

1 *In-situ* Lu – Hf geochronology of calcite

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7 **Abstract.** The ability to constrain the age of calcite formation is of great utility to the Earth Science community, due to the
8 ubiquity of calcite across a wide spectrum of geological systems. Here, we present the first in-situ laser ablation inductively
9 coupled tandem quadrupole mass spectrometry (LA-ICP-MS/MS) Lu–Hf ages for calcite, demonstrating geologically
10 meaningful ages for IOCG and skarn mineralisation, carbonatite intrusion and low grade metamorphism. The analysed samples
11 range in age between ca. 0.9 Ga and ca. 2 Ga with uncertainties between 1.7% and 0.6% obtained from calcite with Lu
12 concentrations as low as ca. 0.5 ppm. The Lu–Hf system in calcite appears to be able to preserve primary precipitation ages
13 over a significant amount of geological time, although further research is required to constrain the closure temperature. The
14 in-situ approach allows calcite to be rapidly dated while maintaining its petrogenetic context with mineralization and other
15 associated mineral processes. Therefore, LA-ICP-MS/MS Lu–Hf dating of calcite can be used to resolve the timing of complex
16 mineral paragenetic sequences that are a feature of many ancient rock systems.

17 **1 Introduction**

18 Calcite (CaCO₃) is the main mineral phase of most carbonate sedimentary rocks and their metamorphic equivalents, is a
19 common diagenetic phase and is a major component of carbonatites. Calcite is also a common product of hydrothermal
20 alteration and constituent of mineralising systems where it may precipitate from fluids during pre-ore, ore-stage, and post-ore
21 forming processes (Debruyne et al., 2016). The ability to directly date calcite unlocks the possibility to constrain the timing of
22 a vast array of geological processes that can be difficult to date using conventional methods.

23
24 Accurate *in-situ* U–Pb geochronology of calcite has been applied to a variety of geological systems (e.g; Li et al., 2014; Ring
25 and Gerdes, 2016; Roberts and Walker, 2016). However, calcite often incorporates significant quantities of Pb during
26 crystallisation (i.e. ‘initial’ or ‘common’ Pb), which can limit the utility of U–Pb geochronology (Rasbury and Cole, 2009).
27 Moreover, Pb is highly fluid mobile (Brugger et al., 2016), so it can be difficult to obtain primary age information with the U–
28 Pb method in hydrothermal or strongly-altered systems (Roberts et al., 2020; Simpson et al., 2021b). Further, given the
29 propensity for calcite to undergo recrystallisation, calcite U–Pb geochronology is rarely applicable to Precambrian systems as
30 the calcite U–Pb system invariably does not remain closed over long timescales (Whitehouse and Russell, 1997).

31

32 Alternative dating systems involving the radioisotopic decay of rare earth elements (REE) such as Sm–Nd and Lu–Hf, have
33 previously been applied to calcite (e.g. Barker et al., 2009; Maas et al., 2020; Nie et al., 1999; Peng et al., 2003), based on the
34 moderate to strong compatibility of REEs in carbonates in many systems (Debruyne et al., 2016; Elzinga et al., 2002; Terakado
35 and Masuda, 1988; Zhong and Mucci, 1995). Although it should be noted that REE compatibility will be dependent on the
36 conditions of calcite formation, and can vary. Importantly for geochronology, experimental evidence indicates that Lu and Hf
37 are highly immobile in many hydrothermal fluids (Brugger et al., 2016; Migdisov et al., 2016), meaning that the Lu–Hf system
38 is potentially during post-formation processes relative to the U–Pb system. However, concentrations of Lu and Hf are generally
39 low (ppm to ppt range) in calcite, necessitating the dissolution of large quantities of material (up to 2g) per sample for
40 conventional Lu–Hf geochronology (Maas et al., 2020). These large quantities significantly reduce the spatial resolution of the
41 technique and have the additional problem of potential contamination from inclusions. Furthermore, age variation is difficult
42 to detect, and bulk samples may produce meaningless average age derived from mixing of age domains.. The dissolution
43 process also removes calcite from its petrological context. The recent development of *in-situ* Lu–Hf geochronology of
44 individual minerals by LA-ICP-MS/MS allows for rapid acquisition of spatially resolved data, and has been demonstrated for
45 garnet (Ribeiro et al., 2021; Tamblyn et al., 2021) and apatite (Glorie et al., 2021)

46

47 In this study, we present the first *in-situ* Lu–Hf dating of calcite from a variety of geological environments. We demonstrate
48 the that *in-situ* calcite Lu–Hf geochronology can produce meaningful ages for complexly deformed and hydrothermally-altered
49 systems, such as mineral deposits, as well as carbonatite intrusions and low grade metamorphism.

50 **2 Geological Background of Samples**

51 The analysed samples were selected; (1) to demonstrate that calcite Lu–Hf can date primary calcite formation in carbonatites;
52 (2) to reveal the potential of the method to unravel complex ore systems or later events, and; (3) to characterize large calcite
53 samples that would make suitable reference materials for *in-situ* analysis.

54 **2.1 Phalaborwa Carbonatite, South Africa**

55 The Phalaborwa Igneous Complex is located ~450 km northeast of Johannesburg, in the Limpopo Province, South Africa. The
56 igneous complex is the result of several distinct pulses of alkaline intrusions that were emplaced into Archean granitic gneiss
57 (Staff, 1976). The Loolekop pipe is located in the centre of the Phalaborwa Igneous Complex and was intruded by two episodes
58 of carbonatite emplaced at the intersection of five major faults and shear zones (Basson et al., 2017; Staff, 1976). The oldest
59 carbonatite is termed the “transgressive banded” carbonatite and has an emplacement age of 2060.0 ± 2.2 Ma (baddeleyite
60 SIMS U-Pb; Wu et al., 2011). This is intruded by a slightly younger carbonatite termed the “banded” carbonatite and has an
61 emplacement age of 2059.8 ± 1.3 Ma (baddeleyite SIMS U-Pb; Wu et al., 2011). The Phalaborwa carbonatite is unique as it is
62 the only known example of a carbonatite containing economic Cu mineralisation (Groves and Vielreicher, 2001). In the banded
63 carbonatite-phoscorite, Cu mineralisation is primarily in the form of bornite inter-grown with valleriite with minor chalcopyrite

64 (Staff, 1976). In the transgressive carbonatite, Cu mineralisation is present as chalcopyrite inter-grown with cubanite and
65 valleriite (Staff, 1976). Cu mineralisation is interpreted to be magmatic-hydrothermal in origin, with Cu leached by high-
66 temperature hydrothermal fluids at depth, precipitating along fractures within the hosting carbonatite (Le Bras et al., 2021).
67 The sample used in this study (P01) is representative of carbonatite hosted Cu-mineralisation from within the Loolekop pipe
68 (Fig. 1). The sample is mineralogically composed of chalcopyrite inter-grown with cubanite and pyrrhotite alongside an
69 assemblage of magnetite, dolomite, calcite, biotite, pyroxene and valleriite. As the Phalaborwa carbonatite has a well
70 constrained crystallisation age, it provides an ideal case study to demonstrate the utility of the *in-situ* Lu–Hf method for dating
71 igneous calcite directly associated with Cu mineralisation.

72 **2.2 The Eastern Fold Belt, Mt Isa Block, Queensland, Australia**

73 The Eastern Fold Belt of the Mount Isa Domain has experienced multiple episodes of deformation, magmatism,
74 metamorphism, mineralisation and pervasive hydrothermal alteration across the Paleo- to Mesoproterozoic, and hence
75 represents one of the most metasomatized crustal blocks on Earth (Oliver et al., 2008). Hydrothermal calcite is common across
76 the Mount Isa region, in the Mary Kathleen Domain (Oliver et al., 1993) and in many of the IOCG deposit of the Cloncurry
77 District. For this study we have selected calcite samples from the Lime Creek calcite quarry and the Mt Elliott IOCG deposit
78 for Lu–Hf analysis. The Mt Isa Domain has both regional and deposit level age constraints, making it a good area to
79 demonstrate the technique.

80

81 The Lime Creek quarry is one of a number of large calcite pods or veins that are exposed in Mary Kathleen Domain. The Lime
82 Creek quarry is hosted within the ca. 1760 Ma Argylla Formation and lies along the steeply dipping NNW-trending Tribulation-
83 Lime Creek Fault, which offsets regional-scale ‘D₂’ folds (Marshall, 2003). Breccias along this fault contain clasts of calc-
84 silicate rocks and metadiorite with a matrix consisting of albite-actinolite-diopside-biotite-titanite-apatite that are subsequently
85 overprinted by the undeformed Lime Creek calcite-dominated veins (Marshall, 2003). These veins are extremely coarse-
86 grained with calcite crystals larger than 1 m³, actinolite crystals over 1 m in length, and apatite, biotite, diopside and titanite
87 grains over 20 cm in diameter (Marshall, 2003; Oliver et al., 1993). Based on cross-cutting relationships, it is interpreted that
88 the Lime Creek vein system, and other calcite pods/veins of this style precipitated post-faulting during late-‘D₃’ deformation
89 (ca. 1550–1500 Ma) of the Isan Orogeny (Giles and Nutman, 2002; Marshall, 2003). This style of veining is common
90 throughout the Mary Kathleen Domain and provides evidence of km-scale fluid transport during late-stage metamorphism
91 (Oliver et al., 1993). Based on C and O isotope analysis of calcite from these veins, they are interpreted to have formed from
92 hydrothermal fluids likely associated with the intrusion of the ca. 1530 to 1500 Ma Williams-Naraku Batholiths (Oliver et al.,
93 1993). Although no direct dating has been completed on the Lime Creek Quarry, titanite from the nearby and cognate Knobby
94 Quarry have produced three titanite U–Pb ages of 1521 ± 5 Ma, 1527 ± 7 Ma, and 1555 ± 5 Ma (Oliver et al., 2004). The
95 sample analysed in this study (LC1) consists of very coarse-grained calcite with coarse-grained diopside collected from a large
96 calcite pod in the Lime Creek Quarry (Fig. 1).

