has a distinct chronologic signature for
427 both <240 Ma and >240 Ma ages (Fig. 8). These chronostratigraphic units correspond to the
428 Mesa Redondo Member, Blue Mesa Member and lower part of the Sonsela Member, upper
429 part of the Sonsela Member, and Petrified Forest Member.

430 **6. U AND Th GEOCHEMISTRY OF CHINLE ZIRCON GRAINS**

431 In an effort to evaluate whether the Triassic zircon grains from the four chronostratigraphic
432 units also have distinct chemical signatures [following Riggs et al. (2012, 2016)], Figure 10
433 summarizes the U concentrations and U/Th values for Triassic zircon grains analyzed from each
434 unit. The patterns exhibited in these plots suggest that (1) zircon grains from the Mesa
435 Redondo Member are significantly different from zircon grains in overlying strata, (2) grains in
436 strata of the Blue Mesa Member and lower Sonsela Member differ from grains in overlying
437 strata of the upper Sonsela Member and Petrified Forest Member, and (3) grains in strata of the

438 upper Sonsela Member and Petrified Forest Member have distinctive and slightly different
439 bimodal patterns. Plots showing U concentrations and U/Th values for individual samples are
440 included in DR Table 3.

441 **7. PROVENANCE INTERPRETATIONS**

442 Detrital zircon geochronology has previously been used to reconstruct the provenance of
443 Permian and Triassic strata of the Colorado Plateau by Riggs et al. (1996, 2003, 2012, 2013,
444 2016), Dickinson and Gehrels (2003, 2008), Gehrels et al. (2011), Lawton et al. (2015), and
445 Marsh et al. (2019). The results of most of these chronological studies, and a large number of
446 stratigraphically based analyses, have recently been summarized by Dickinson (2018). The
447 following sections compare our new results with this existing information.

448 The following comparisons are based in part on qualitative comparison of age distributions of
449 the strata that we have analyzed and of age distributions from five potential source areas
450 (summarized on Figure 3). As described by Gehrels (2000), such comparisons focus on the
451 degree to which two age distributions contain similar proportions of similar ages. Comparisons
452 are also based on the results of statistical analyses (DR Table 4) that compare our results with
453 the age distributions of possible source areas, and on graphic displays of these comparisons
454 using MDS plots (Fig. 9).

455 **7.1 Coconino Sandstone**

456 Lawton et al. (2015) and Dickinson (2018) suggest that lower Permian strata of the Colorado
457 Plateau comprise a regional blanket of eolian strata that was shed predominantly from the
458 Appalachian and/or Ouachita orogens, with increasing input in northern regions from local
459 basement rocks exposed in the Uncompahgre or Ute Uplift (Fig. 1). These interpretations are
460 supported by the age distributions shown on Figures 5 and 11, with southern strata (Coconino,
461 Cedar Mesa, and White Rim sandstones) forming a distinct group dominated by
462 Appalachian/Ouachita detritus, and northern strata (Castle Valley and Cutler strata) forming a
463 separate group with increasing proportions of ca 1.44 Ga grains. The age distribution from our
464 Coconino Sandstone sample (390-1) fits well with other strata from the southern Colorado
465 Plateau in having abundant 1.2-1.0 and 670-300 Ma (Appalachian-Ouachita) grains and a low
466 proportion of ~1.44 Ga grains (Figures 5 and 11).

467 **7.2 Moenkopi Formation**

468 As summarized on Figure 6, the detrital zircon ages from our four Holbrook Member samples
469 are generally similar to ages from a Holbrook Member sandstone reported by Dickinson and
470 Gehrels (2008). Dominant >300 Ma age groups and interpreted source terranes include ~1.44
471 Ga and subordinate ~2.0-1.6 Ga grains derived from Laurentian Precambrian basement and
472 ~670-300 Ma grains derived from Ouachita/Gondwana sources. Based on comparison with
473 detrital zircon ages from strata that accumulated in proximity to the East Mexico and southern
474 Cordilleran arcs (Fig. 3), 300-260 Ma grains (PDP peak ages of 285, 284, 265, 260, and 279) are

475 interpreted to have been shed from the East Mexico arc (peak age of 284 Ma), whereas 260-
476 230 Ma grains (peak ages of 250, 248, 228, 245, and 239 Ma) were likely shed from Early-
477 Middle Triassic parts of the Cordilleran magmatic arc in California and northwestern Mexico
478 (peak ages of 243, 236, and 226 Ma) (Fig. 3). Statistical analyses (DR Table 4) suggest nearly
479 equal contributions from the Ouachita orogen, local basement rocks, and the East Mexico arc.

480 More detailed analysis of the age distributions (Fig. 6) and MDS patterns (Fig. 9) suggest that
481 the lower two samples (349-3 and 335-1) [plus sample CP8 of Dickinson and Gehrels (2008)] are
482 dominated by ~1.44 Ga and ~285 Ma grains, whereas the upper two samples (327-2 and 319-2)
483 are dominated by ~620-590 Ma and ~250-230 Ma grains. The age distributions (Fig. 6) and
484 comparison metrics (Fig. 9C; DR Table 4) suggest that the lower samples were shed mainly from
485 local basement rocks (KS-D=0.35), whereas the upper samples were shed largely from the
486 Ouachita orogen (KS-D=0.23).

487 **7.3 Chinle Formation**

488 Our results from detrital zircon grains recovered from strata of the Chinle Formation are
489 consistent with the provenance and paleogeographic reconstructions offered by Riggs et al.
490 (1996, 2003, 2012, 2013, 2016), Dickinson (2018), and Marsh et al. (2019). Given the observed
491 age distributions (Fig. 7) and the location of our study site relative to Late Triassic
492 paleogeographic and paleotectonic features of southwestern North America (Fig. 12), likely
493 sources for pre-Triassic grains include rocks exposed in the Ouachita orogen to the southeast
494 and the Ancestral Mogollon highlands to the south and southwest. Given the abundance of ash
495 layers, bentonitic mudstone, and near-depositional-age zircon grains in strata of the Chinle
496 Formation, and the existence of arc-related plutons and volcanic rocks of Triassic age in Sonora
497 and southern California (Barth and Wooden, 2006, 2011, 2013; Saleeby and Dunne, 2015; Riggs
498 et al., 2016), Stewart et al. (1986), Riggs et al. (2012, 2016), Dickinson (2018), Marsh et al.
499 (2016), and many other researchers conclude that Triassic grains in Chinle strata were derived
500 from the active arc built along the southern Cordilleran margin. The occurrence in fore-arc and
501 back-arc strata of very similar distributions of ages (Fig. 3) is inconsistent with interpretations
502 (e.g., Hildebrand, 2009, 2013) that the early Mesozoic arc was located far from southwestern
503 North America.

504 Although our data are entirely consistent with the provenance interpretations outlined above,
505 the density of our sampling and the large number of analyses from most samples provide
506 opportunities to reconstruct temporal changes in Triassic provenance in greater detail, and with
507 the benefit of statistical analyses to quantify conclusions. Following are interpretations based
508 on strata belonging to each of the different members of the Chinle Formation.

509 **7.4 Mesa Redondo Member**

510 The provenance of strata belonging to the Mesa Redondo Member is similar to that of the
511 underlying Moenkopi Formation, with our sample (305-2) containing abundant ~640-300 Ma
512 grains derived from Ouachita/Gondwana sources as well as ~290-260 Ma grains derived from

513 the East Mexico arc (Fig. 8). Statistical analysis confirms higher similarity of >240 Ma grains with
514 Ouachita sources (0.58) than with Appalachian (0.35) or local basement (0.15) sources (DR
515 Table 4). This sample also yields a significant proportion of Triassic ages that approximate the
516 depositional age for these strata (Fig. 7). These young grains, with a PDP age peak of 223 Ma,
517 are interpreted to have been transported primarily by aeolian processes from the active
518 magmatic arc to the west (Fig. 12). Statistical analysis demonstrates that the Triassic ages in
519 these samples are significantly different from ages in overlying strata (Fig. 9A) and that the
520 >240 Ma ages are similar to those in some strata of the Petrified Forest Member (Fig. 9C).

521 **7.5 Blue Mesa Member**

522 Our three samples from strata of the Blue Mesa Member yield a large proportion of Triassic
523 zircon grains (Figures 7 and 8) that were derived from the active Cordilleran magmatic arc to
524 the west (Fig. 12), and a small proportion of pre-240 Ma grains that were shed from local
525 basement rocks and the Ouachita and/or Appalachian orogens (Fig. 8). Statistical analysis
526 confirms that the Triassic ages in all these samples are quite similar (Fig. 9A), whereas the age
527 distributions of >240 Ma grains in the three samples are more variable (Fig. 9C; DR Table 4).

528 **7.6 Lower Sonsela member**

529 The lower six samples from the Sonsela Member yield a large proportion of Triassic grains
530 derived from the Cordilleran magmatic arc, and fewer ages derived from local basement rocks
531 and Ouachita/Gondwana sources (Figures 7 and 8). Distinctive among the older grains is a
532 significant proportion of ~1.44 Ga grains that most likely signal increased input from the
533 Ancestral Mogollon highlands to the southwest (Marsh et al., 2019) (Fig. 12). MDS analysis
534 demonstrates that the <240 Ma and >240 Ma ages in these samples are quite similar, with the
535 main difference being the larger number of ~1.1 Ga grains in sample 243-3 (Figures 7 and 9C).

536 **7.7 Upper Sonsela Member**

537 The upper six samples from the Sonsela Member reveal a continued low contribution from the
538 Ouachita orogen, and a significant increase in the proportion of ~1.08 Ga and 260-240 Ma
539 grains (Figures 7 and 8). The ~260-240 Ma grains were likely derived from Permian-Early Triassic
540 igneous rocks along the southern Cordilleran margin (Saleeby and Dunne, 2015; Riggs et al.,
541 2016), exposed in the Ancestral Mogollon Highlands (Fig. 12). The prominent ~1.44 and 1.08 Ga
542 grains in these samples may also have been shed from highland sources to the south and
543 southwest. Triassic grains in these samples record a slightly younger (230 to 204 Ma, peak age
544 of 215 Ma) phase of magmatism along the Cordilleran margin. Significant changes in both <240
545 Ma and >240 Ma ages occur between samples 196-3 and 195-2 (Figure 7). MDS analysis
546 demonstrates that patterns of both <240 Ma and >240 Ma ages are consistent among the six
547 upper Sonsela Member samples, but are distinct from ages in all other parts of the Chinle
548 Formation (Figures 7 and 9).

549 **7.8 Petrified Forest Member**

550 Strata of the Petrified Forest Member record an important shift in provenance, with
551 significantly greater detrital input from the East Mexico arc (~287 Ma) and the Ouachita orogen
552 (~640-300 Ma), and a broader range of >1.0 Ga basement sources (Figures 7 and 8). Triassic
553 grains in these strata are also significantly younger, with ages of 228 to 200 Ma (peak age of
554 209 Ma).

555 An exception to these patterns is recorded by ages from the coarse-grained sandstone of
556 sample 131-2, which has Precambrian grains that are mainly ~1.1-1.0 and 1.44 Ga (like upper or
557 lower Sonsela Member; Fig. 9C), and Triassic grains that are ~221 Ma (like strata of the lower
558 Sonsela Member and Blue Mesa Member; Fig. 9A). This lower Petrified Forest Member sample
559 is interpreted to have been reworked mainly from lateral equivalents of underlying strata of the
560 Sonsela Member and Blue Mesa Member, with little or no input from the active arc to the west.

561 **8. MAXIMUM DEPOSITIONAL AGES**

562 The depositional age of Triassic strata on the Colorado Plateau is of considerable interest
563 because of the rich faunal and paleoclimatic records preserved within the Moenkopi Formation
564 and Chinle Formation, and as the zircon-based geochronological framework for the early
565 Mesozoic when coupled with paleomagnetic polarity stratigraphy and astrochronology (Olsen
566 et al., 2018, 2019; Kent et al., 2018, 2019; Rasmussen et al., 2020). There accordingly have been
567 many prior attempts to determine the depositional age of these strata by dating igneous zircon
568 grains in ash beds or volcanic cobbles and detrital zircon grains in clastic strata (e.g., Riggs et al.,
569 1996, 2003, 2012, 2013, 2016; Heckert et al., 2009; Dickinson and Gehrels, 2009; Irmis et al.,
570 2011; Ramezani et al., 2011, 2014; Atchley et al., 2013; Nordt et al., 2015). As part the Colorado
571 Plateau Coring Project, Kent et al. (2018) and Rasmussen et al. (2020) report the results of CA-
572 TIMS analyses on many of the same samples reported herein. All of the available CA-TIMS ages,
573 and the preferred age models of Kent et al. (2019) and Rasmussen et al. (2020), are shown on
574 Figure 13.