98 Mt Elliott is an IOCG deposit located in the Eastern Fold Belt of the Mount Isa Inlier (Duncan et al., 2011). The deposit is
99 situated within northwest striking splays of the Mount Dore Fault (Duncan et al., 2011; Wang and Williams, 2001), and is
100 hosted within skarn-altered and deformed phyllites and schists (Garrett, 1992; Wang and Williams, 2001). The host rocks were
101 metamorphosed to lower amphibolite facies during the ca. 1600-1580 Ma 'D₂' deformation of the Isan Orogeny (Garrett, 1992;
102 Wang and Williams, 2001). The formation of early albite-hematite (red rock) alteration enhanced brittle fracturing and
103 brecciation of the shale (Garrett, 1992). This was infilled by two stages of open-space skarn development: (1) diopside-
104 magnetite-hematite-calcite-titanite-allanite-phlogopite, and: (2) actinolite-scapolite-magnetite-andradite-calcite-epidote-
105 allanite-chlorite and biotite (Garrett, 1992; Wang and Williams, 2001). Sulphides in the second stage include chalcopyrite,
106 pyrrhotite and pyrite (Garrett, 1992). Although the two skarn assemblages are difficult to distinguish mineralogically, the
107 second episode is the most widespread and represents the major Cu–Au event (Garrett, 1992; Wang and Williams, 2001). A
108 variety of geochronological techniques have been applied to constrain the age of Cu–Au mineralisation. The earliest phase of
109 skarn development has been dated at 1530 ± 11 Ma (U–Pb titanite; Duncan et al., 2011). The second stage of skarn development
110 associated with Cu–Au mineralisation has been dated at 1513 ± 5 Ma (molybdenite Re–Os; Duncan et al., 2011) and $1510 \pm$
111 3 Ma (actinolite Ar–Ar; Wang and Williams, 2001). Two outcrop samples from the Mount Elliot Cu–Au deposit were selected
112 for Lu–Hf geochronology. Mt Elliott 1 (ME 1) consists of coarse-grained pink-coloured calcite that is coeval with the formation
113 of diopside, scapolite and magnetite (Fig. 1). Although the paragenesis of this sample is relatively unconstrained, the lack of
114 sulphides may indicate that this sample belongs to the early pre-mineralisation skarn assemblage. Calcite from sample Mt
115 Elliott 2 (ME 2) is coeval with the formation of andradite, pyrite, chalcopyrite, pyrrhotite and magnetite (Fig. 1). The close
116 relationship between calcite and chalcopyrite in this sample indicates that it is associated with the main Cu–Au bearing skarn
117 assemblage.

118 **2.3 Flin Flon Volcanic Massive Sulphide (VMS) Deposit, Canada**

119 The Flin Flon Greenstone Belt stretches across central Manitoba through to east central Saskatchewan and hosts several world-
120 class Zn–Cu VMS deposits including the Flin Flon, Callinan and 777 deposits (Koo and Mossman, 1975). Zn–Cu
121 mineralisation is interpreted to have formed contemporaneously with deposition of the 1888.9 ± 1.6 Ma Millrock member
122 during the Trans-Hudson Orogeny (Gibson et al., 2012; Koo and Mossman, 1975; Rayner, 2010). The Flin Flon Zn–Cu orebody
123 is recognised to have undergone six distinct deformation events that have affected the shape of the deposit (Lafrance et al.,
124 2016; Schetselaar et al., 2017). 'D₁' and 'D₂' were associated with the intra-oceanic accretion of the Flin Flon Arc to other
125 volcanic terrains before ca. 1872 Ma (Lafrance et al., 2016). 'D₃' occurred from 1847–1842 Ma as a response to the final
126 accretion of the Flin Flon Terrane to the Glennie Terrane, producing west-verging folds within stacked, east dipping thrust
127 sheets of basement and cover rocks bounded by NNW-striking thrust faults (Lafrance et al., 2016). 'D₄' resulted from the
128 collision between the Flin Flon and Glennie complex with the Sask Craton and is broadly coeval with the ca. 1840 Ma Phantom
129 Lakes dyke (Gibson et al., 2012; Lafrance et al., 2016). 'D₅' deformation produced a penetrative regional cleavage (S₅) that is

130 defined by a continuous chloritic foliation ubiquitous in the volcanic basement rocks (Gibson et al., 2012; Lafrance et al.,
131 2016). EWE–WNW directed compression during ‘D₆’ deformation produced a second regional penetrative cleavage and
132 reactivated a variety of regional-scale faults (Gibson et al., 2012; Lafrance et al., 2016). Regional greenschist to granulite facies
133 metamorphism is associated with D_{5–6} deformation at ca. 1820–1790 Ma (Schneider et al., 2007). The Flin Flon mine horizon
134 was imbricated during ‘D₃’ thrusting with the shape of the ore lenses moulded during ‘D₄’ and ‘D₅’ deformation (Schetselaar
135 et al., 2018). Regional greenschist to amphibolite grade metamorphism occurred between 1820 – 1790 Ma (U-Pb monazite;
136 Schneider et al., 2007), with rocks in the Flin Flon deposit reaching greenschist facies (Koo and Mossman, 1975). The sample
137 selected for this study is from the hydrothermally altered and sheared footwall of the Flin Flon VMS deposit. This sample is
138 composed of highly foliated chlorite and calcite with disseminated pyrite and residual titanomagnetite. A band of highly
139 foliated calcite was selected for Lu–Hf analysis (sample FF014; Fig. 1) to constrain the age of syn-metamorphic shearing of
140 the deposit.

141 **2.3 Yates U-Th prospect, Otter Lake Area, Grenville Province, Canada**

142 The Otter Lake area is located in SE Ontario within the Grenville Province. The Grenville Province can be distinguished from
143 surrounding provinces based on various structural, metamorphic and isotopic signatures attributed to the overprinting ca. 1080–
144 980 Ma Grenvillian Orogeny (Rivers, 2015). This orogenic event produced widespread metamorphism from granulite to
145 amphibolite facies (Breemen and Corriveau, 2005), accompanied by widespread hydrothermal alteration in the Otter Lake area
146 (Kretz et al., 1999). The Yates U-Th prospect is located approximately 100 km northwest of Ottawa, and is renowned for the
147 occurrence of pegmatites that contain large euhedral crystals of apatite set within a matrix of predominantly orange-pink
148 calcite, with diopside, allanite, titanite, fluorite, thorite and phlogopite (Schumann et al., 2019). A wide range of dates have
149 been produced from the Yates mine; including titanite Pb–Pb and U–Pb ages between ca. 1020 and 998 Ma (Frei et al., 1997;
150 Kennedy et al., 2011); apatite Pb–Pb, and U–Pb ages of 913 ± 7 Ma (Barfod et al., 2005), 933 ± 12 Ma and 920 – 850 Ma
151 (Chew et al., 2011; Xiang et al., 2021) respectively, and an apatite Lu–Hf age of 1031 ± 6 Ma (Barfod et al., 2005). In addition,
152 Simpson et al. (2021a) obtained an *in-situ* Lu-Hf apatite age of 1000 ± 11 Ma (when corrected for laser induced elemental
153 fractionation). Importantly, the apatite Lu–Hf and Pb–Pb ages have been obtained from the same large apatite crystal,
154 indicating that the Lu–Hf and U–Pb systems have been decoupled (as opposed to multiple generations of apatite growth).
155 Barfod et al. (2005) argued that late-stage fluid interactions may have affected Pb retentivity in the apatite, as the apatite was
156 unlikely to be above the apatite Pb closure temperature at ca. 913 Ma. Calcite from a specimen containing coarse-grained
157 euhedral apatite with pink calcite, quartz and diopside was selected for calcite Lu–Hf analysis (OL-MB, Fig. 1). The apatite is
158 enclosed in the sampled calcite, and is interpreted to have crystallised just prior to the calcite, but during the same hydrothermal
159 event.

160 3 Method

161 The samples were mounted in 2.5 cm diameter epoxy mounts and screened for Lu concentration by LA-ICP-MS to determine
162 suitability for Lu–Hf analysis. Mineral liberation analysis (MLA) maps were obtained using a Hitachi SU3800 scanning
163 electron microscope (SEM) to reveal the petrogenetic context of the analysed calcite.

164 3.1 In-Situ Lu-Hf Dating Method

165 Analysis was conducted at Adelaide Microscopy, The University of Adelaide. Calcite samples were analysed using a
166 RESOLution 193 nm laser ablation system (Applied Spectra) with a S155 sample chamber (Laurin Technic). The laser ablation
167 system was coupled to an Agilent 8900 tandem mass spectrometer (ICP-MS/MS). The methodology largely follows that of
168 Simpson et al. (2021a) including an initial instrument tune conducted with no NH₃ in the reaction cell to achieve robust plasma
169 conditions (U/Th = 1.00-1.05) and minimal oxide interferences (ThO/Th <0.2%). A carrier gas of 3.5 mL/min N₂ was added
170 after the sample cell in order to increase sensitivity (Hu et al., 2008). Analytical conditions are included in appendix Table 1.

171
172 Methods for separation of ¹⁷⁶Hf from ¹⁷⁶Lu and ¹⁷⁶Yb follow that of Simpson et al. (2021a). In more detail, the Agilent 8900x
173 utilises a reaction cell between two quadrupole mass analysers, which can be used to separate isobaric interferences. The first
174 quadrupole is used as a mass filter (e.g., when set to mass 176, only ¹⁷⁶Lu, ¹⁷⁶Yb, and ¹⁷⁶Hf can pass), thereby minimizing
175 potential background interferences and other, unwanted reactions. Following this, a mixture of 10% NH₃ and 90% He is added
176 to the reaction cell (at a rate of 3 mL/min). This mixture is optimized to promote formation of the Hf reaction product
177 Hf((NH)(NH₂)(NH₃)₃)⁺ and the second quadrupole is set to 82 amu higher than the first (e.g., Q1 = 176 amu and Q2 = 258
178 amu). This method minimizes the equivalent Lu and Yb reaction products (~0.03% for Lu, and below detection for Yb), such
179 that the isobaric interferences on ¹⁷⁶Hf are negligible (Simpson et al., 2021a). Lens voltages were tuned to increase sensitivity
180 on the Hf reaction product (Simpson et al., 2021a). In order to calculate Lu/Hf ratios, ¹⁷⁶Hf (+82) was measured directly, ¹⁷⁵Lu
181 was measured as a proxy for ¹⁷⁶Lu, and ¹⁷⁸Hf (+82) measured as a proxy for ¹⁷⁷Hf (Simpson et al., 2021a). ¹⁷⁶Hf/¹⁷⁶Lu,
182 ¹⁷⁶Lu/¹⁷⁷Hf, and ¹⁷⁶Hf/¹⁷⁷Hf ratios were calculated as part of the normalization to NIST610, as opposed to separately converting
183 measured ¹⁷⁵Lu and ¹⁷⁸Hf into ¹⁷⁶Lu and ¹⁷⁸Hf. In more detail, if we assume that the ¹⁷⁶Lu/¹⁷⁵Lu ratio (or ¹⁷⁷Hf/¹⁷⁸Hf ratio) is
184 identical between NIST SRM 610 and all analysed samples, a correction factor calculated from the % difference between the
185 ¹⁷⁵Lu/¹⁷⁸Hf ratio measured in NIST SRM 610 and the published ¹⁷⁶Lu/¹⁷⁷Hf will correct the unknowns for matrix independent
186 fractionation and differences in isotopic abundance. ⁴³Ca was measured for internal normalization of trace element abundances,
187 and the following isotopes were measured to monitor for inclusions: ²⁷Al, ⁴⁷Ti, ⁸⁹Y, ⁹⁰Zr, ¹⁴⁰Ce, and ¹⁷²Yb.

188

189 Lutetium abundances in most calcite samples are low (< 6 ppm), so we employed a large laser diameter of 257 μm, and a
190 repetition rate of 10 Hz to maximize sensitivity. High sensitivity is important in order to either measure common Hf (in this
191 case ¹⁷⁸Hf), or demonstrate that ¹⁷⁶Hf is sufficiently above detection limits that the effects of common Hf are negligible. Smaller

192 spot sizes could be employed for higher Lu and/or higher Hf samples. An extra 20 seconds delay was added after each interval
193 of sample ablation in order to ensure the washout had reached background levels. NIST SRM 610 glass ($^{176}\text{Lu}/^{177}\text{Hf}$: 0.1379
194 \pm 0.005, $^{176}\text{Hf}/^{177}\text{Hf}$: 0.282122 \pm 0.000009; Nebel et al., 2009) was used as the primary reference material and was analyzed
195 using a spot size of 43 μm . The smaller spot size was required to ensure that ^{175}Lu was measured in pulse counting mode (<4
196 Mcps). Consistent with observations in Simpson et al. (2021a), Lu and Hf showed no measureable down-hole fractionation in
197 the analysed carbonates (fig. 2), as such, no down-hole correction was applied to the data.

198

199 A side effect of the use of large ablation spots is ‘plasma loading’ for which the introduction of a large amount of material
200 reduces the ionizing efficiency of the plasma (Kroslakova and Günther, 2007). Plasma loading was observed in the time
201 resolved signals, with a reduction in signal intensity for all isotopes after ~10 to 15 seconds of ablation. Following this, the
202 signal stabilized after ~18 seconds of ablation (Fig. 2). Importantly, this variation in signal intensity was not observed in the
203 calculated time resolved isotope ratios (Fig. 2), which means that identical ratios were calculated whether this decrease in
204 signal intensity was included or not in the ratio calculation. Importantly, plasma loading can be affected by sample matrix
205 (Kroslakova and Günther, 2007), especially for minerals containing easily ionized elements such as Ca. This necessitates
206 matrix matched calibration, despite the observed lack of down-hole changes in Lu–Hf ratios (Simpson et al., 2021a).

207

208 The large ablation volume caused accumulation of ablated material in the tubing and on the interface cones during the first
209 analytical session, which coincided with a decrease in signal intensity over time. Consequently, session 1 records slightly more
210 signal drift compared to session 2. However, there was no measureable corresponding drift observed in the calculated isotopic
211 ratios, apart from a slight decrease in precision due to the lower sensitivity toward the end of the run. Therefore, we recommend
212 that cones are cleaned prior to analysis, and suggest a maximum session duration of approximately 7 hours when using spot
213 diameters of >200 μm . In addition to this, the accumulated material was sometimes mobilized in later analyses, potentially
214 contaminating data. This was observed by increases in Al during the start of ablation that decayed down to background levels.
215 Importantly, similar Al spikes were not observed during background measurement, indicating contamination due to material
216 remobilized during ablation is likely, hence why the additional 20 seconds of washout did not fix this. This contamination did
217 not generally produce a measurable effect on calculated Lu/Hf ratios. However, we stress that this contamination is important
218 to monitor as Hf concentrations are sometimes in the ppt level. As such we recommend close monitoring of signals, particularly
219 Al concentrations, and the removal of 1-3 seconds of each analysis after signal stabilisation if necessary.

220

221 3.2 Data Processing

222 For both LA-ICP-MS and LA-ICP-MS/MS analysis, a stoichiometric Ca concentration of 40.04 wt.% for calcite was used for
223 internal normalization of trace element concentrations. Although the high Ca cps for all analysed samples indicate that they
224 are close to stoichiometric calcite, there may be slight inaccuracies in calculated element concentrations due to major element

225 substitutions from Mg, Fe, and Mn that are common in carbonates. However, element concentrations were largely used as
226 relative proxies to monitor for inclusions.

227

228 Background subtractions, element concentrations and ratio calculations were performed using LADR software (Norris and
229 Danyushevsky, 2018). Where ^{178}Hf was measured above detection limits (~ 2 ppt for ^{178}Hf), common Hf corrections were
230 applied to the data after background subtractions, but prior to normalization to the standard. In more detail, the ^{178}Hf cps
231 measurement for each sampling cycle of the analysis period of each laser spot was used to calculate the common Hf component
232 of the corresponding ^{176}Hf cps measurement, using the following equation:

$$233 \quad i^{176}\text{Hf}_r = i^{176}\text{Hf}_m - \left(\frac{i^{176}\text{Hf}}{i^{178}\text{Hf}} \right) c \times i^{178}\text{Hf}_m$$

234 Where $^{176}\text{Hf}_r$ = radiogenic ^{176}Hf , $^{176}\text{Hf}_m$ = measured ^{176}Hf , $^{178}\text{Hf}_m$ = the measured ^{178}Hf and $\left(\frac{i^{176}\text{Hf}}{i^{178}\text{Hf}} \right) c$ = the initial or ‘common’
235 $^{176}\text{Hf}/^{178}\text{Hf}$ ratio. These corrections were applied using an assumed initial $^{176}\text{Hf}/^{178}\text{Hf}$ ratio of 0.192 ± 0.004 , which is equivalent
236 to a $^{176}\text{Hf}/^{177}\text{Hf}$ ratio of 0.2816 ± 0.006 . This value is based on the Hf evolution of the crust, with uncertainty that comfortably
237 covers likely natural variation. The uncertainty on the initial $^{176}\text{Hf}/^{178}\text{Hf}$ ratio used for the common Hf corrections has been
238 propagated to the final ages, in order to account for any inaccuracies introduced by value used. However, as most analyses
239 have $<1\%$ common Hf (Table 1), any inaccuracy related to the initial $^{176}\text{Hf}/^{178}\text{Hf}$ ratio is negligible compared to the total
240 uncertainty estimates given Hf isotopes do not vary significantly with time (Fisher and Vervoort, 2018; Vervoort, 2014). Such
241 corrections, however, should be used with caution for samples with higher common Hf, although the dataset presented in this
242 study is not sufficient to determine what an appropriate cutoff should be.

243

244 Subsequent to this correction, isotopic ratios were corrected using an external reference material bracketing approach
245 (commonly used in LA-ICP-MS geochronology), with primary and secondary reference materials interspaced with unknowns
246 through each analytical session. The data was normalized to NIST SRM 610 glass to correct for drift and matrix independent
247 fractionation. The Lu–Hf isotopic ratios published in Nebel et al. (2009) were used for the NIST610 SRM normalisation.
248 Following this, $^{176}\text{Hf}/^{176}\text{Lu}$, $^{176}\text{Lu}/^{177}\text{Hf}$, and $^{176}\text{Lu}/^{176}\text{Hf}$ ratios were corrected to MKED calcite. Although the age of MKED
249 calcite is currently not independently constrained, calcite is interpreted from textural evidence to have formed with the MKED
250 titanite reference material, and therefore the titanite TIMS U–Pb age was used (1517.32 ± 0.32 Ma; Spandler et al., 2016).
251 Further details are outlined in appendix A. This correction method is similar to that used by Roberts et al. (2017) for calcite
252 U–Pb, where the observed analytical offset between the measured and expected Lu–Hf ratio in the standard is applied (as a %
253 correction factor) to the ratios of the unknowns. This offset is inferred to be due to a combination of laser induced (matrix-
254 dependent) elemental fractionation and plasma loading. The uncorrected ages for MKED calcite as well as for ME 1 across
255 four analytical sessions are constant within uncertainty, indicating the age offset is a systematic analytical bias that is applicable

256 to the calcite samples of unknown age (Fig. A2). Weighted average ages were calculated using ISOPLOT (Vermeesch, 2018),
257 using the ^{176}Lu decay constant determined by Söderlund et al. (2004); $0.00001867 \pm 0.00000008 \text{ Myr}^{-1}$
258
259 Correct handling of uncertainties in geochronology is important in order to draw accurate conclusions about the resulting ages.
260 As per the recommendations for LA-ICP-MS U-Pb uncertainty propagation in Horstwood et al. (2016), uncertainties are
261 categorised as random, in which case they are propagated to individual analyses, or systematic, in which case they are
262 propagated to the final calculated age. As such, the uncertainties associated with the measurement of the primary standard
263 (NIST SRM 610) have been propagated to the uncertainties of individual analyses. The following systematic uncertainties
264 have been propagated to the final ages: Measurement uncertainty on the secondary standard (MKED C), uncertainty on the
265 titanite U-Pb age used as the reference age for MKED C, uncertainties associated with the ^{176}Lu decay constant, the reference
266 $^{176}\text{Hf}/^{177}\text{Hf}$ ratios for NIST SRM 610. Although for completeness it would be good to propagate uncertainty relating to potential
267 differences in $^{175}\text{Lu}/^{176}\text{Lu}$ and $^{177}\text{Hf}/^{178}\text{Hf}$ between NIST SRM 610 and samples (i.e. natural variation in these ratios), currently
268 there appears to be no data on this. These uncertainties are likely to be negligibly small relative to the overall uncertainty
269 estimates for the analyses. The uncertainty associated with the reference $^{176}\text{Lu}/^{177}\text{Hf}$, $^{176}\text{Lu}/^{176}\text{Hf}$, and $^{176}\text{Hf}/^{176}\text{Lu}$ ratios of NIST
270 SRM 610 are not propagated, as the correction factor associated with NIST610 SRM is cancelled during the correction to
271 MKED calcite (as the NIST610 SRM correction factor is applied equally to MKED calcite and the unknowns samples, and
272 thus becomes redundant). Uncertainty relating to long term reproducibility of the standards has not been propagated, as the
273 standard data for all sessions does not show scatter outside of what would be expected from a single population. More data,
274 however, is required to fully constrain this.