575 Maximum depositional ages (MDA's) have been determined using the minimum age model of
576 Vermeesch (2020). The possibility that this maximum depositional age has been compromised
577 by Pb loss is evaluated mainly by determining whether there is a correlation between U
578 concentration and age. One criterion is whether the youngest single age has higher U
579 concentration than the average of the youngest cluster – if yes than the youngest analysis (and
580 perhaps other analyses within the youngest cluster) may have experienced Pb loss. A second
581 criterion is whether analyses within the youngest cluster display an inverse correlation between
582 U concentration and age – if yes, then the higher U and younger analyses within the cluster may
583 have experienced Pb loss. An additional criterion is whether the youngest date is excluded from
584 the cluster determined by Tuffzirc analysis. Samples in which all three methods suggest the
585 presence of Pb loss are shown with red arrows on Figure 13. Rasmussen et al. (2020) document
586 Pb loss in zircon grains from several of our samples by showing that CA-TIMS ages are
587 commonly older than LA-ICPMS ages from the same crystals.

588 **8.1 Coconino Sandstone**

589 Our analyses do not provide a useful MDA for strata of the Coconino Sandstone (sample 390-1)
590 because few late Paleozoic ages were recovered from this sample.

591 **8.2 Holbrook Formation of the Moenkopi Formation**

592 Of our four samples from the Holbrook Member of the Moenkopi Formation, three yield MDA's
593 that young upward from 248.05 (± 1.82) Ma to 246.63 (± 1.92) Ma to 236.78 (± 9.92) Ma (DR
594 Table 6). These MDA's are consistent with the inferred Early-Middle Triassic age of the strata
595 and the corresponding ~ 251 -237 Ma range for Early and Middle Triassic time on the Geologic
596 Time Scale (Cohen et al., 2013). All three samples show patterns of U concentration that
597 suggest the possibility of Pb loss (DR Table 6).

598 **8.3 Mesa Redondo Member of the Chinle Formation**

599 Our one sample (305-2) from strata of the Mesa Redondo Member yields an MDA of $223.24 \pm$
600 1.50 Ma (DR Table 6). Patterns of U concentration do not indicate the presence of Pb loss (DR
601 Table 6). This MDA overlaps with CA-TIMS ages of ~ 224.7 -221.7 Ma from the same sample but
602 is slightly older than the preferred single-grain age of ~ 221.7 Ma (Rasmussen et al., 2020).
603 However, the LA-ICPMS MDA of 223.24 ± 1.50 is younger than CA-TIMS ages of ~ 225.2 Ma
604 (Ramezani et al., 2011) and ~ 227.6 (Atchley et al., 2013) from outcrop samples of the Mesa
605 Redondo Member.

606 **8.4 Blue Mesa Member of the Chinle Formation**

607 Our three samples (297-2, 287-2, 261-1) from strata of the Blue Mesa Member yield MDA's of
608 219.68 ± 0.46 , 218.62 ± 0.98 , and 221.23 ± 1.02 Ma (DR Table 6). All samples yield MSWD
609 values >1.0 (average of 2.4), which indicates the presence of multiple age populations and/or
610 Pb loss (DR Table 6). Patterns of U concentration suggest the possible presence of Pb loss in all
611 three samples, and likely Pb loss in sample 287-2. As shown on Figure 13, these MDA's are
612 slightly younger than CA-TIMS ages of ~ 221.8 Ma [from sample 297-2; Rasmussen et al. (2020)],
613 and ~ 220.5 Ma [from sample 287-2; Rasmussen et al. (2020)]. From upper strata, our age is
614 similar to a CA-TIMS age from outcrop of ~ 220.1 Ma (Atchley et al., 2013) but significantly
615 younger than a CA-TIMS age of ~ 223.0 Ma (Ramezani et al., 2011), also from outcrop.

616 **8.5 Lower part of the Sonsela Member**

617 Our six samples from the lower part of the Sonsela Member (243-3 to 196-3) yield MDA's of
618 219.27 ± 0.44 Ma (sample 243-3), 220.81 ± 0.44 Ma (sample 227-3), 221.30 ± 0.48 Ma (sample 215-
619 2), 219.21 ± 0.66 Ma (sample 210-1), and 221.06 ± 0.50 Ma (sample 201-1). The sixth, uppermost
620 sample (196-3) yields younger ages with an MDA of 217.93 ± 0.56 Ma. MSWD values for these
621 samples are all high (average of 2.6), which demonstrates the presence of multiple age
622 components. There is evidence for Pb loss in analyses from samples 243-3 and 210-1.

623 As shown on Figure 13, these MDA's are 1-3 m.y. older than most CA-TIMS ages from
624 equivalent strata. From oldest to youngest, the CA-TIMS ages include ~220.1 Ma [from outcrop;
625 Atchley et al. (2013)] from near the base, through ~218.8 Ma [sample 243-3; Rasmussen et al.
626 (2020)], ~217.7 Ma [sample 227-3; Rasmussen et al. (2020)], ~219.3 Ma [from outcrop;
627 Ramezani et al. (2011)], ~217.8 Ma [sample 215-2; Rasmussen et al. (2020)], ~218.0 Ma [from
628 outcrop; Ramezani et al. (2011)], and ~215.7 Ma and 214.4 Ma [samples 201-1 and 196-3;
629 Rasmussen et al. (2020)] at the top. The LA-ICPMS-based MDA's ages are also older than a
630 ~216.6 Ma MDA determined on LA-ICPMS ages from an outcrop sample of sandstone in the
631 middle part of the lower Sonsela Member, exposed ~132 km north of the CPCP core site (Marsh
632 et al., 2019).

633 **8.6 Upper part of the Sonsela Member**

634 The lower five samples from the upper Sonsela Member yield similar preferred MDA's of
635 214.36 ± 0.68 Ma (sample 195-2), 216.32 ± 0.72 Ma (sample 188-2), 216.19 ± 0.62 Ma (sample 182-
636 1), 214.81 ± 0.70 Ma (sample 177-1), and 217.07 ± 0.86 Ma (sample 169-1). An upper sample
637 yields a younger MDA of 214.18 ± 0.54 Ma (sample 158-2). All samples yield MSWD values
638 greater than 1.0 (average of 2.6) (DR Table 6), demonstrating the presence of multiple age
639 components. Most samples have patterns of U concentration that suggest the possibility of Pb
640 loss. The lower five MDA's are 2-3 m.y. older than CA-TIMS ages from equivalent strata, which
641 include outcrop ages of ~213.9 (Ramezani et al., 2011), ~213.6 Ma (Nordt et al., 2015), and
642 ~213.1 Ma (Ramezani et al., 2011), and CPCP core ages of ~214.0 Ma [samples 182-1 and 177-1;
643 Rasmussen et al. (2020)]. A CA-TIMS age of ~213.5 Ma for the upper sample [158-2; Rasmussen
644 et al. (2020)] is nearly identical to our age determination.

645 **8.7 Petrified Forest Member**

646 Our seven samples from the Petrified Forest Member yield three sets of MDA's. The lowest unit
647 (sample 131-2) yields an MDA of 221.54 ± 0.44 Ma, which is significantly older than MDA's in
648 adjacent strata. Four samples near the middle of the unit yield similar MDA's of 211.53 ± 3.26
649 Ma (sample 116-1), 209.90 ± 1.56 Ma (sample 104-3), 210.42 ± 1.08 Ma (sample 92-2), and
650 211.86 ± 0.94 Ma (sample 84-2). The MDA's for two of these samples overlap with an ID-TIMS
651 age of ~211.9 Ma (Irmis et al., 2011) from equivalent strata in outcrop, the other two younger
652 MDA's may be compromised by Pb loss (Fig. 13).

653 Two upper samples, from the Black Forest bed, yield preferred MDA's of 208.26 ± 3.38 Ma
654 (sample 66-1) and 209.75 ± 0.42 Ma (sample 52-2). These MDA's are similar to CA-TIMS ages of
655 ~210.2 Ma from core [sample 52-2; Rasmussen et al. (2020)] and ~209.9 Ma from outcrop
656 (Ramezani et al., 2011), but are significantly younger than outcrop-based ID-TIMS ages of
657 ~211.0 Ma (Heckert et al., 2009) and ~213.0 Ma (Riggs et al., 2003). Most of our samples yield
658 MSWD values greater than 1.0 (average of 1.5), suggesting the presence of multiple age
659 components, and have patterns of U concentration that suggest the presence of Pb loss.

660 **9. COMPARISON OF LA-ICPMS, CA-TIMS, AND MAGNETOSTRATIGRAPHIC CONSTRAINTS ON**
661 **DEPOSITIONAL AGE OF CHINLE FORMATION STRATA**

662 Our maximum depositional ages for strata of the Chinle Formation range from ~223.2 to ~208.3
663 Ma, which is similar to the ~227.6 to ~209.9 Ma range of CA-TIMS ages (Fig. 13). All available U-
664 Pb data therefore suggest that the analyzed Chinle Formation strata are Late Triassic, and
665 probably Norian in age (Dickinson, 2018), given the assigned ages of ~237 to ~201.3 for Late
666 Triassic time (Cohen et al., 2013) and ~227 to ~208.5 Ma (Cohen et al., 2013) or ~205.7 Ma
667 (Kent et al., 2017) for Norian time.

668 Figure 13 presents a comparison of our LA-ICPMS-based average ages and maximum
669 depositional ages, all available ID- and CA-TIMS ages [from Riggs et al. (2003), Heckert et al.
670 (2009), Ramezani et al. (2011), Irmis et al. (2011), Atchley et al. (2013), Nordt et al., (2015), Kent
671 et al. (2018), and Rasmussen et al. (2020)], and two age models that are based on
672 magnetostratigraphic and CA-TIMS geochronologic information (Kent et al., 2019; Rasmussen et
673 al., 2020). As shown on this figure, our LA-ICPMS ages reveal two first-order patterns. The first
674 pattern is that the LA-ICPMS-based ages overlap with most CA-TIMS ages and both age models
675 for most strata belonging to the Blue Mesa Member and Petrified Forest Member, but are
676 significantly older for strata of the Sonsela Member. The second pattern is that most LA-ICPMS
677 based ages belong to five main clusters (~223 Ma, ~222-220 Ma, ~217-215 Ma, ~212-211, and
678 ~210 Ma), whereas the other chronologic records show a relatively simple pattern of upward
679 younging (Fig. 13). The following discussion explores these two patterns – details of the
680 magnetostratigraphic information, CA-TIMS data, and age models are discussed by Kent et al.
681 (2018, 2019) and Rasmussen et al. (2020).

682 As shown on Figure 13, the LA-ICPMS-based average ages and MDA's presented herein overlap
683 with the other chronometers for sequences which are dominated by fine-grained strata (e.g.,
684 Blue Mesa Member and Petrified Forest Member), but are several million years too old for
685 sequences which are dominated by coarse-grained strata (Sonsela Member) (Fig. 13). This
686 pattern appears to hold for member-scale stratigraphic units (e.g., strata from the Petrified
687 Forest Member), although some individual samples clearly do not follow this pattern. For
688 example, of the six samples from the Petrified Forest Member that yield LA-ICPMS ages which
689 overlap with the other chronometers, four are mudstone-siltstone and two are sandstone. In
690 the lower Sonsela Member, of the six samples with LA-ICPMS ages that predate the other
691 chronometers, five are sandstone and one is siltstone. These exceptions suggest that the
692 dominant lithic characteristics and depositional environment of a member (e.g., dominantly
693 fine-grained floodplain deposits for the Petrified Forest Member versus dominantly coarse-
694 grained channel deposits of the Sonsela Member [Woody, 2006]), are more important than the
695 grain size of an individual horizon in controlling the recognition of near-depositional-age zircon
696 grains.

697 The observed pattern that predominantly fine-grained strata of the Mesa Redondo, Blue Mesa,
698 and Petrified Forest members yield reliable LA-ICPMS ages, whereas predominantly coarse-

699 grained sandstones of the Sonsela Member do not, is surprising for two reasons. First, in terms
700 of provenance (as described above), strata of the Mesa Redondo, Blue Mesa, and Petrified
701 Forest members are interpreted to have been shed mainly from the Ouachita orogen, which
702 lacks Triassic igneous rocks, whereas strata of the Sonsela Member were shed from the
703 Cordilleran magmatic arc to the southwest, which contains abundant Permian and Triassic
704 igneous rocks (Fig. 3). Second, as shown in the margins of Figures 7 and 8, Triassic zircon grains
705 are significantly (~2x) more abundant in strata of the Sonsela Member than in underlying and
706 overlying strata. Based on these two observations, one might expect that strata of the Sonsela
707 Member would yield reliable MDA's, whereas strata from the Mesa Redondo Member, Blue
708 Mesa Member, and Petrified Forest Member would not.

709 We suggest that these counter-intuitive relations result in large part from our analytical method
710 of only analyzing zircon grains that are >60 um, combined with the maximum size of zircons
711 that can be transported in fine-grained versus coarse-grained sediments. For coarse-grained
712 sediment, >60 um zircon grains could include both transported (detrital) components that
713 predate deposition, as well as zircons that are air-fall in origin and approximately of
714 depositional age. A MDA calculated from a mix of these grains would accordingly pre-date
715 deposition. In contrast, Triassic zircon grains from fine-grained strata would tend to be mostly
716 air-fall in origin given that the older, transported grains are too small to analyze. An MDA
717 calculated from zircons that are primarily of air-fall origin would accordingly approach the true
718 depositional age.