275 **4 Lu-Hf Results**

276 The analysed calcite generally contain < 1% common Hf, apart from sample P01, which contains up to 13% common Hf in
277 individual analyses (Table 1). Consequently, the common Hf corrections are small (or effectively non-existent), and the
278 resultant ages are not significantly affected by the assumed initial $^{176}\text{Hf}/^{177}\text{Hf}$ ratio. Corrected and uncorrected data are included
279 in Supplementary File 1. The inverse isochron and weighted mean single-spot Lu–Hf ages, reported below, are corrected
280 against MKED calcite for matrix-dependant fractionation and common-Hf corrected (where relevant) (Fig. 3). For analyses
281 with inclusions, the signals have been cropped to remove inclusions or, in the case of more significant signal disturbances,
282 excluded from age calculations. Inclusions were detected in the following samples: MKED calcite (6), LC1 (1), P01 (19) and
283 FF014 (6). Excluded data points are included in supplementary file 1. Due to the large number of inclusions, P01 was analysed
284 over two sessions. Data is presented as inverse isochrons (Li and Vermeesch, 2021), and as common Hf corrected weighted
285 average ages (Fig. 3).

286 5 Discussion

287 The Phalaborwa carbonatite sample produced a hf corrected weighted average Lu-Hf age of 2050 ± 30 Ma (Fig. 3), consistent
288 with previous baddeleyite U–Pb SIMS ages (~ 2060 Ma; Wu et al., 2011). Importantly, the consistency between the calcite Lu-
289 Hf age and existing constraints on carbonatite formation demonstrates that calcite Lu-Hf dating can produce primary age
290 information for early Paleoproterozoic calcite. This result also demonstrates that calcite Lu–Hf geochronology is a viable
291 technique to directly date carbonatite magmatism and associated mineralisation, even in the case of old calcite samples with
292 only ~ 0.5 ppm Lu.

293

294 The weighted average Lu–Hf ages for samples ME 1 and ME 2 are 1538 ± 9 Ma and 1504 ± 13 Ma, respectively (Fig. 3). The
295 ages of these samples are consistent with the paragenetic timing of alteration at Mt Elliott, providing evidence for calcite
296 precipitation during at least two temporally distinct alteration events. Sample ME 1 is from a coarse calcite-diopside-scapolite-
297 magnetite vein that does not contain sulfides (Fig. 1), the age is, therefore, consistent with formation prior to the major ~ 1510
298 Ma Cu-Au mineralisation event (Duncan et al., 2011; Wang and Williams, 2001). In addition, this age overlaps with a titanite
299 U–Pb age from the Mt Elliott deposit (1530 ± 11 Ma; Duncan et al., 2011), and is potentially related to regional Na-Ca
300 alteration between ca. 1555 and ca. 1521 Ma (Oliver et al., 2004). The 1504 ± 13 Ma age obtained from sample ME 2 that has
301 an ore stage paragenesis conforms with the $^{207}\text{Pb}/^{206}\text{Pb}$ age of cogenetic andradite (1507 ± 35 Ma; appendix B), and overlaps
302 with the ca. 1510 Ma main mineralisation event (Duncan et al., 2011; Wang and Williams, 2001). Additionally, data for ME1
303 was pooled from all four analytical sessions in order to test reproducibility. Similar to the standard (MKED1; appendix A),
304 ME1 does not show excess scatter between sessions (fig. 3).

305

306 Sample LC1, from the Lime Creek quarry, Eastern Fold Belt, Mt Isa Inlier produced an age of 1513 ± 26 Ma, consistent with
307 published titanite U–Pb ages (1521 ± 5 Ma 1527 ± 7 Ma) from the nearby Knobby Quarry (Oliver et al., 2004). Additionally,
308 this age is consistent with the intrusion of the ca. 1530-1500 Ma Williams-Naraku batholiths, which is interpreted to be the
309 source of the fluids from which the calcite precipitated (Oliver et al., 1993; Page and Sun, 1998). Our results for this sample
310 further demonstrate that calcite Lu–Hf geochronology is an effective technique for constraining the age of calcite
311 mineralisation.

312

313 Sample OL-MB from Otter Lake produced a Lu–Hf age of 892 ± 12 Ma (Fig. 3). This age is significantly younger than the
314 apatite solution Lu–Hf age of 1030 ± 6 Ma (Barfod et al. (2005) and the *in-situ* apatite Lu-Hf age of 1000 ± 11 Ma (Simpson
315 et al. 2021), but is similar to the apatite Pb–Pb age of 913 ± 7 Ma (Barfod et al. 2005) and the latest stage of extensional activity
316 on the nearby Bancroft Shear Zone (1045 - 893 Ma, Ar–Ar phlogopite; Cosca et al., 1995). Given the similarity between the
317 ca. 0.9 Ga ages, obtained by different methods, it seems likely that the calcite either grew or records Lu–Hf isotopic resetting
318 during the same event that induced resetting of the apatite Pb–Pb system. The slight difference between the calcite Lu–Hf age

319 (894 ± 12 Ma) and apatite Pb-Pb age (913 ± 7 Ma) may be due to analytical (i.e. mixing of age domains in the solution Pb-Pb
320 age) rather than geological reasons, particularly given an individual crystal of apatite from the Yates mine produced a U-Pb
321 age range of 920-850 Ma (Xiang et al., 2021). The age difference may also be due to underestimation of uncertainties. Large
322 (~3cm) apatite crystals such as the one analysed by Barfod et al. (2005) are expected to have Pb closure temperatures of up to
323 600 °C (Barfod et al., 2005; Krogstad and Walker, 1994), giving a possible upper limit to Lu–Hf closure in calcite. We note
324 that this is significantly higher than the closure temperature of Ar–Ar in phlogopite (ca. 400 °C), indicating that the Otter Lake
325 area potentially had a different thermal history and/or that isotopic resetting in the apatite and calcite was aided by late fluid
326 interactions, as hypothesised by Barfod et al., (2005). As such, further work is required to constrain the Lu–Hf closure
327 temperature in calcite.

328

329 The *in-situ* Lu–Hf age of 1810 ± 18 Ma for the cleavage-hosted calcite vein from the Flin Flon VMS deposit (FF14; Fig. 3),
330 as expected, is younger than the timing of initial mineralisation at the deposit (Koo and Mossman, 1975; Rayner, 2010; Stern
331 et al., 1995). Instead, the age is in excellent agreement with ca. 1820-1790 Ma regional peak greenschist to amphibolite grade
332 metamorphism (Schneider et al., 2007), suggesting the calcite precipitated during metamorphism related to deformation stage
333 ‘D₅’ or ‘D₆’, associated with the final collision between the Flin Flon-Glennie Complex and the Sask Craton (Lafrance et al.,
334 2016). This regional event locally reached maximum greenschist-facies metamorphism (Koo and Mossman, 1975), suggesting
335 the calcite grew under low-grade metamorphic conditions. Sample FF014, therefore, demonstrates that calcite Lu–Hf
336 geochronology has the potential to date low grade metamorphism, which has been difficult using traditional dating methods
337 (e.g. Henrichs et al., 2018).

338

339 In summary, we demonstrate that *in-situ* Lu–Hf geochronology can produce geologically meaningful ages for calcite from a
340 variety of mineralisation styles (e.g. IOCG, carbonatite, and skarn alteration) as well as greenschist-facies metamorphism. The
341 technique also has great potential to date a range of other geological settings and processes (e.g., chemical sedimentation,
342 carbonation reactions) provided calcite contains sufficient Lu for analysis.

343 **5.1 Limitations**

344 The success rate of the *in-situ* Lu–Hf dating approach in calcite is intrinsically related to; (1) the concentration of Lu, and; (2)
345 the ingrowth time for radiogenic Hf (Fig. 4). Generally, the method is more suitable for REE-rich calcite typically observed in
346 mineral deposits and carbonatites, and/or for Precambrian samples. In addition, the currently available mass-spectrometers
347 require large laser beam diameters (257 µm) for successful calcite Lu–Hf dating, limiting spatial resolution compared to most
348 laser-ablation dating techniques. We note that for high Lu samples, such as ME 1 (or samples that incorporate common Hf),
349 smaller spot sizes are feasible. Additionally, particularly in hydrothermal settings, calcite often forms large, mm to cm scale
350 crystals, reducing the need for small ablation volumes. While individual calcite crystals in other settings can sometimes be <
351 260 µm, the total amount of calcite is often large enough that aggregates of pure (or close to pure) calcite can be ablated.

352 Caution should be used with such analysis however, as this may affect laser induced fractionation, individual crystals may be
353 of different ages, and there may be micro inclusions of other minerals.

354 **5.2 Advantages of in-situ Lu-Hf dating of calcite compared to other geochronological methods**

355 The Previous dissolution-based Lu–Hf geochronology has produced scattered isochrons, indicative of isotopic disturbances
356 (Maas et al., 2020). While individual data points are significantly less precise than dissolution based methods, the ability to
357 gain spatially resolved data on a much smaller scale (>260 microns), as well as obtain a large number of analyses in a single
358 session can make data interpretation easier (Simpson et al., 2021a). Importantly, trace element data can be obtained
359 simultaneously to interrogate each data point for inclusions or age zonation. Furthermore, calcite Lu–Hf dating can overcome
360 two issues often encountered during U–Pb dating; (1) in contrast to Pb, calcite does not incorporate significant concentrations
361 of common Hf, and; (2) Lu is comparatively resistant to thermal diffusion in calcite (Cherniak, 1998), increasing the likelihood
362 of primary precipitation ages to be preserved. Although it should be acknowledged that fluid mobility and re-crystallisation of
363 the calcite may affect Lu-Hf ages, and are difficult to predict. This opens the possibility that time constraints can be obtained
364 for carbonates from the first three-quarters of Earth history that are generally difficult to date by other methods. Importantly,
365 calcite is commonly associated with ore formation, meaning *in-situ* Lu–Hf dating affords the possibility to directly constrain
366 the age of mineralising events and the temporal evolution of mineral deposit systems.

367

368 From our work, we suggest samples ME1 and OL-MB calcite could be developed as primary reference materials due to being
369 (1) common-Hf free, (2) homogenous in age across crystals up ~1cm cm in size, and (3) available in large quantities. We aim
370 to characterize such reference materials and make them available to the wider geochronology community.

371 **6 Conclusions and Future Directions**

372 Calcite is among the most common of rock-forming minerals, meaning that *in-situ* Lu–Hf geochronology of calcite has
373 enormous potential to constrain the age of formation and/or alteration of a range of igneous, sedimentary, metamorphic, and
374 hydrothermal rock systems, including rock-types that are considered very difficult to date (e.g., marbles). This technique has
375 particular application to mineral deposits as it allows for the ability to constrain the age of pre-ore, ore-stage and post-ore
376 events (e.g., Fig. 3). Furthermore, given the successful dating of old (~2 Ga) calcite with <1 ppm Lu (e.g., sample P01; Table
377 1), this technique has the potential to date old calcite from a variety of settings with relatively low HREE concentrations. *In-*
378 *situ* Lu–Hf dating of calcite can be regarded as a complimentary, and in some cases alternative, technique to carbonate U–Pb
379 dating, where Lu–Hf dating is well suited for older samples, or to obtain primary precipitation ages for systems affected by Pb
380 mobility. Coupling *in-situ* Lu–Hf dating with other isotopic systems (U–Th–Pb, C, O, Sr, Nd) may be particularly powerful for
381 constraining the origin, nature and redox conditions of the fluids or melts from which the calcite precipitated.

382

383 **Author contributions**

385 Author contributions: A Simpson (corresponding author): conceptualisation, method development, experimentation,
 386 manuscript drafting. S Glorie: conceptualisation, manuscript drafting, primary supervision. M Hand: conceptualisation,
 387 manuscript drafting, secondary supervision. C Spandler: Conceptualisation, sampling, manuscript drafting. S Gilbert: Method
 388 development, experimentation, manuscript drafting. B Cave. Experimentation, manuscript drafting.

389

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396

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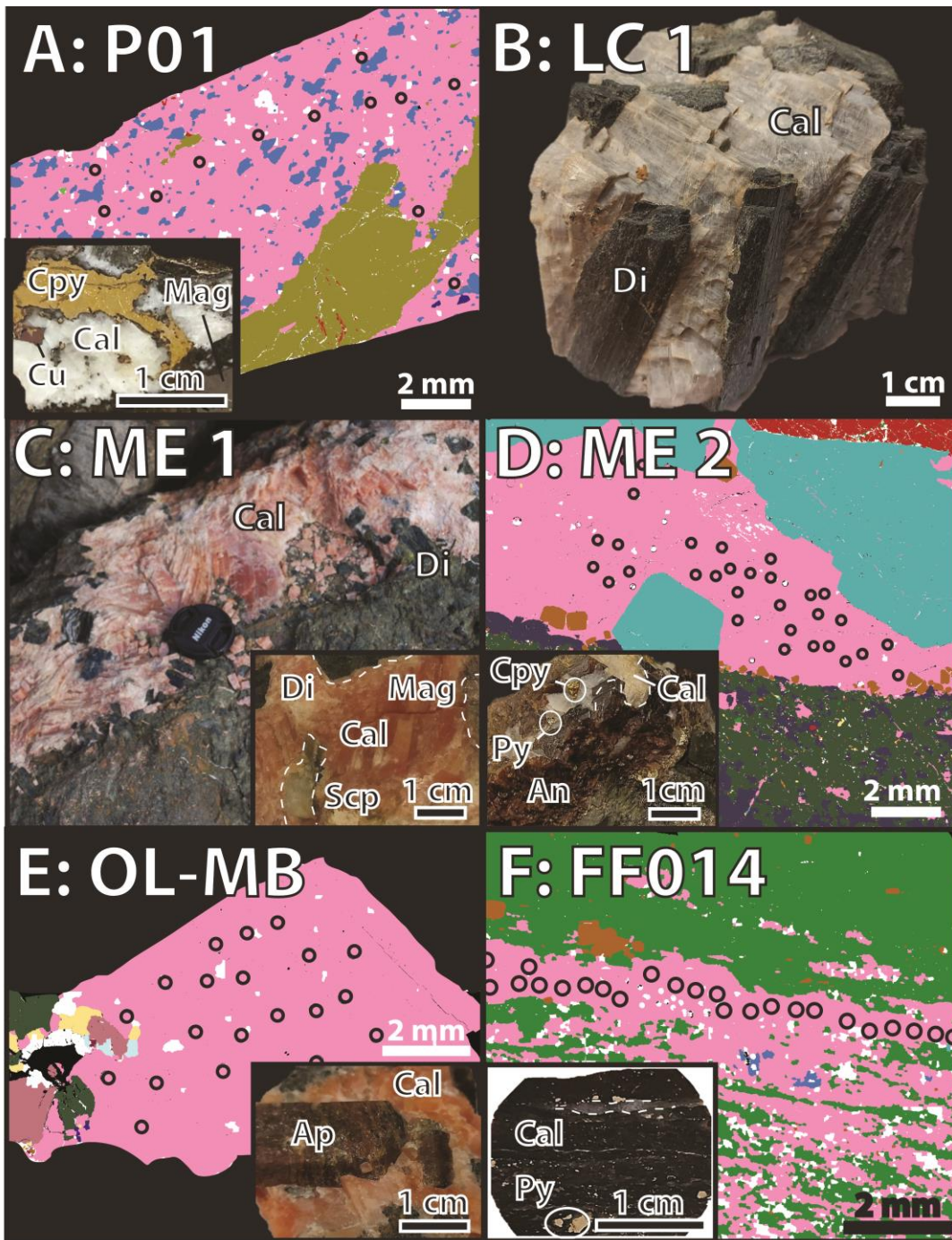
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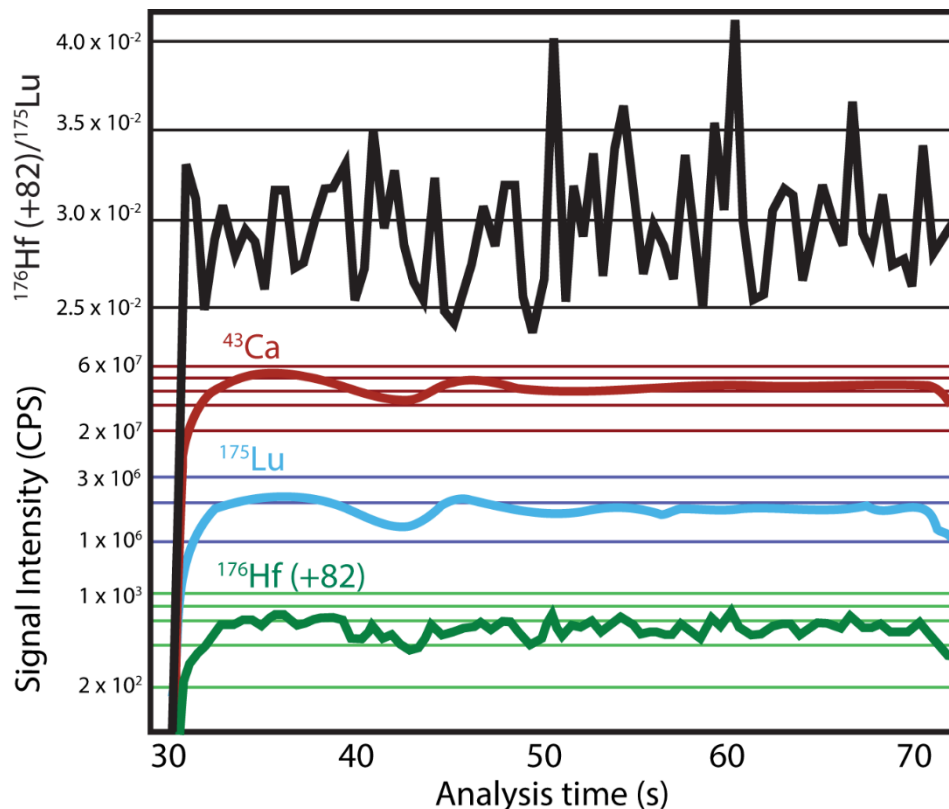
580 **Figures and Tables**