719 The relations described above suggest that convergence versus divergence of the chronologic
720 records results from connections between depositional setting, grain size, provenance, and
721 analytical methods, which together conspire to control the proportions of air-fall (near-
722 depositional age) versus slightly older detrital zircon grains recognized in our samples. We
723 suggest that the three chronometric records agree (to within ~2 m.y.) for strata of the lower
724 Blue Mesa Member and middle-upper Petrified Forest Member because of the availability of
725 zircon grains of air-fall origin, which are near depositional age and both <60 um and >60 um in
726 size, versus the scarcity of pre-depositional-age Triassic grains of sufficient size for analysis due
727 to the lack of Triassic rocks in the source region (mainly the Ouachita orogen) and the small
728 (<60 um) grain size of most sediment. In contrast, for the Sonsela Member, the LA-ICPMS
729 average ages and MDA's are interpreted to pre-date the other chronologic records because the
730 sediment was derived from the south, where abundant igneous rocks of Permian-Triassic age
731 were exposed, and the grain size of the detrital (pre-depositional-age) zircons was sufficiently
732 large that many would have been analyzed.

733 A test of this hypothesis is provided by MSWD values of the weighted means calculated for ages
734 from samples belonging to the various stratigraphic units. As shown in DR Table 6, average
735 MSWD values for samples from dominantly fine-grained strata of the Mesa Redondo-Blue Mesa
736 and Petrified Forest units are 1.7 and 1.3 (respectively), whereas coarser grained strata of the
737 lower and upper Sonsela units yield higher MSWD values of 2.6 and 2.1 (respectively). These

738 values are consistent with the interpretation that Triassic zircon grains in coarser-grained units
739 have a greater range of ages than Triassic zircon grains in finer-grained units.

740 These interpreted connections may also provide an explanation for the patterns of offset of the
741 CA-TIMS ages of Rasmussen et al. (2020) relative to the LA-ICPMS ages and
742 magnetostratigraphic age models in the Sonsela Member (Fig. 13). For strata of the upper
743 Sonsela Member, the CA-TIMS and magnetostratigraphic records converge because the
744 methods of grain selection were apparently successful in identifying populations of syn-
745 depositional age zircon grains. For strata of the lower Sonsela Member, however, these
746 methods were unsuccessful in identifying a sufficient number of depositional-age zircon grains
747 to determine a reliable MDA, presumably because of their low abundance relative to older
748 transported grains.

749 The second main pattern exhibited by the three chronometers is that most of the LA-ICPMS-
750 based average ages and MDA's belong to five main clusters (~223 Ma, ~222-220 Ma, ~217-215
751 Ma, ~212-211, and ~210 Ma), whereas the other chronologic records show a relatively simple
752 pattern of upward younging (Fig. 13). For the ~222-220 Ma cluster, a plausible interpretation,
753 following from the connections described above, is that ~222-220 Ma zircon grains of air-fall
754 origin accumulated in fine-grained strata of the lower Blue Mesa Member, and were then
755 recycled from age-equivalent strata into predominantly coarser grained channel sands of the
756 upper Blue Mesa Member and lower Sonsela Member. Grains from these same sources appear
757 to have also been recycled into sandstone sample 131-2 of the lower Petrified Forest Member
758 (Fig. 13). The ~212-211 Ma cluster may have formed in a similar fashion, with initial
759 accumulation of near-depositional-age air-fall zircons in mudstones of sample 116-1, followed
760 by recycling of these grains from age-equivalent strata into coarser-grained strata of samples
761 104-3, 92-2, and 84-2 (Fig. 13).

762 The source of zircon grains that belong to the ~217-215 Ma cluster is less obvious given the lack
763 of recognized fine-grained strata dominated by zircons of this age (Fig. 13). One possibility is
764 that ~217-215 Ma grains were eroded from fine-grained strata exposed elsewhere [perhaps
765 near Sonsela Buttes (Marsh et al., 2019) or near the Cordilleran magmatic arc] that are
766 dominated by grains of this age. A second possibility is that fine-grained strata dominated by
767 ~217-215 Ma ages were originally present in the lower Sonsela Member, but were removed by
768 erosion and recycled into strata of the upper Sonsela Member. Previous workers have
769 suggested the existence of a hiatus or hiatuses (Ramezani et al., 2011) or an erosional event
770 (Rasmussen et al., 2020) at approximately this stratigraphic level, as shown by the preferred
771 age model of Rasmussen et al. (2020) on Figure 13. The occurrence of very different <240 Ma
772 ages, >240 Ma ages, and U/Th values in samples 196-3 and 195-2 suggests that this shift in
773 provenance, accumulation of a condensed section, or formation of an unconformity likely
774 coincides with the proposed boundary between strata of the lower Sonsela Member and upper
775 Sonsela Member. As discussed by Ramezani et al. (2011) and Rasmussen et al. (2020), the
776 possibility of an unconformity or condensed section near this stratigraphic position has

777 important implications for Chinle stratigraphy and fundamental Late Triassic biotic and climatic
778 changes. It should be noted, however, that no stratigraphic evidence for such an unconformity
779 was recognized in the CPCP core.

780 **10. IMPLICATIONS FOR THE STRATIGRAPHY OF THE CHINLE FORMATION**

781 The interpreted connections between the three geochronologic records and Chinle stratigraphy
782 provide an opportunity to reconstruct the depositional history of the Chinle Formation.

783 Fundamental assumptions in reconstructing this history are that:

784 (1) Chinle Formation strata encountered in the CPCP core record nearly continuous deposition
785 as described in the age model of Kent et al. (2019), perhaps with a period of erosion or very
786 slow deposition in the middle part of the Sonsela Member (Rasmussen et al., 2020).

787 (2) LA-ICPMS ages recovered from strata of the Chinle Formation belong to five separate groups
788 (red vertical bars of Figure 13) due to the hypothesized connections between stratigraphy, grain
789 size, and proportions of near-depositional-age (air-fall) versus older (recycled) zircon ages.

790 (3) Late Triassic igneous activity in the Cordilleran magmatic arc provided a nearly continuous
791 supply of zircon grains of air-fall origin to the Chinle deposystem. This assumption is supported
792 by the relatively continuous distribution of U-Pb ages within the Cordilleran magmatic arc and
793 back-arc (upper curves of Figure 13).

794 The interpreted stratigraphic evolution is summarized below and shown schematically on
795 Figure 14. Important phases in this evolution are as follows:

796 A: An LA-ICPMS MDA of ~223.3 Ma from our one sample from the Mesa Redondo Member
797 (305-2) agrees with the magnetostratigraphic information, the two age models, and the set of
798 CA-TIMS ages from this sample, presumably because these fine-grained strata are dominated
799 by zircon grains of air-fall origin. Older CA-TIMS ages of ~225.2 Ma (Ramezani et al., 2011) and
800 ~227.6 (Atchley et al., 2013) from outcrops of the Mesa Redondo Member may be
801 compromised by an abundance of recycled zircon grains.

802 B: LA-ICPMS average ages of ~221-220 Ma for most grains from fine-grained strata in the lower
803 part of the Blue Mesa Member are also near depositional age, presumably because the >60 um
804 zircon grains in these fine-grained strata are dominated by air-fall (or slightly reworked)
805 components. Minimum ages for these samples are somewhat younger, presumably due to Pb
806 loss.

807 C: LA-ICPMS ages from strata of the upper Blue Mesa Member significantly pre-date deposition,
808 presumably because these strata are dominated by recycled zircons. The predominance of 221-
809 220 Ma LA-ICPMS ages suggests that most zircon grains were recycled from lateral equivalents
810 of underlying strata in the lower part of the Blue Mesa Member. CA-TIMS ages also pre-date
811 deposition, presumably because of the difficulty of isolating near-depositional-age grains of air-
812 fall origin.

813 D: This pattern continues up through most of the lower Sonsela Member, with LA-ICPMS ages
814 remaining at 221-220 Ma (except where compromised by Pb loss) due to recycling of strata
815 from lateral equivalents of the lower Blue Mesa Member. Most CA-TIMS ages predate the age
816 of deposition because depositional-age (air fall) grains were diluted by recycled components.

817 E: The age patterns from sandstones of the upper Sonsela Member are somewhat puzzling
818 given that the dominant ~217-215 Ma LA-ICPMS ages pre-date deposition, but fine-grained
819 strata that could have sourced grains of these ages are not present in the lower Sonsela
820 Member (Fig. 13). One possibility, as described above, is that the ~217-215 Ma grains were
821 eroded from fine-grained strata exposed elsewhere [perhaps near Sonsela Buttes (Marsh et al.,
822 2019) or from the Cordilleran magmatic arc] that are dominated by grains of this age. A second
823 possibility is that fine-grained strata dominated by ~217-215 Ma ages were originally present in
824 the underlying lower Sonsela Member, but were removed by erosion and recycled into strata of
825 the upper Sonsela Member. An erosional event of the appropriate age and stratigraphic
826 position has been described by Ramezani et al. (2011) and by Rasmussen et al. (2020), as shown
827 by their age model on Figure 13. The occurrence of very different <240 Ma ages, >240 Ma ages,
828 and U/Th values in samples 196-3 and 195-2 suggests that this change in provenance,
829 condensed section, or unconformity most likely coincides with the boundary between lower
830 and upper Sonsela Member strata. As discussed by Rasmussen et al. (2020), the possibility of an
831 unconformity or condensed section near this stratigraphic position has important implications
832 for Chinle stratigraphy and fundamental Late Triassic biotic and climatic changes.

833 F: The dominance of pre-depositional-age grains in sample 131-2 provides strong evidence for
834 recycling of detrital zircons from lateral equivalents of underlying strata of the Blue Mesa
835 Member or lower Sonsela Member.

836 G: All chronometers agree for strata of sample 116-1, presumably because these fine-grained
837 strata are dominated by air-fall (or slightly reworked) detrital zircons.

838 H: LA-ICPMS ages from sandstones of the middle Petrified Forest Member (samples 104-3, 92-
839 2, and 84-2) slightly predate deposition (except where compromised by Pb loss) because they
840 were recycled from lateral equivalents of immediately underlying fine-grained strata (e.g.,
841 sample 116-1).

842 I: Most LA-ICPMS ages agree with the other chronometers for strata of the Black Forest bed
843 because this unit is dominated by air-fall (or slightly reworked) detrital zircon grains. The
844 minimum age for sample 66-1 is somewhat younger, presumably due to Pb loss.

845 **11. CONCLUSIONS**

846 First-order conclusions that result from our U-Pb geochronologic analyses of detrital zircon
847 grains from the Coconino Sandstone, Moenkopi Formation, and Chinle Formation are as
848 follows:

849 1. The provenance of strata belonging to the Coconino Sandstone and Moenkopi Formation can
850 be reconstructed by comparison of our LA-ICPMS ages (Figures 5 and 6) with age distributions
851 that characterize potential source regions (Figure 3). As shown on Figures 5 and 11, data from
852 our sample of the Coconino Sandstone and equivalent sandstones of the southern Colorado
853 Plateau suggest that these strata belong to an eolian blanket that was derived largely from the
854 Ouachita and/or Appalachian orogens, whereas strata from the northern Colorado Plateau
855 consist mainly of sediment derived from local basement uplifts (Fig. 1; Dickinson and Gehrels,
856 2003; Gehrels et al., 2011; Lawton et al., 2015). Lower-Middle Triassic strata of the Moenkopi
857 Formation record a very different dispersal system, with most detritus derived from the
858 Ouachita orogen, the East Mexico arc, and early phases of the Cordilleran magmatic arc (Figures
859 6 and 9).

860 2. LA-ICPMS ages from strata of the Chinle Formation belong to five groups that generally
861 correspond to the main stratigraphic units (Figures 7, 8, and 13). Maximum depositional ages
862 calculated from <240 Ma ages and provenance interpretations derived from >240 Ma ages are
863 as follows:

864 -- Strata of the Mesa Redondo Member yield a preferred MDA of ~223.3 Ma, and were derived
865 mainly from the Ouachita orogen.

866 -- Strata of the Blue Mesa Member yield MDA's of ~221.2 to ~218.6 Ma, and were derived from
867 local basement and Ouachita sources.

868 -- Strata in the lower part of the Sonsela Member yield similar MDA's of ~221.3 to ~219.2 Ma
869 (plus an uppermost sample with an MDA of ~217.9 Ma). Detritus was derived mainly from local
870 basement (especially ~1.44 Ga) sources, perhaps located in the ancestral Mogollon highlands to
871 the south.

872 -- Strata in the upper part of the Sonsela Member yield younger MDA's of ~217.1 to ~214.4 Ma,
873 plus an uppermost sample with an MDA of ~214.2 Ma. Grains with >240 Ma ages were derived
874 mainly from Precambrian basement (mainly ~1.44 Ga) and Grenville-age rocks to south, as well
875 as the East Mexico arc.

876 -- Strata of the Petrified Forest Member yield LA-ICPMS ages that belong to three separate
877 groups. The lowest sample yields an MDA of ~221.5, which is significantly older than ages from
878 adjacent strata. The middle four samples yield MDA's of ~211.9 to ~209.9 Ma, whereas the
879 upper two samples yield MDA's of ~209.8 and ~208.3 Ma. All six upper samples contain
880 abundant >240 Ma grains that were shed from a broad range of Ouachita, local basement, and
881 East Mexico arc sources.