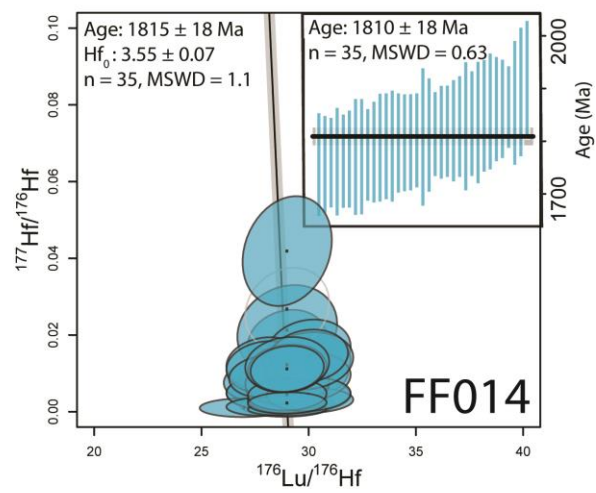
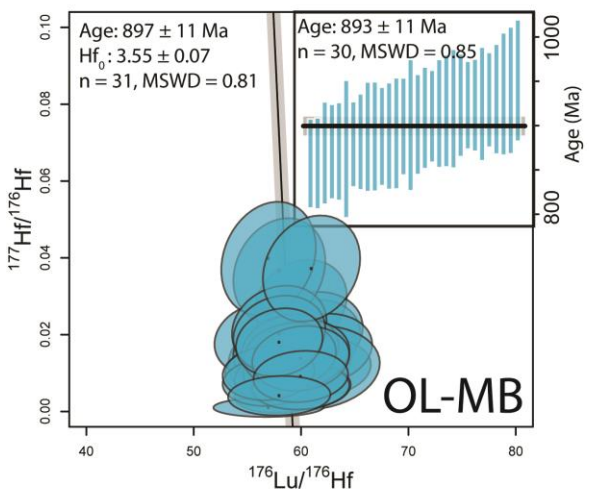
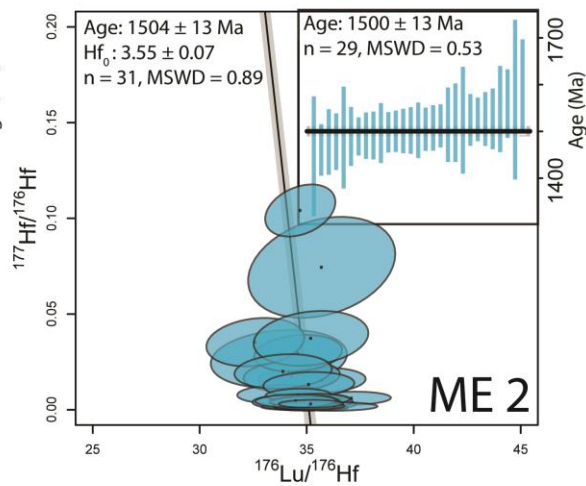
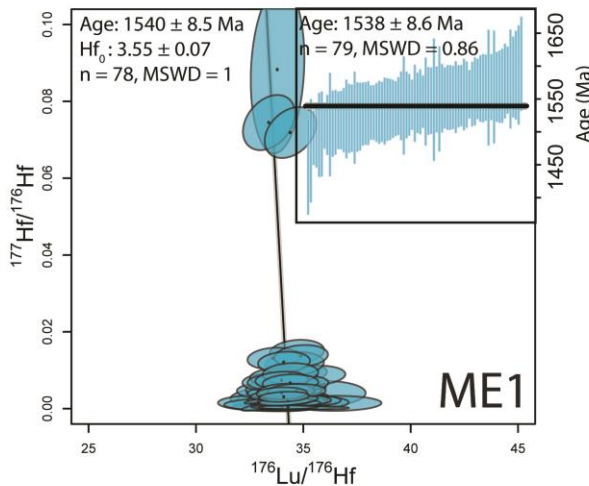
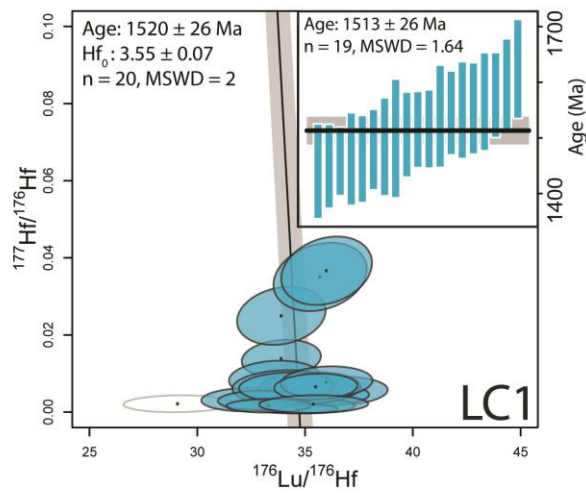
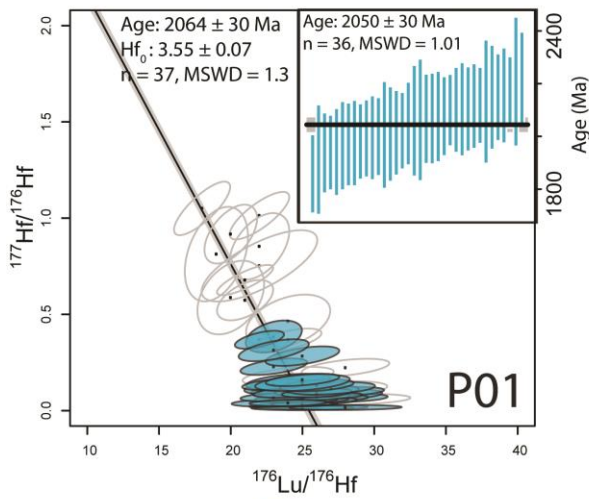
Mineral Key

	Calcite		Dolomite		Andradite		Chlorite		Quartz
	pyrite		Chalcopyrite		Augite		Magnetite		Aegirine/ Augite
	Biotite		Microcline		Albite				

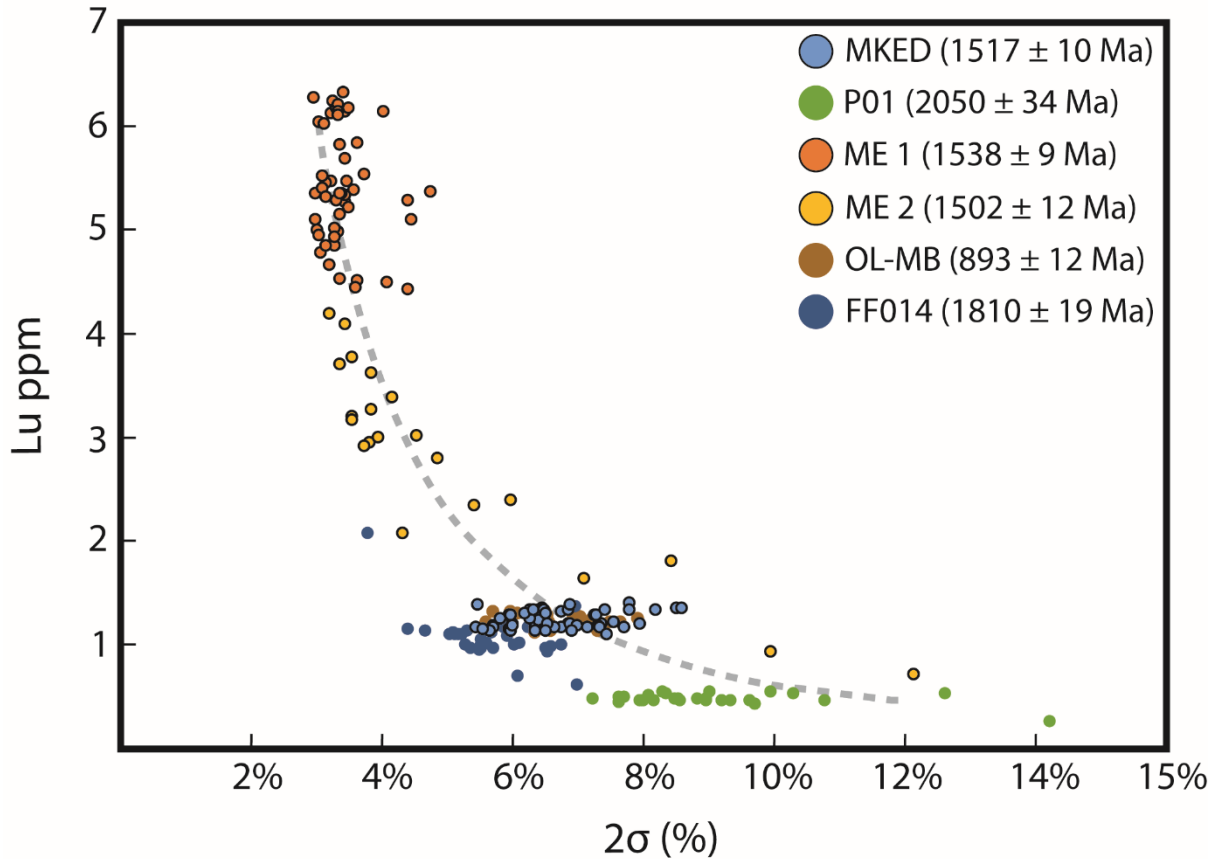
582 Figure 1: A combination of SEM mineral maps (A,C,D,E) and photos of analysed samples. (A) P01 (Phalaborwa Carbonatite) shows calcite in
 583 petrogenetic context to chalcopyrite. (B) Photo of LC 1 hand sample. (C) Photo of the ME 1 sample in outcrop, with inset showing the mineralogy of
 584 the analysed sample. (D) calcite from ME 2 (Mt Isa) in contact with hematite, pyrrhotite and andradite, with inset showing hand sample (D) OL-MB
 585 (Otter Lake), showing analysed calcite with associated minerals, with inset showing relationship between apatite (Ap) and calcite (Cal) in hand
 586 sample. (E) FF014 (Flin Flon deposit) shows calcite vein in chlorite matrix with disseminated pyrite, with inset showing analysed block (dark coloured
 587 matrix is composed of chlorite). Black circles represent laser spot locations. Mineral abbreviations: Cal: Calcite, Cpy: chalcopyrite, Py: pyrite, Mag:
 588 magnetite, Cu: cubanite, Di: diopside, Scp: scapolite, An: andradite. Larger size sample images are included in appendix C.
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 591 Figure 2: Time resolved signals for $^{175}\text{Lu}/^{(176)}\text{Hf} (+82)$, ^{43}Ca , ^{175}Lu , and $^{(176)}\text{Hf} (+82)$ demonstrating the effects of plasma loading on
 592 the signal intensities (i.e. dip in signal intensities at ~10-15s ablation), but not for the $^{176}\text{Hf}/^{175}\text{Lu}$ ratio, which remains constant down-
 593 hole. The time resolved intensity of each analyte has been offset in the graph for better comparison, therefore the Y axis scale is not
 594 continuous. Green horizontal lines show the scale for $^{176}\text{Hf} (+82)$, blue horizontal lines show the scale for ^{175}Lu , red horizontal lines
 595 show the scale for ^{43}Ca , and black horizontal lines show the scale for the $^{176}\text{Hf} (+82)/^{175}\text{Lu}$ ratio. Presented data is from an analysis
 596 of MKED calcite.



598 **Figure 3: Anchored inverse isochron and weighted average ‘single spot’ ages for analysed samples, corrected for matrix-induced**
 599 **fractionation against MKED1 calcite. Isochrons have been anchored to an initial $^{177}\text{Hf}/^{176}\text{Hf}$ ratio of 3.55 ± 0.07 . Ellipses represent**
 600 **data points and 2σ uncertainty. Uncoloured ellipses represent data points excluded from the regression calculation. Weighted**
 601 **average ages are corrected for common-Hf where relevant (see Table 1 and text). Blue bars represent 2σ uncertainties. Black lines**
 602 **represent weighted average ages, with grey boxes representing the 95% confidence interval uncertainty.**



603 **Figure 4: Lu ppm vs 2σ uncertainty for each calcite analysis. The grey curve shows a function fitted to the data from samples with**
 604 **ages between 1500 and 1540 Ma (samples ME1, ME2, and MKED, with symbols outlined in black). Only data points with similar**
 605 **ages were used to construct this guiding curve as the obtained precision is age-dependant. The Lu-Hf ages for older samples (e.g.**
 606 **P01 and FF014) are more precise relative to younger samples for a given Lu concentration (assuming no common Hf). Note: MKED**
 607 **is the calcite Lu-Hf standard used to correct the analysed samples. All data for MKED is included in supplementary file 1.**
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619 TABLE 1: AGES AND LU AND HF CONCENTRATION INFORMATION FOR THE ANALYSED SAMPLES.

Sample	Age	95%	n	Min.	Av %	Max.	Min	Av	Max	Min	av Hf*	Max
	(Ma)	CI		% Hf	Hf	% Hf	Lu	Lu	Lu	Hf*	ppb	Hf
				corr	corr	corr	ppm	ppb	ppb	ppb		ppb
P01	2054	1.5%	36	0%	2.65%	13%	270	505	557	0.03	0.270	1.00
LC 1	1513	1.7%	19	0.16%	0.46%	1.25%	1900	1600	3600	0.011	0.090	0.360
ME 1	1540	0.6%	79	0%	0.28%	2.4%	4300	5325	6300	0.002	0.150	1.80
ME 2	1500	0.9%	29	0.16%	0.58%	3.04%	700	3159	5500	0.002	0.110	0.340
OL-MB	892	1.2%	30	0.16%	0.58%	1.25%	1100	1200	1300	0.003	0.050	0.120
FF014	1807	1.0%	35	0.16%	0.39%	1.3%	616	1020	1170	0.004	0.052	0.210

620 *Note: 95% CI refers to the 95% confidence interval uncertainty on the calculated age. n refers to the number of analyses used for the*621 *age calculation. % Hf corr refers to the average % decrease in age due to the common Hf correction. Hf* concentrations have been*622 *calculated from ¹⁷⁸Hf and assume no radiogenic ingrowth of ¹⁷⁶Hf and thus represent the ‘common’ Hf concentration for each sample.*

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624 **Appendix A: MKED calcite sample description**625 A sample of orange-pink calcite associated with the MKED1 titanite U-Pb standard (1517.32 ± 0.32 Ma, U-Pb TIMS; Spandler

626 et al., 2016) was analysed as a matrix matched secondary standard in order to correct unknown samples for matrix related

627 analytical offsets, such as laser induced elemental fractionation and plasma loading effects. The calcite was sampled from the

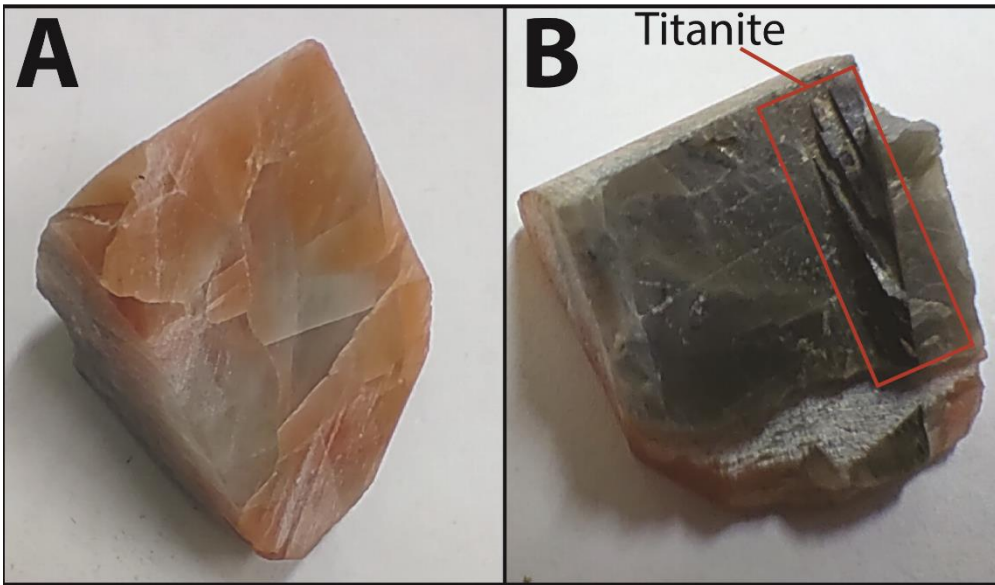
628 same drill core from which the titanite standard was taken (full details can be found in Spandler et al., 2016). The sample

629 consists of massive calcite surrounding large (~8cm) euhedral titanite crystals. The titanite is interpreted to have grown in the

630 same fluid as the calcite, but just prior to calcite crystallization. The average age across all 4 analytic sessions is 1560 ± 10 Ma

631 (fig. A2), suggesting that matrix fractionation during laser ablation produces ages that are systematically approximately 3%

632 too old

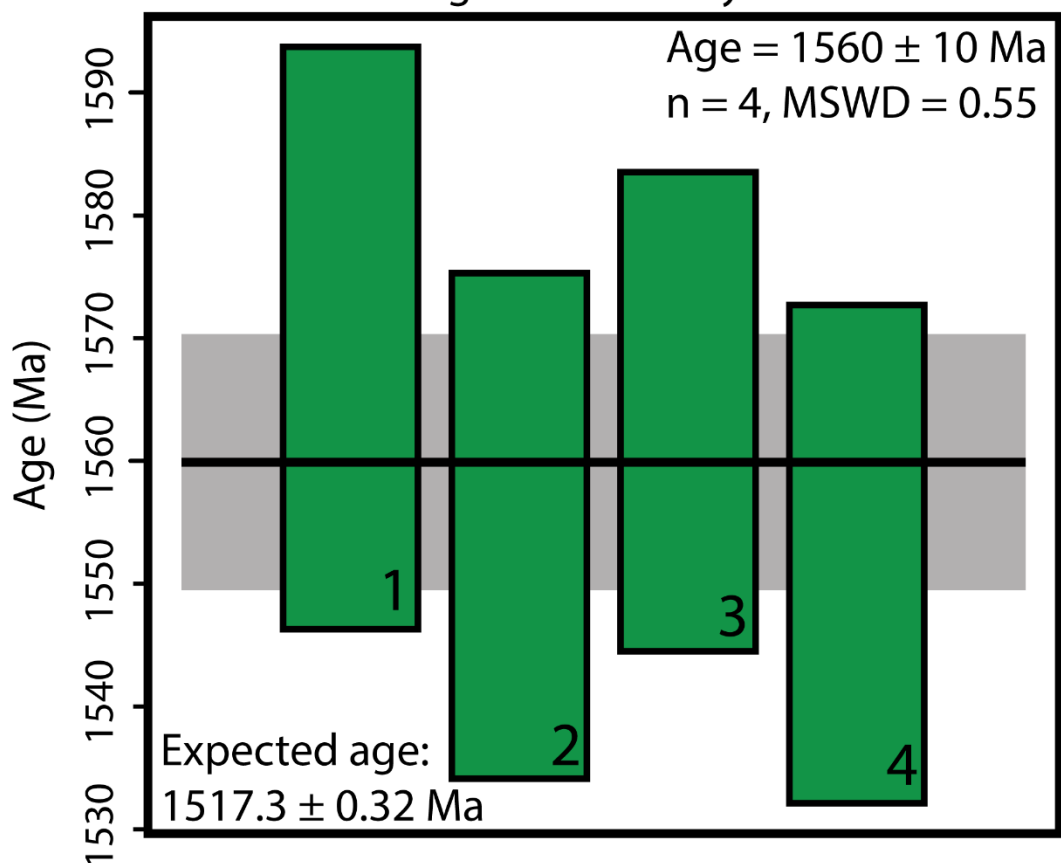


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634 **Figure A1: images of MKED calcite. A shows calcite chip from where the analysed sample was taken. B shows underside of the same**
635 **chip, where a large titanite crystal has been removed. Red box shows remnant fragments of titanite.**

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MKED Calcite ages for all analytical sessions



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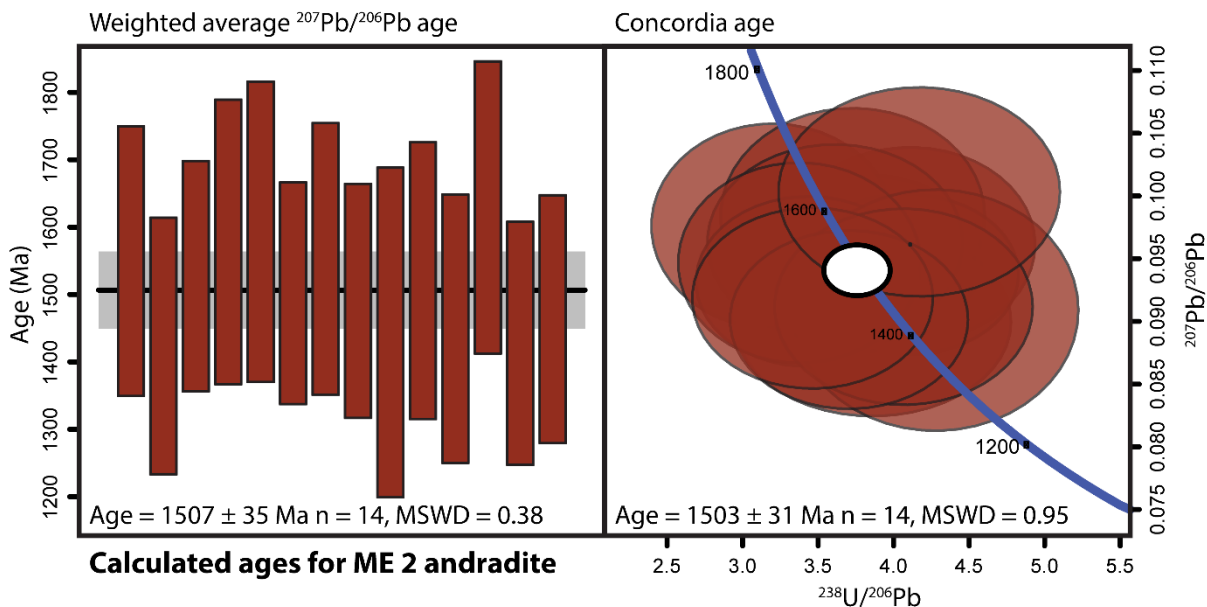
638 **Figure A2: Demonstration of the systematic analytical offset observed for calcite Lu-Hf ages. Green rectangles are 95% confidence**
639 **intervals around weighted mean ages for each session, with session number in the bottom right corner of each rectangle. Horizontal**
640 **black line shows the weighted average age of all analytical sessions, with grey rectangle showing 95% confidence interval uncertainty.**
641 **The combined weighted average age for all analytical sessions is shown in the top right corner. The expected age is from Spandler**
642 **et al. (2016). Weighted mean ages were calculated using ISOPLOTR (Vermeesch, 2018).**

643 **Appendix B: Mt Elliott Andradite U-Pb data:**

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645 Cogenetic andradite was analysed from the Mt Elliott 2 calcite sample (fig. 1). The sample was analysed using the same laser
646 system as used for Lu-Hf analysis, but coupled with an Agilent 7900 quadrupole mass spectrometer. As the University of
647 Adelaide does not currently possess an andradite U-Pb standard, U-Pb and Pb-Pb ratios were corrected to NIST610 SRM,
648 using ratios from (REF). A large aspect ratio ablation spot (120 microns in diameter, drilling approximately 30 microns deep)

649 was used to minimise the effects of downhole fractionation (REF – Sylvester?) however, it is possible that calculate U-Pb ages
 650 are inaccurate due to the lack of matrix matched primary standard. As the data appears to be concordant, however, a weighted
 651 average age can be calculated from the $^{207}\text{Pb}/^{206}\text{Pb}$ ratios (fig. B1), which should not be significantly affected by laser induced
 652 matrix fractionation (REF). As such, the calculated age is considered accurate within uncertainty.