882 3. Patterns of U and Th concentration in Triassic zircon grains from the Chinle Formation belong
883 to four distinct groups that generally coincide with the chronostratigraphic units described
884 above. Changes in U and Th concentrations are interpreted to record variations in the chemistry

885 of arc magmatism through time, as has been documented previously by Barth and Wooden
886 (2006, 2011, 2013) and Riggs et al. (2010, 2012, 2016).

887 4. Comparison of the Chinle Formation MDA's with magnetostratigraphic information (Kent et
888 al., 2018, 2019) and CA-TIMS geochronologic information (Rasmussen et al., 2020) from the
889 CPCP core, plus CA-TIMS ages reported from outcrop samples, indicates that LA-ICPMS MDA's
890 approximate depositional ages for most strata of the Mesa Redondo Member, Blue Mesa
891 Member, and Petrified Forest Member (except where compromised by Pb loss), but
892 significantly pre-date deposition for strata of the Sonsela Member (Fig. 13). The correlation of
893 age patterns with stratigraphy is interpreted to reflect the proportions of air-fall (or slightly
894 reworked) versus recycled (older) zircon grains: fine-grained strata are dominated by near-
895 depositional ages because most zircon grains are air-fall (or slightly reworked) in origin,
896 whereas coarse-grained strata are dominated by pre-depositional ages because recycled zircon
897 grains dilute the abundance of air-fall crystals.

898 5. This hypothesized connection between stratigraphy and the three geochronologic records
899 supports the following depositional history for Chinle Formation strata encountered in the CPCP
900 core (Figures 13 and 14):

901 -- LA-ICPMS ages and magnetostratigraphic information (Kent et al., 2019) indicate that the
902 sampled part of the Mesa Redondo Formation was deposited at ~223.3 Ma. CA-TIMS ages of
903 ~225.2 Ma (Ramezani et al., 2011) and ~227.6 (Atchley et al., 2013) from outcrop samples
904 suggest that strata of the Mesa Redondo Member in other areas are dominated by older
905 recycled components.

906 -- Magnetostratigraphic information (Kent et al., 2019) suggests that strata of the Blue Mesa
907 Member and lower Sonsela Member accumulated between ~222 Ma and ~214 Ma, whereas
908 LA-ICPMS MDA's are consistently 222-220 Ma for the same strata (except for the uppermost
909 sample of ~218 Ma). This suggests that most zircons in strata of the upper Blue Mesa Member
910 and lower Sonsela Member were recycled from lateral equivalents of strata of the lower Blue
911 Mesa Member. The observation that most CA-TIMS ages from these strata also pre-date
912 deposition is interpreted to result from the dilution of air-fall zircon crystals by older recycled
913 zircon grains.

914 -- Strata of the upper Sonsela Member accumulated between ~215 and ~213 Ma, as
915 constrained by magnetostratigraphic information and CA-TIMS ages. LA-ICPMS MDAs from
916 these strata are ~217-215 Ma, which indicates that they are dominated by zircons recycled
917 from older units. The lack of samples in the lower Sonsela Member that are dominated by
918 ~217-215 Ma grains suggests that zircon grains of this age in upper Sonsela Member strata may
919 have been transported from sections of the Chinle Formation exposed outside of the PEFO
920 area. It is also possible that such strata were exposed in the PEFO area, but were removed
921 during an erosional event inferred by Rasmussen et al. (2020) from the pattern of CA-TIMS ages
922 in the upper Sonsela Member (Fig. 3). Significant changes in <240 Ma ages, >240 Ma ages, and

923 U-Th values suggest that this unconformity, if present, occurs between samples 196-3 and 195-
924 2.

925 -- All available evidence suggests that mudstone and subordinate sandstone of the middle
926 Petrified Forest Member accumulated at ~212-211 Ma, and the Black Forest bed in the upper
927 part of the unit accumulated at ~210 Ma. In contrast, LA-ICPMS ages recovered from sample
928 131-2, from the lower part of the Petrified Forest Member, are dominantly ~221 Ma, suggestive
929 of recycling from lateral equivalents of strata of the Blue Mesa Member and lower Sonsela
930 Member.

931 6. Comparisons of our LA-ICPMS ages, the available CA-TIMS data, and magnetostratigraphic
932 information provide insights into methods for determining the depositional age of fluvial strata.
933 Our results show that the most reliable information comes from sequences dominated by fine-
934 grained clastic strata (mudstone and siltstone) given that these strata have a low abundance of
935 pre-depositional-age zircon grains of the appropriate size (>60 μm diameter) for routine
936 analysis by LA-ICPMS. Mudstone-siltstone samples may accordingly yield a high proportion of
937 >60 μm zircon grains that are air-fall in origin (or only slightly reworked) and thereby record the
938 age of deposition. In contrast, sedimentary sequences dominated by sandstone could well yield
939 abundant >60 μm zircon grains that predate deposition, thereby diluting syn-depositional-age
940 zircon grains. Future attempts to determine depositional ages from fluvial strata should
941 accordingly focus on sequences dominated by fine-grained strata, rather than sandstones, in
942 spite of the challenges of extracting and analyzing the smaller zircon crystals.

943 **12. AUTHOR CONTRIBUTION**

944 NG and GG generated the LA-ICPMS data reported in this paper. All coauthors were involved in
945 acquiring the samples that were analyzed and/or interpreting the data. GG prepared this
946 manuscript with input from all co-authors.

947 **13. COMPETING INTERESTS**

948 The authors declare that they have no conflict of interest.

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1222 **FIGURE CAPTIONS**

1223 **Figure 1.** Map showing the main basement provinces of southern North America and Mexico.
1224 Also shown are locations of the study area within the Colorado Plateau, outlines of Ancestral
1225 Rocky Mountains uplifts, and the Permian-Triassic magmatic arc along the continental margin
1226 of southwestern North America. Modified from Gehrels et al. (2011).

1227 **Figure 2.** Strata encountered in the Colorado Plateau Coring Project (adapted from Olsen et al.,
1228 2018). Sampled horizons are shown relative to core depth, stratigraphic depth, and
1229 stratigraphic nomenclature relevant for the Petrified Forest region. Detailed descriptions of
1230 samples are provided in DR Table 1; images of the sampled material are presented in Appendix
1231 1.

1232 **Figure 3.** Normalized probability density plots of U-Pb (zircon) ages from source terranes.
1233 Distinctive age groups include 1750-1620 Ma and 1520-1360 Ma ages from southwest Laurentia
1234 basement provinces, 1240-960 Ma ages from Grenville-age provinces exposed in the
1235 Appalachian and Ouachita orogens, 640-570 Ma and 480-370 Ma ages characteristic of the
1236 Appalachian orogen, 670-300 Ma ages from the Ouachita orogen, 300-260 Ma ages from the
1237 East Mexico arc, and 260-200 Ma ages belonging to the Cordilleran magmatic arc of
1238 southwestern North America. See text for sources of information.

1239 **Figure 4.** Plot showing the accuracy of $^{206}\text{Pb}^*/^{238}\text{U}$ dates of secondary standards analyzed
1240 during the current study. Each pair of symbols represents the weighted mean age and 2σ
1241 uncertainty of R33 and FC-1 analyses conducted with each sample, expressed as % offset from
1242 reported ID-TIMS dates of 1099.9 Ma for FC-1 (Paces and Miller, 1993) and 419.26 Ma for R33
1243 (Black et al., 2004). For FC-1, 1065 analyses are reported, with MSWD = 0.95 for all analyses. For
1244 R33, 295 analyses are reported, with MSWD = 0.92 for all analyses. Data are reported in DR
1245 Table 7.

1246 **Figure 5.** Normalized probability density plots of detrital zircon ages from our sample of the
1247 Coconino Sandstone and from other lower Permian sandstones of the Colorado Plateau.
1248 Numbers of constituent analyses are shown for each sample. Data are from ¹Dickinson and
1249 Gehrels (2003), ²Gehrels et al. (2011), ³Lawton et al. (2015), and ⁴this study. Shown for
1250 reference are age ranges from the Appalachian orogen (purple bands) and from local basement
1251 rocks (blue bands) (from Figure 3), which are interpreted by previous researchers to have
1252 sourced most of the detritus in these units. Also shown is our sample 383-2, which is
1253 interpreted to belong to the Wupatki Member of the Moenkopi Formation, but has an age
1254 signature characteristic of lower Permian strata of the Colorado Plateau.

1255 **Figure 6.** Probability density plots of detrital zircon ages from four samples from the Moenkopi
1256 Formation (lower four curves) as well as a Moenkopi sample from Dickinson and Gehrels
1257 (2008). Numbers of constituent analyses are shown for each sample. Samples 349-3, 335-1,
1258 327-2, and 319-2, plus the sample from Dickinson and Gehrels (2008), are all from the Holbrook

1259 Member. Sample 383-2 is interpreted to belong to the Wupatki Member, but has an age
1260 distribution that resembles lower Permian strata. Source regions are interpreted to include
1261 local basement rocks (blue bands), the Ouachita orogen (green bands), the East Mexico arc (red
1262 band), and the Late Permian-Triassic arc built along the Cordilleran margin (orange band).

1263 **Figure 7.** Normalized probability density plots of detrital zircon ages from twenty-four samples
1264 from the Mesa Redondo, Blue Mesa, Sonsela, and Petrified Forest Members of the Chinle
1265 Formation. Numbers of constituent analyses are shown for each sample. Age distributions older
1266 than 240 Ma are exaggerated by 10x. Black tick marks indicate the interpreted maximum
1267 depositional ages for each sample (from DR Table 6). Source regions are interpreted to include
1268 local basement rocks (blue bands), the Ouachita orogen (green bands), the East Mexico arc (red
1269 band), and the Late Permian-Triassic arc built along the Cordilleran margin (orange band).
1270 Percent of all grains that are <240 Ma in age are shown for each sample on the left.

1271 **Figure 8.** Normalized probability density plots of detrital zircon ages from each set of samples
1272 analyzed in this study. Numbers of constituent analyses are shown for each sample. Age
1273 distributions older than 240 Ma for Chinle strata are exaggerated by 10x relative to <240 Ma
1274 ages. Age distributions for Moenkopi and Coconino Sandstones are exaggerated by 5x relative
1275 to Chinle ages. Source regions are interpreted to include local basement rocks (blue bands), the
1276 Ouachita orogen (green bands), the East Mexico arc (red band), and the Late Permian-Triassic
1277 arc built along the Cordilleran margin (orange band). Results from sample 383-2 are not
1278 included in this plot because of its uncertain stratigraphic position. Data from sample 131-2 are
1279 omitted because they differ from ages present in other samples from the Petrified Forest
1280 Member. Percent of all grains that are <240 Ma in age are shown for each sample on the left.

1281 **Figure 9.** MDS plot (Vermeesch, 2013) comparing age distributions of samples analyzed herein
1282 with each other and with possible source areas. MDS (metric) analyses were conducted using
1283 the software of Saylor et al. (2018). Data from samples analyzed herein are in DR Table 3. Ages
1284 for source regions are from the sources cited in the text. Stars represent MDS values for sets of
1285 examples, with the exception that sample 131-2 is not included with other Petrified Forest
1286 samples.

1287 **Figure 10.** Density distributions of U concentration versus U/Th for Triassic grains in the four
1288 chronostratigraphic units recognized in this study. Plots made with Hf density plotter software
1289 of Sundell et al. (2019).

1290 **Figure 11.** MDS plot comparing age distributions of Permian strata of the Colorado Plateau with
1291 each other and with potential source regions including the Appalachian orogen, Ouachita
1292 orogen, and basement rocks of southwestern North America. Data sources are described in
1293 Figures 3 and 4. The data support the interpretation of Lawton et al. (2015) that the Coconino,
1294 Cedar Mesa, and White Rim sandstones (cool shades) belong to a regional blanket of eolian
1295 strata that was derived largely from the Appalachian and/or Ouachita orogen, where strata of

1296 the Castle Valley and Cutler formations (warm shades) include greater proportions of detritus
1297 derived from local basement sources.

1298 **Figure 12.** Sketch map of relevant tectonic features in southwestern Laurentia during Late
1299 Triassic time [adapted from Figure 42 of Dickinson (2018)].

1300 **Figure 13.** Plot showing the available chronologic information for strata of the Chinle Formation
1301 from the study area. LA-ICPMS results are shown using red crosses for interpreted maximum
1302 depositional ages [using the minimum age approach of Vermeesch (2020)], and various symbols
1303 for the four age estimates (and the average) of the youngest cluster. Red arrows indicate that
1304 LA-ICPMS ages may be compromised by Pb loss (DR Table 6). CA-TIMS and ID-TIMS ages are
1305 shown in approximate stratigraphic position (as shown by Kent et al., 2019), with outcrop
1306 samples in gray symbols and core samples using black symbols. Smaller symbols represent ID-
1307 TIMS ages or CA-TIMS ages based on a single age or of uncertain reliability. Stratigraphic units
1308 are keyed to dominant rock type, with brown = mudstone and siltstone, yellow = sandstone,
1309 pink = bentonite. Average grain size of each sample is shown with bars on left (from Appendix 1
1310 and DR Table 1). PDP curves to right show 2.0 Ga to 240 Ma ages, as plotted on Figure 7. Also
1311 shown are age models of Kent et al. (2019) and Rasmussen et al. (2020). Vertical red bands
1312 show interpreted ages of main clusters of LA-ICPMS ages.

1313 Curves across top of diagram show the distribution of ages from (1) fore-arc strata of the
1314 Barranca and El Antimonio Groups in Sonora (Gonzalez-Leon et al., 2009; Gehrels and Pecha,
1315 2014) and the Great Valley Group in California (DeGraaff-Surpless et al., 2002; Surpless et al.,
1316 2006; Wright and Wyld, 2007), (2) Permian-Triassic igneous rocks in California (Chen and
1317 Moore, 2982; Miller et al., 1995; Tobisch et al., 2000; Barth and Wooden, 2006, 2011, 2013;
1318 Saleeby and Dunne, 2015), and (3) strata of the Chinle Formation in other parts of the Colorado
1319 Plateau (Dickinson and Gehrels, 2008; Riggs et al., 2012; Marsh et al., 2019). Diamond-shaped
1320 symbols beneath curves represent individual ages.

1321 **Figure 14.** Depositional model of strata of the Chinle Formation encountered in the CPCP core.
1322 Each time slice contains information about the dominant grain size of the host sedimentary
1323 rock, the abundance of syn-depositional-age zircon grains that are interpreted to be air-fall in
1324 origin, and the abundance of recycled zircon grains that pre-date deposition.

Figure 1 (NAmap)

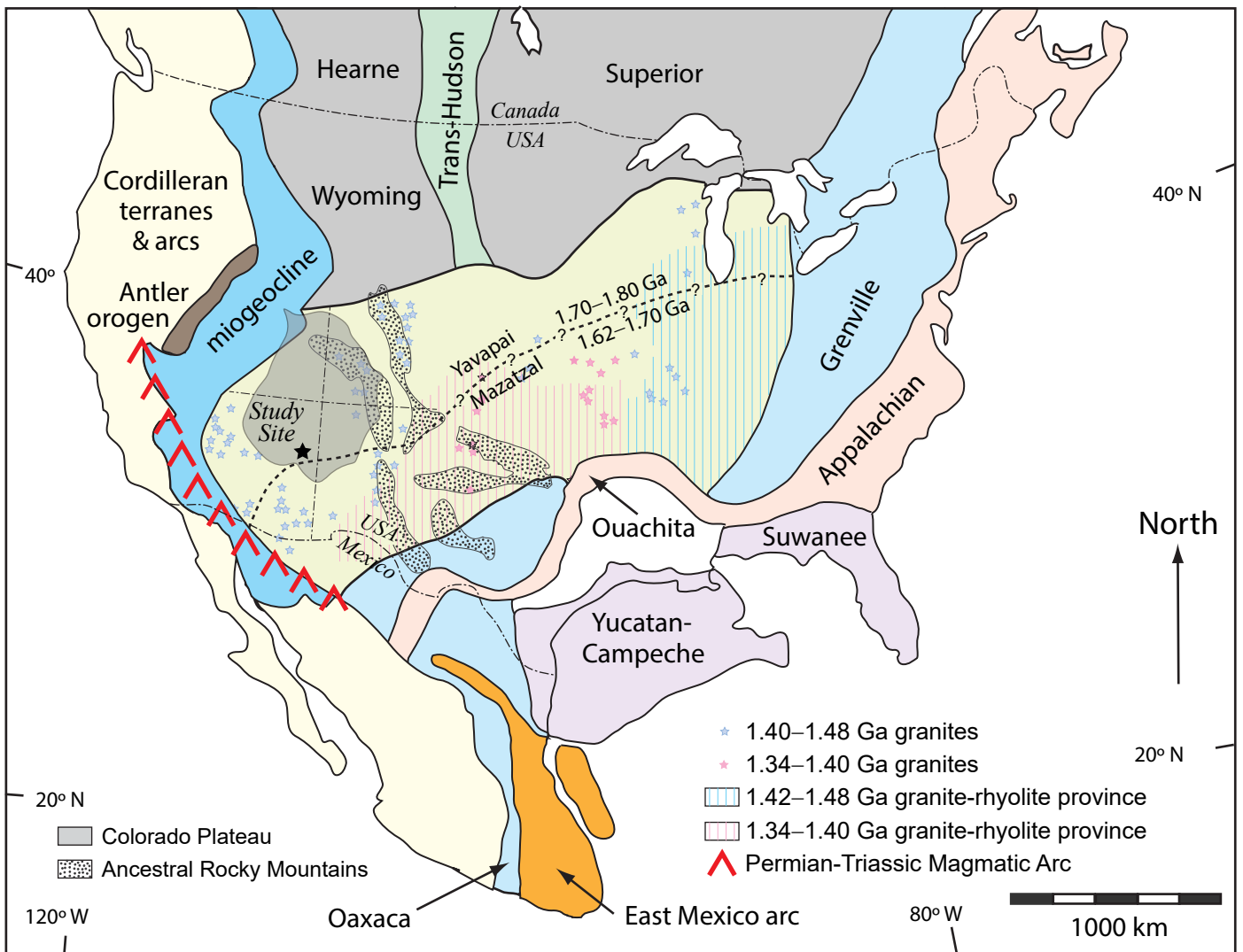


Figure 2 (Strat Column)

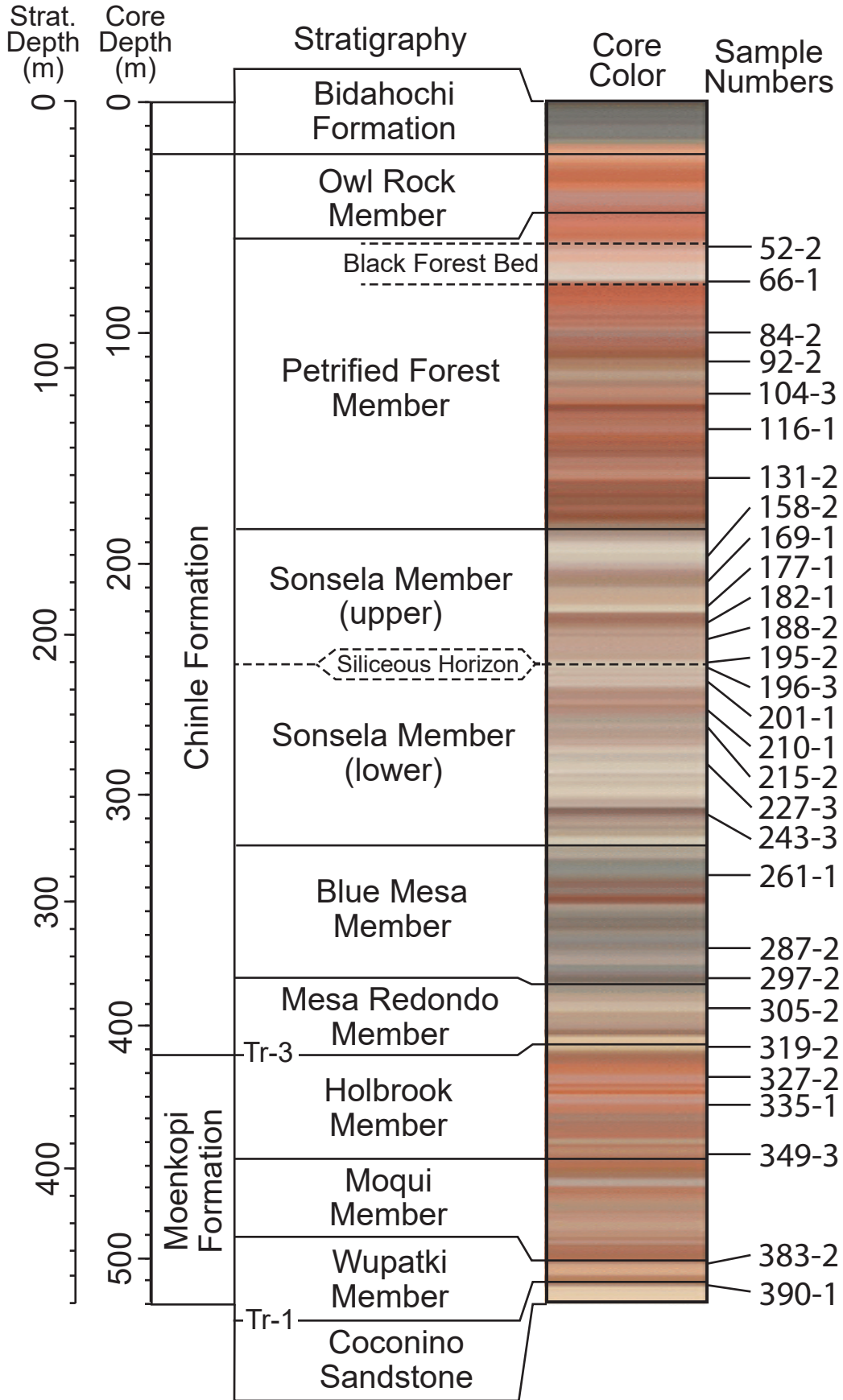


Figure 3 (App-Ouach-Bsmt-EMArc-CordArc PDP)

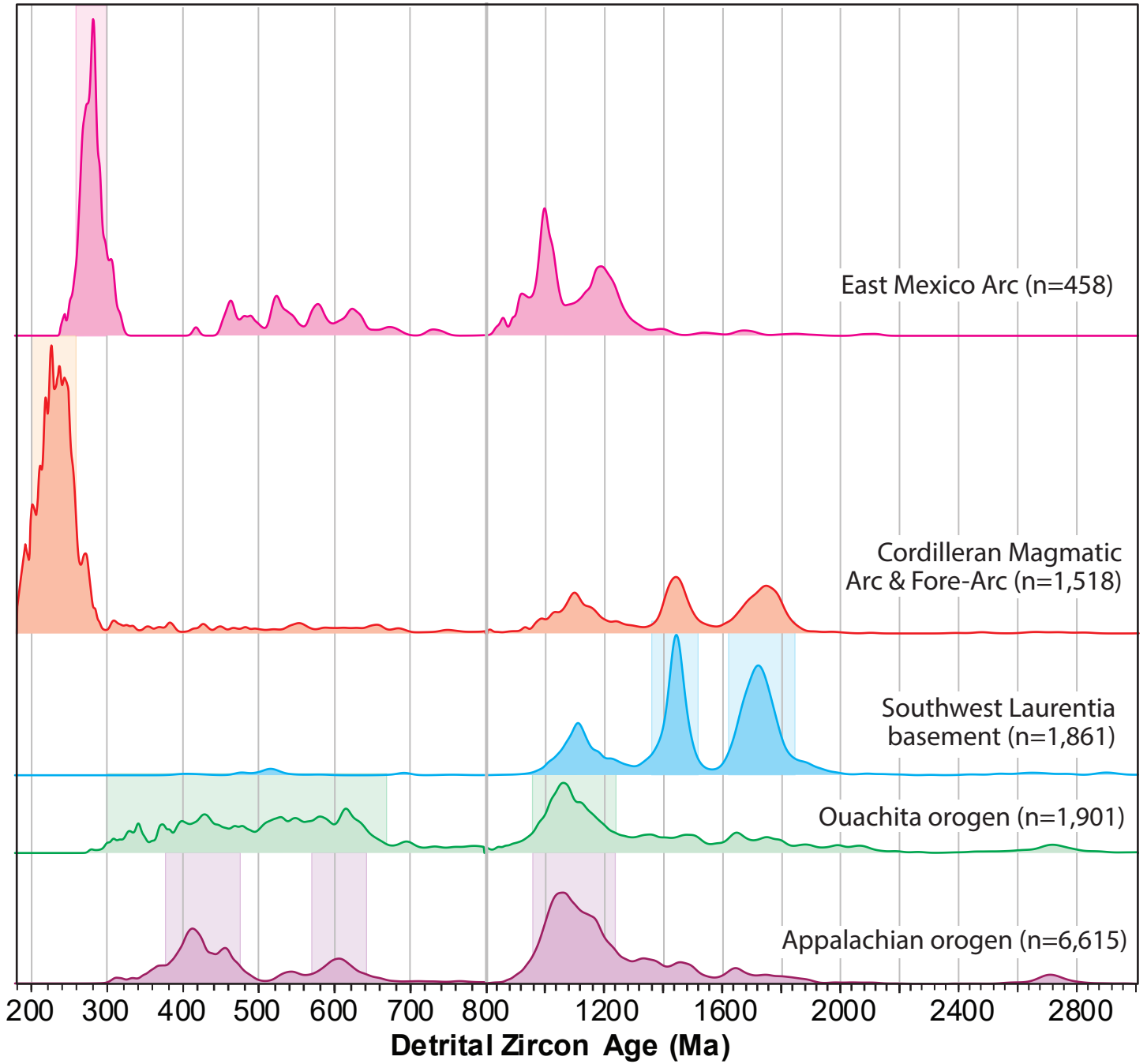


Figure 4 (Standard Ages)

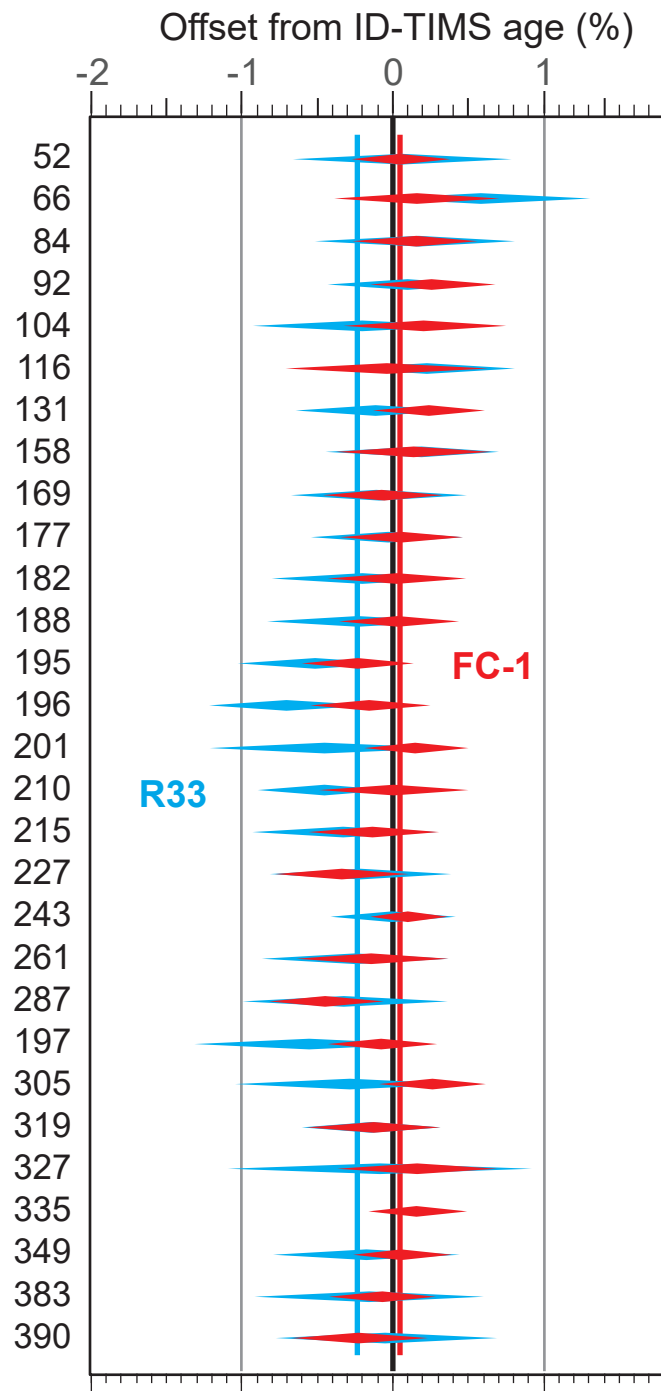


Figure 5 (Coconino PDP)

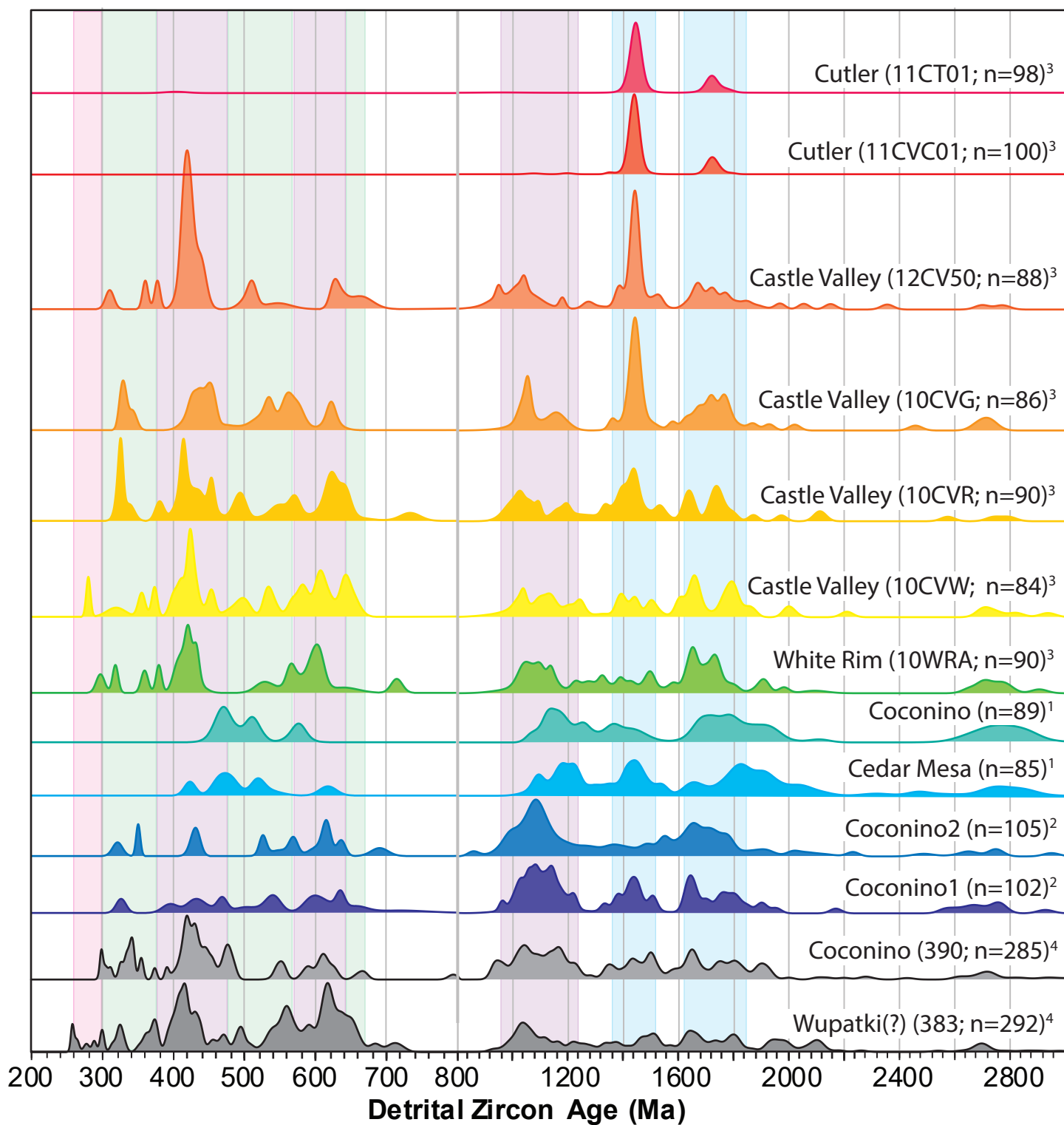


Figure 6 (Moenkopi PDP)

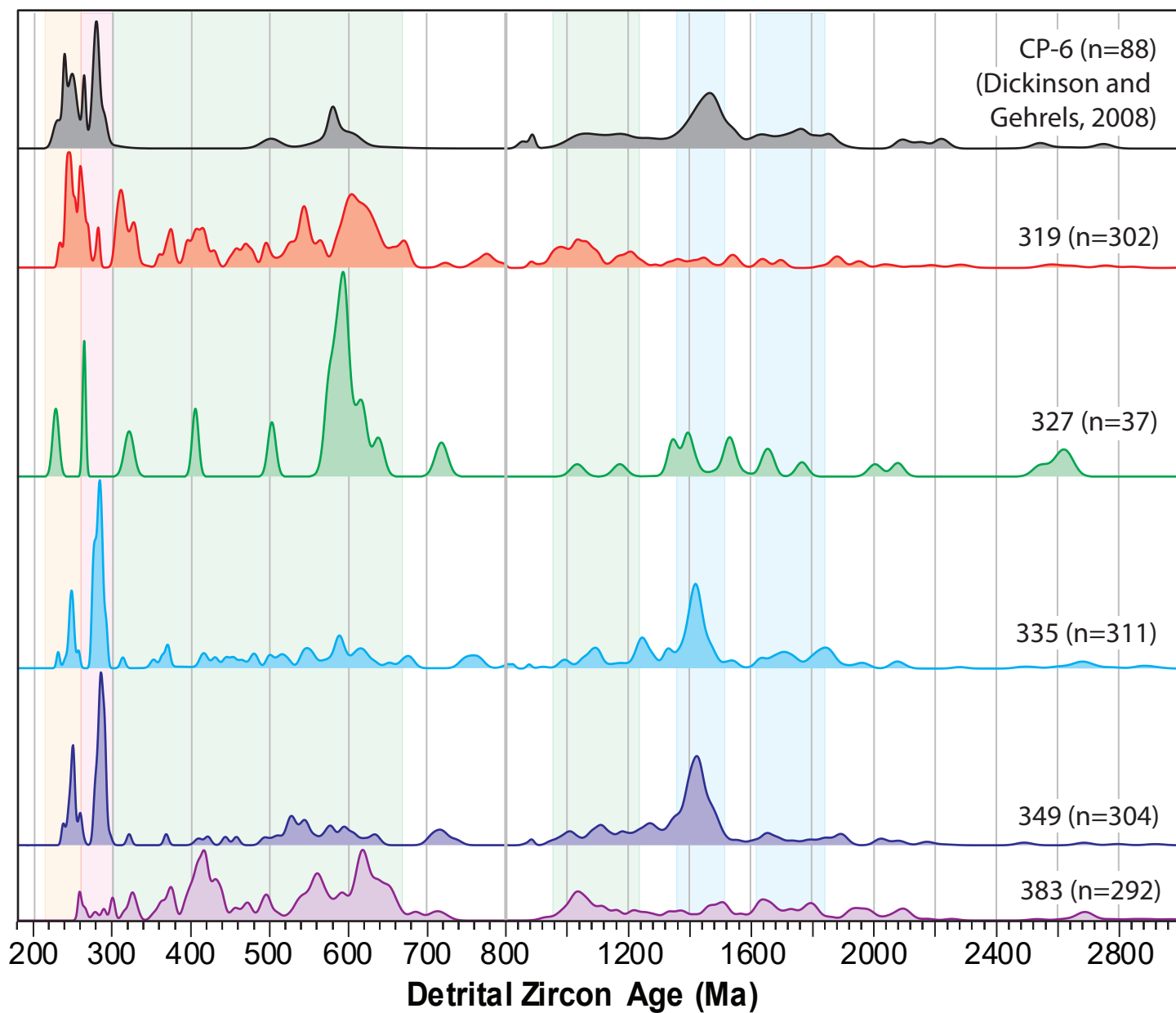


Figure 7 (Chinle PDP)

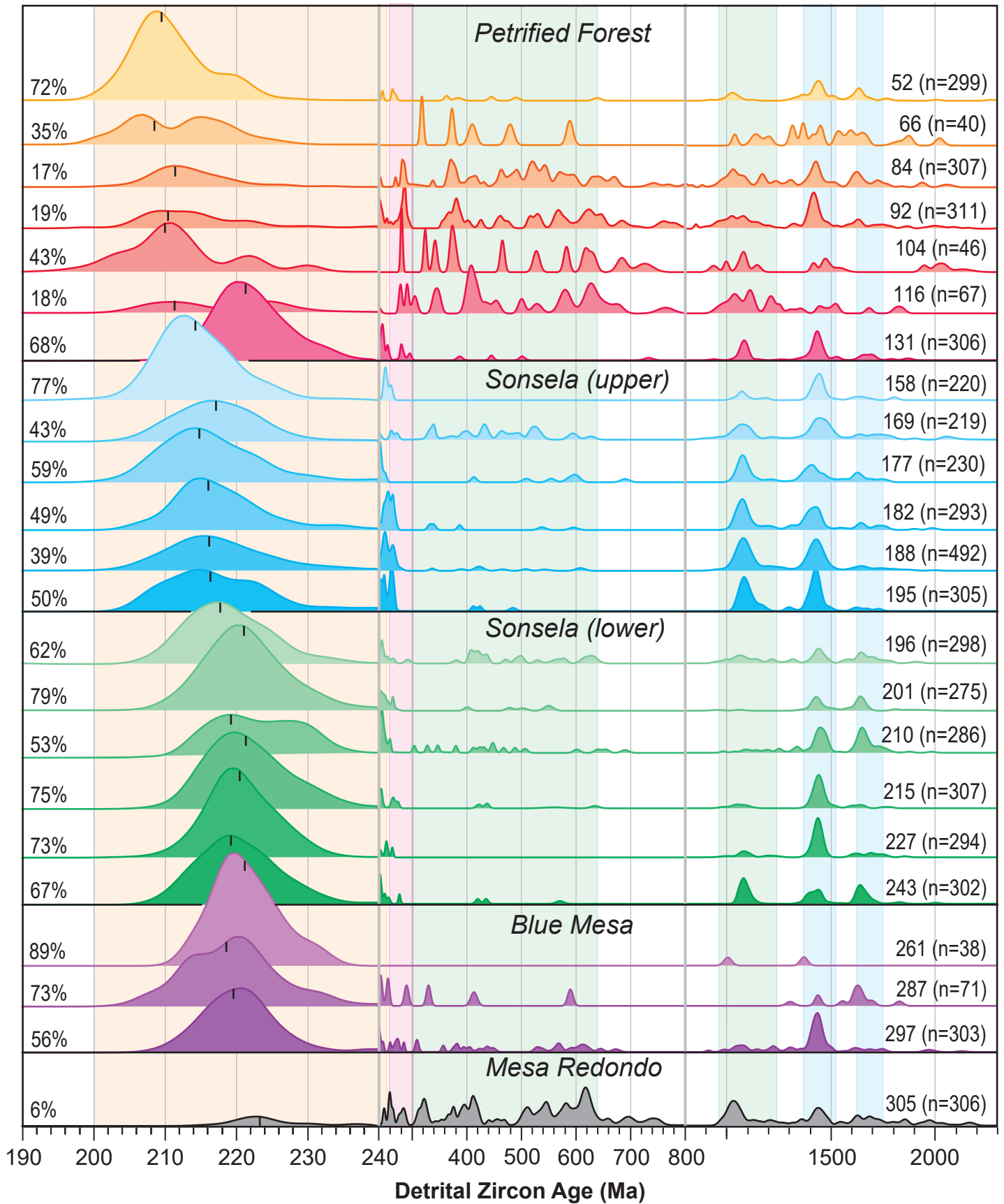


Figure 8 (Coco-Moen-Chin PDP)

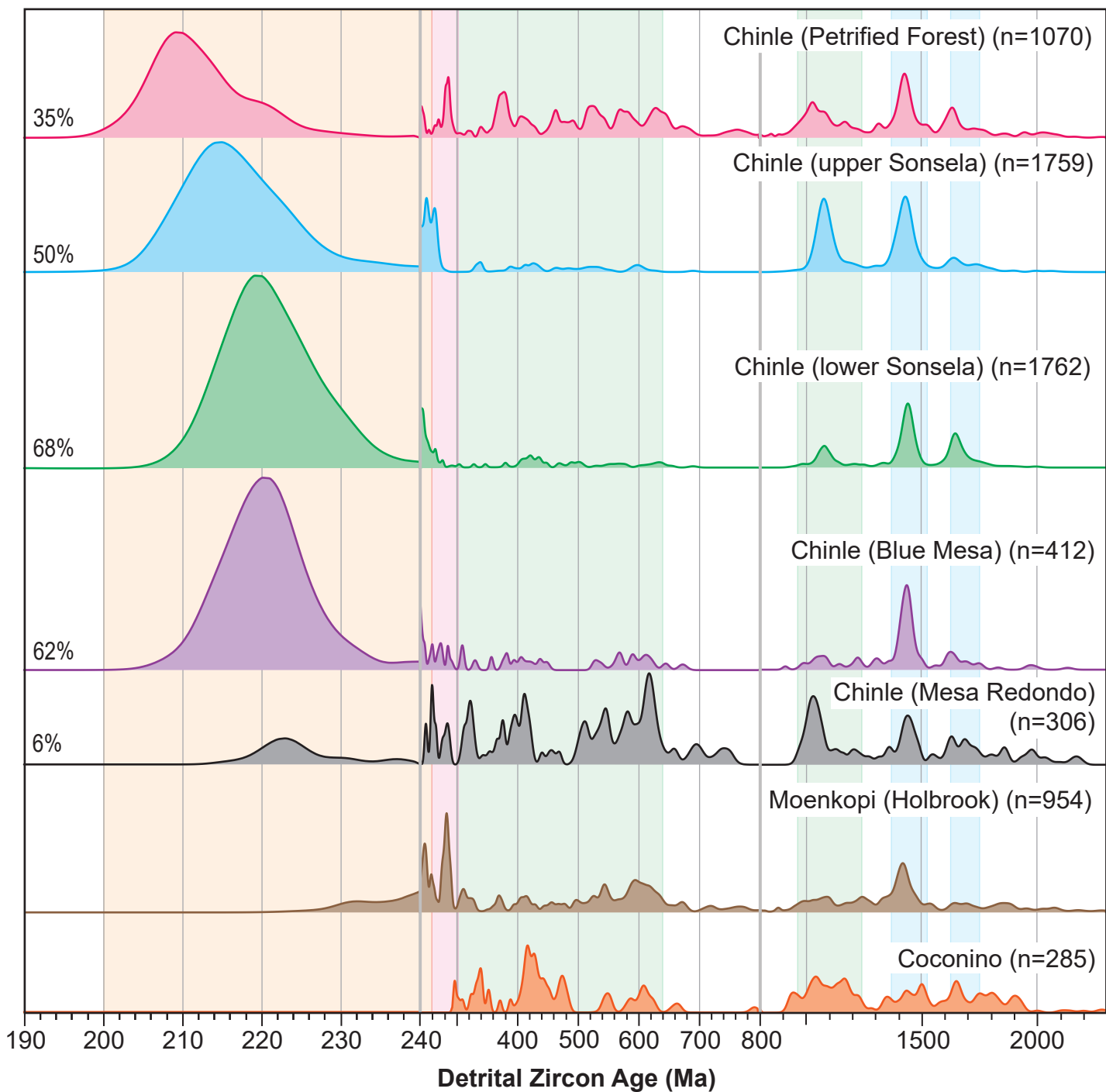


Figure 9 (MDS Plots)

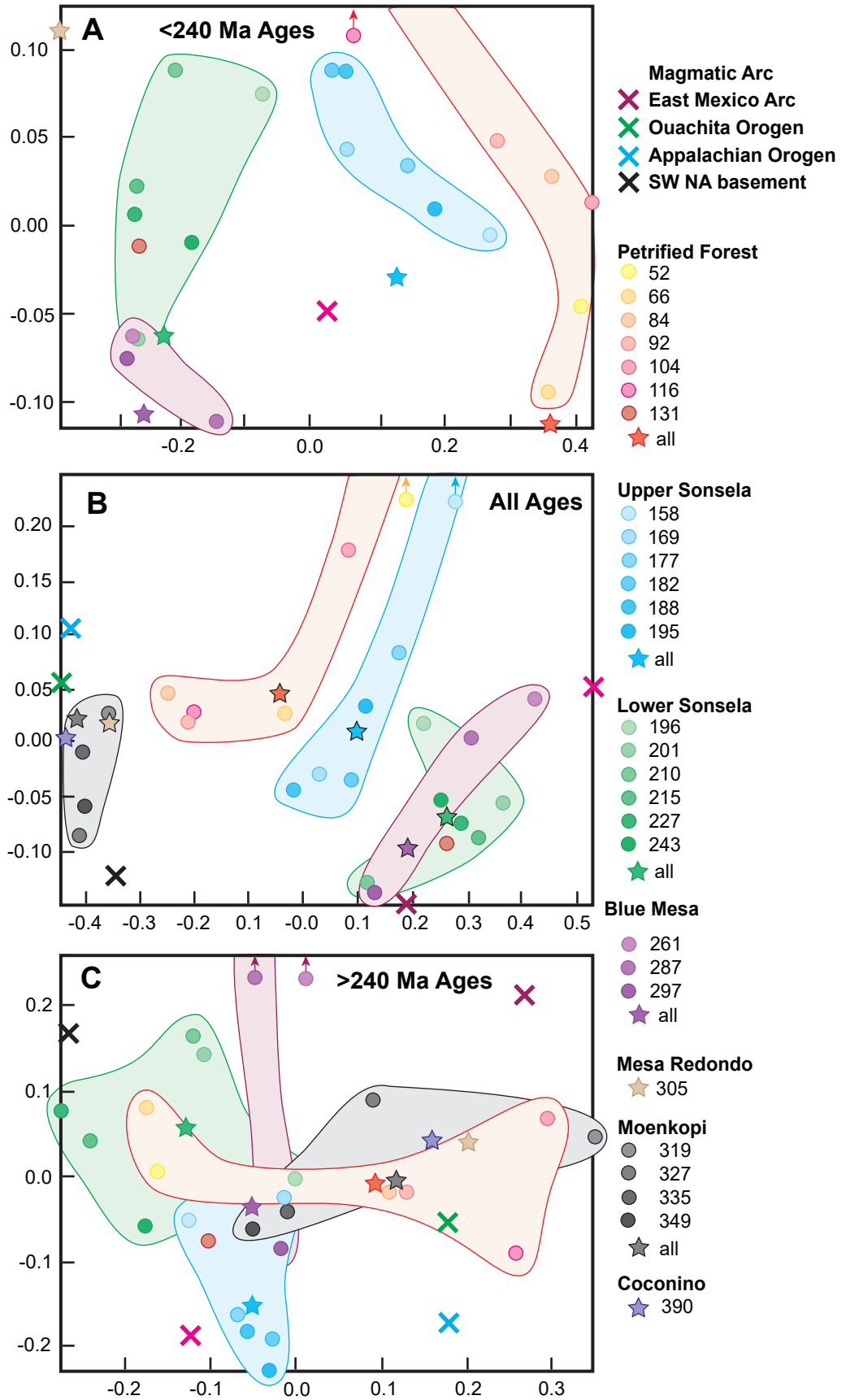


Figure 10 (Uconc-UTh plot)

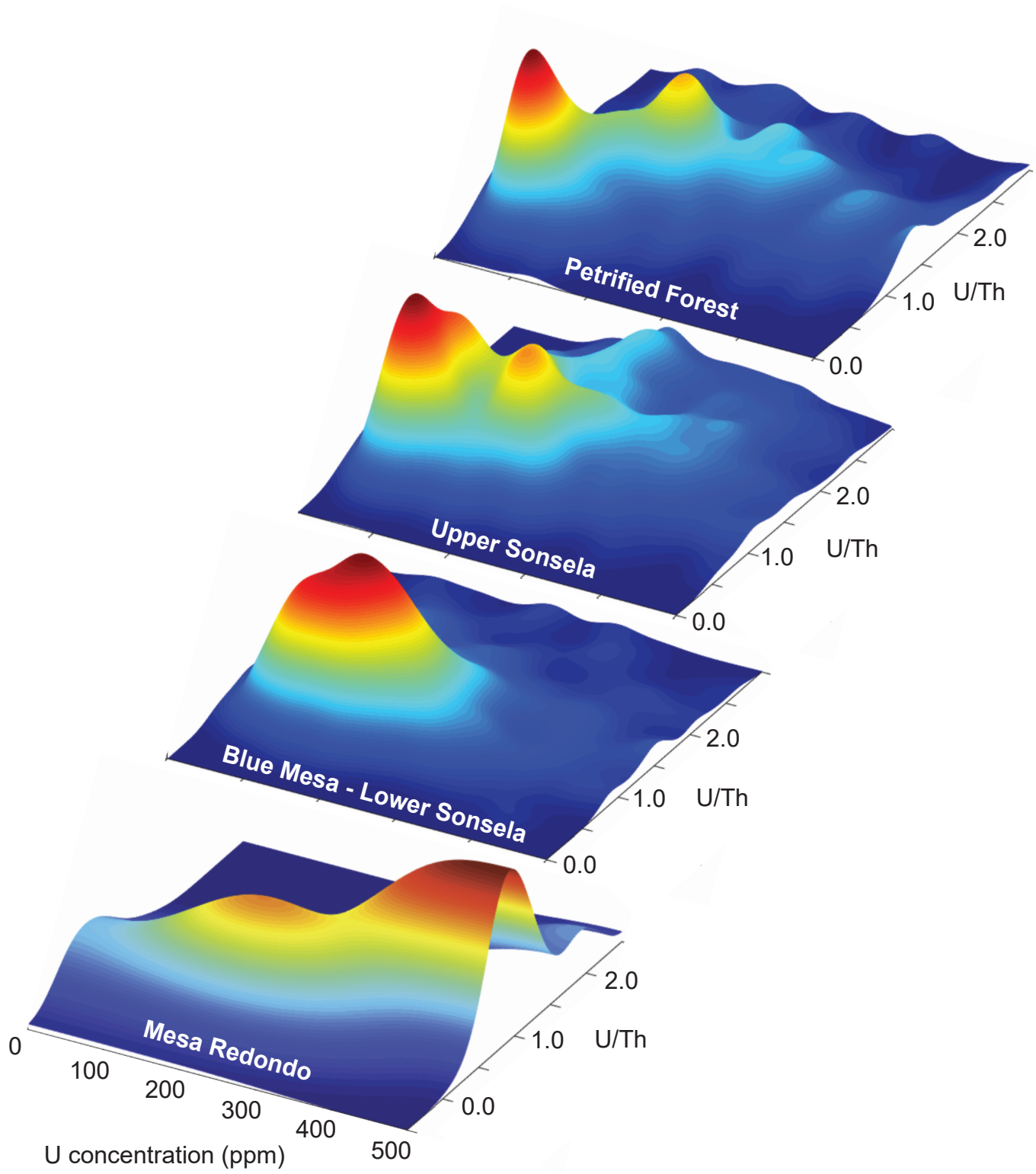


Figure 11 (AOB CO MDS plot)

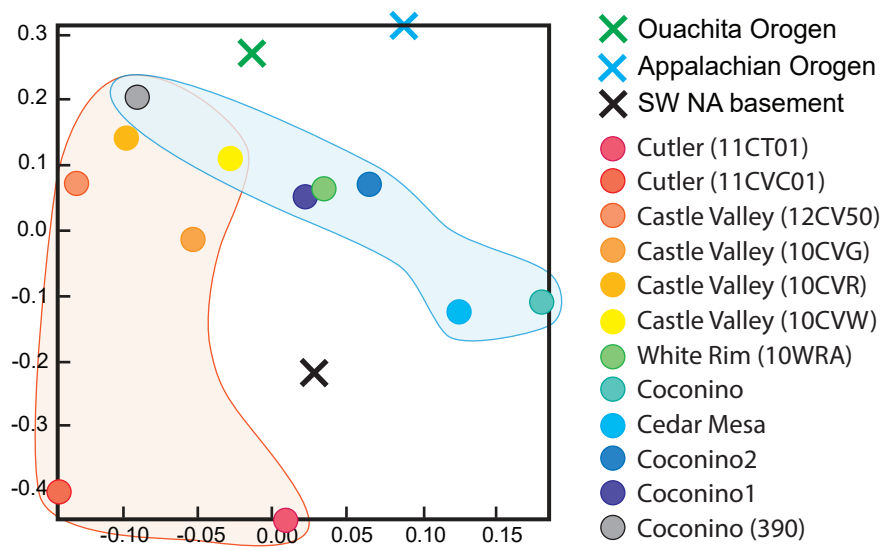


Figure 12 (Triassic Paleogeography)

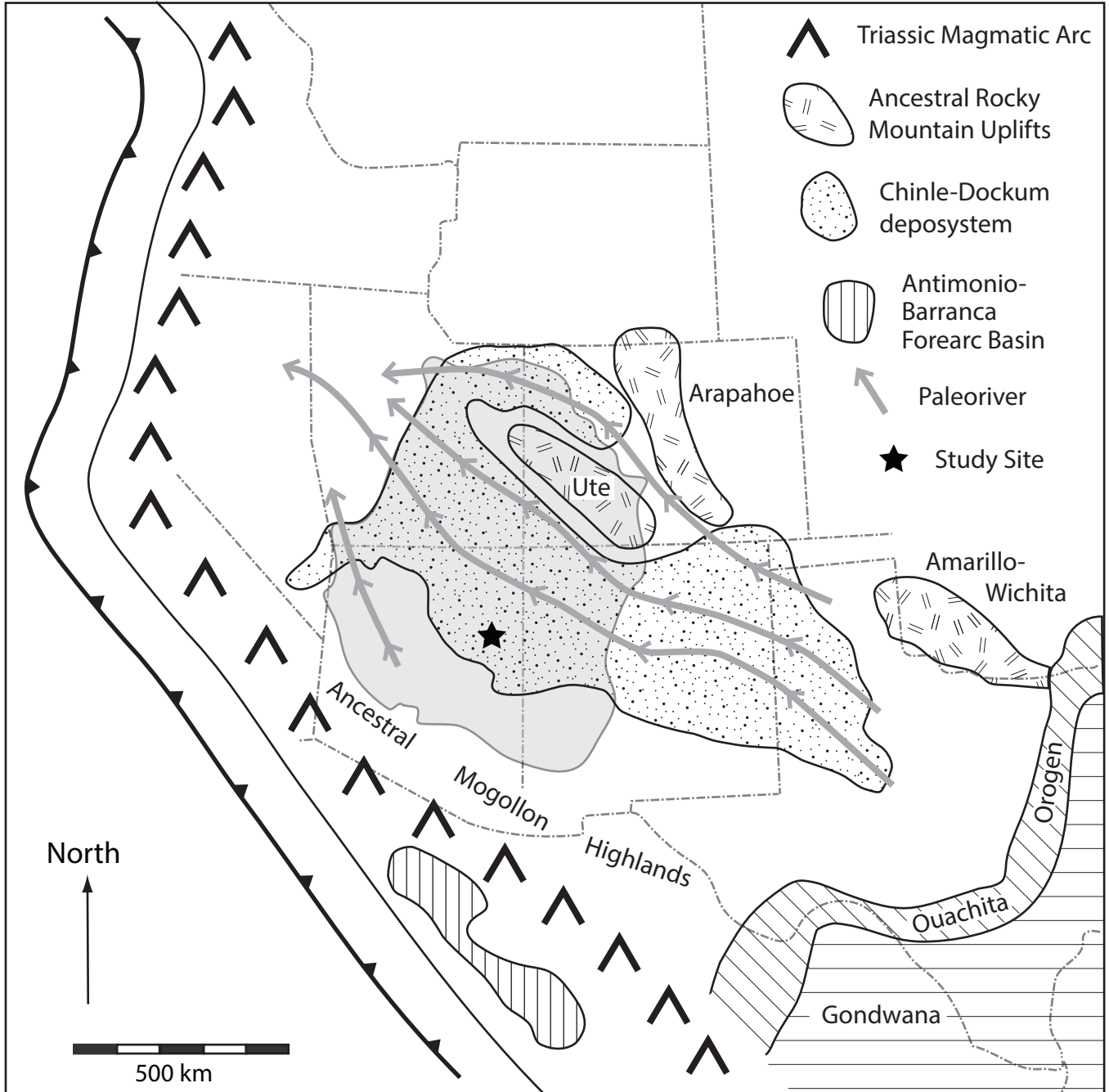


Figure 13 (DZ MDA plot)

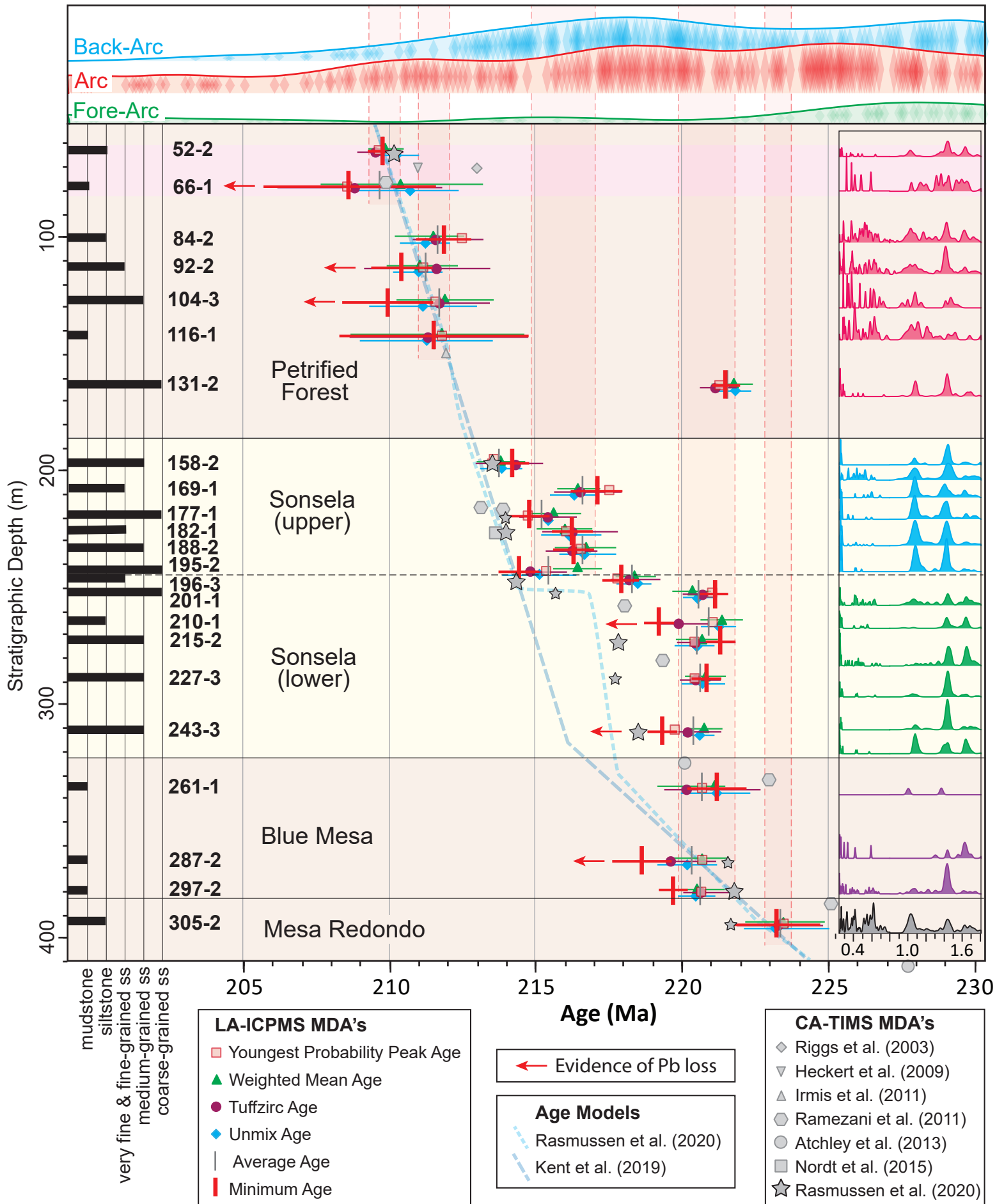
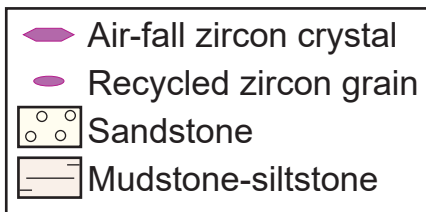
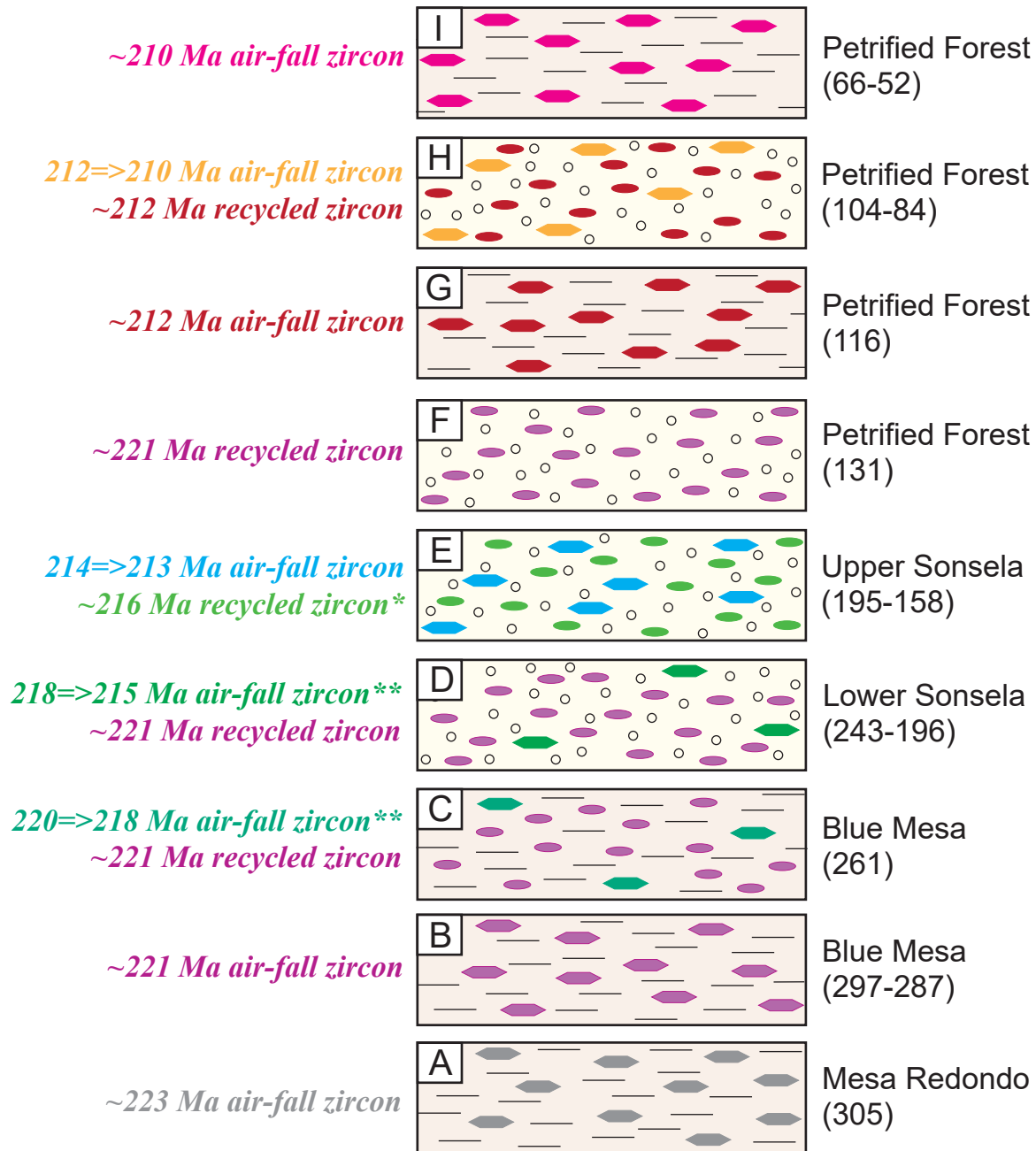


Figure 14 (Chinle Strat)



**Uncertain source for ~216 Ma recycled grains in upper Sonsela*

***Low abundance of 220-215 Ma air-fall grains in lower Sonsela*