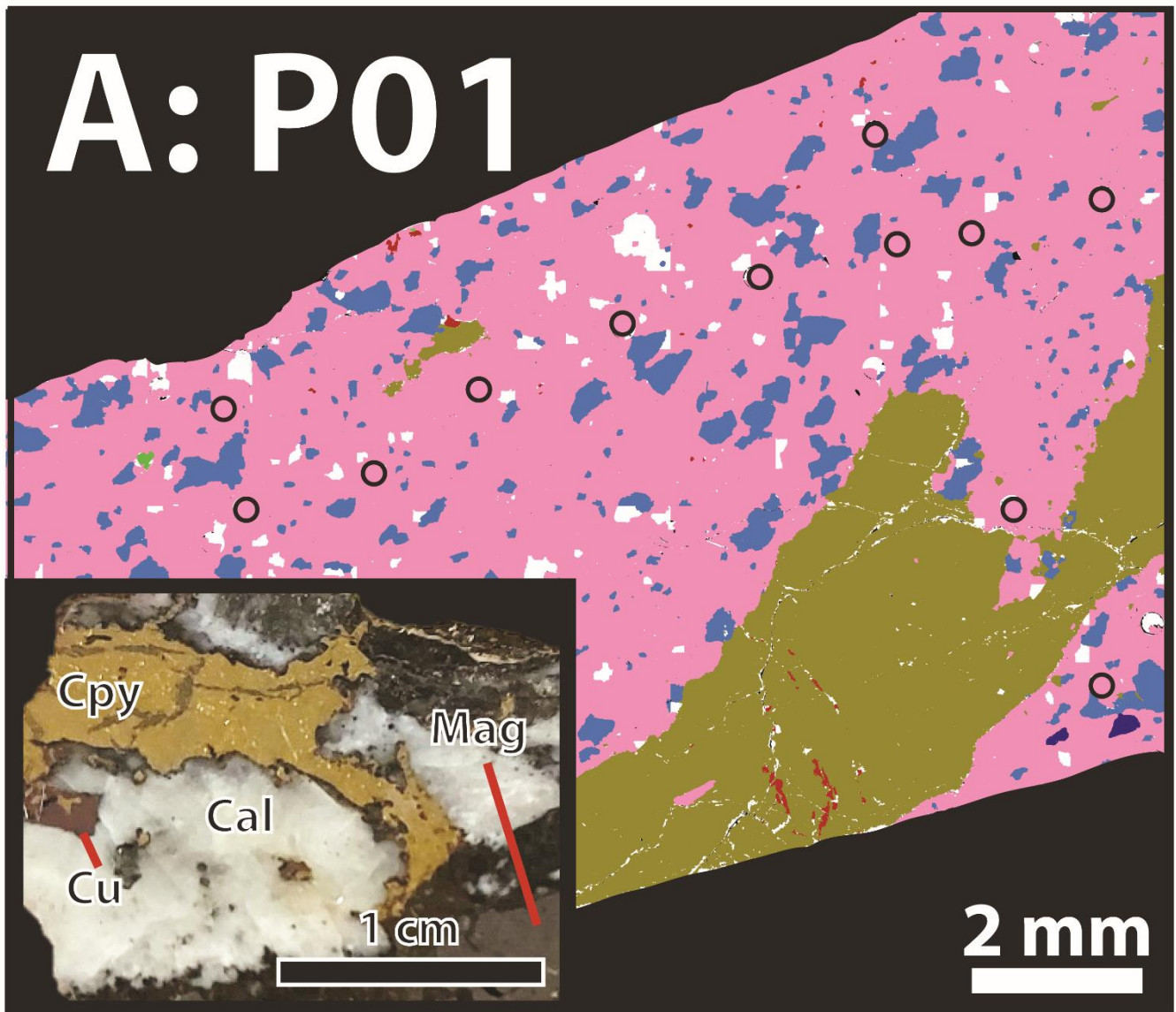
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654 **Figure B1: U-Pb age of ME 2 andradite.** Left panel shows the weighted average $^{207}\text{Pb}/^{206}\text{Pb}$ age and the right panel shows the
 655 concordia age on a Tera-Wasserburg Cconcordia plot. For the weighted average, vertical rectangles are 2σ uncertainties around
 656 calculated single spot ages, with the black bar showing calculated weighted mean age, and the grey rectangle showing associated
 657 95% confidence interval uncertainty. For the concordia plot, each ellipse shows the 2σ uncertainty around each analysis, with white
 658 ellipse representing 95% confidence interval uncertainty around the calculated concordia age. Weighted mean age and concordia
 659 age were calculated using ISOPLOTR (Vermeesch, P., 2018).

660 Appendix C: Large sample images

661 The following are larger versions of the sample images from figure 1. Mineral abbreviations are: Cal: Calcite, Cpy:
 662 chalcopyrite, Py: pyrite, Mag: magnetite, Cu: cubanite, Di: diopside, Scp: scapolite, An: andradite

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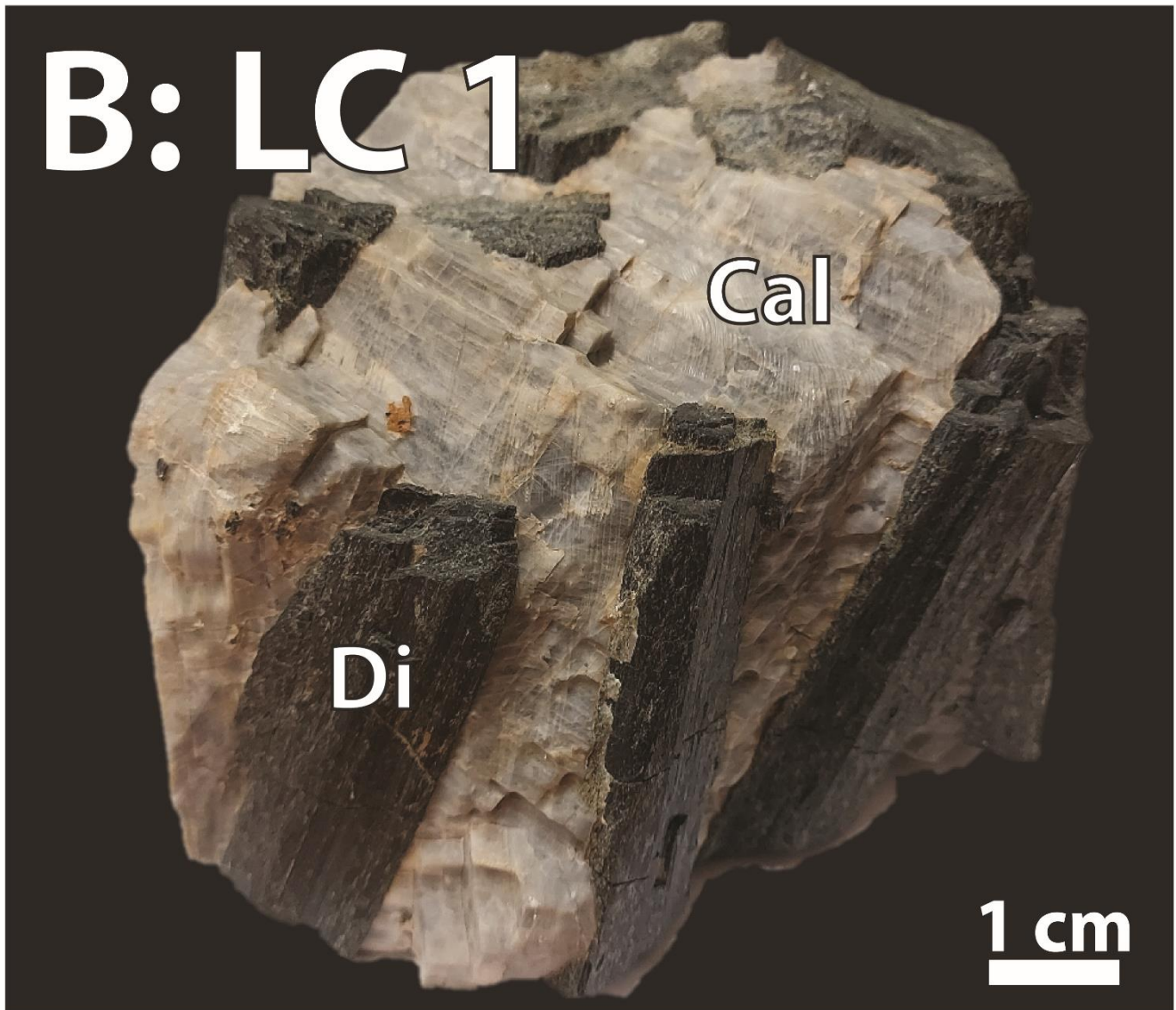
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665 Figure C1: Large image of sample P01 from Phalaborwa carbonatite, South Africa. Large image shows SEM mineral map. Inset

666 shows hand sample photo. Cpy: Chalcopyrite, Cu: Cubanite, Cal: calcite, Mag: Magnetite.

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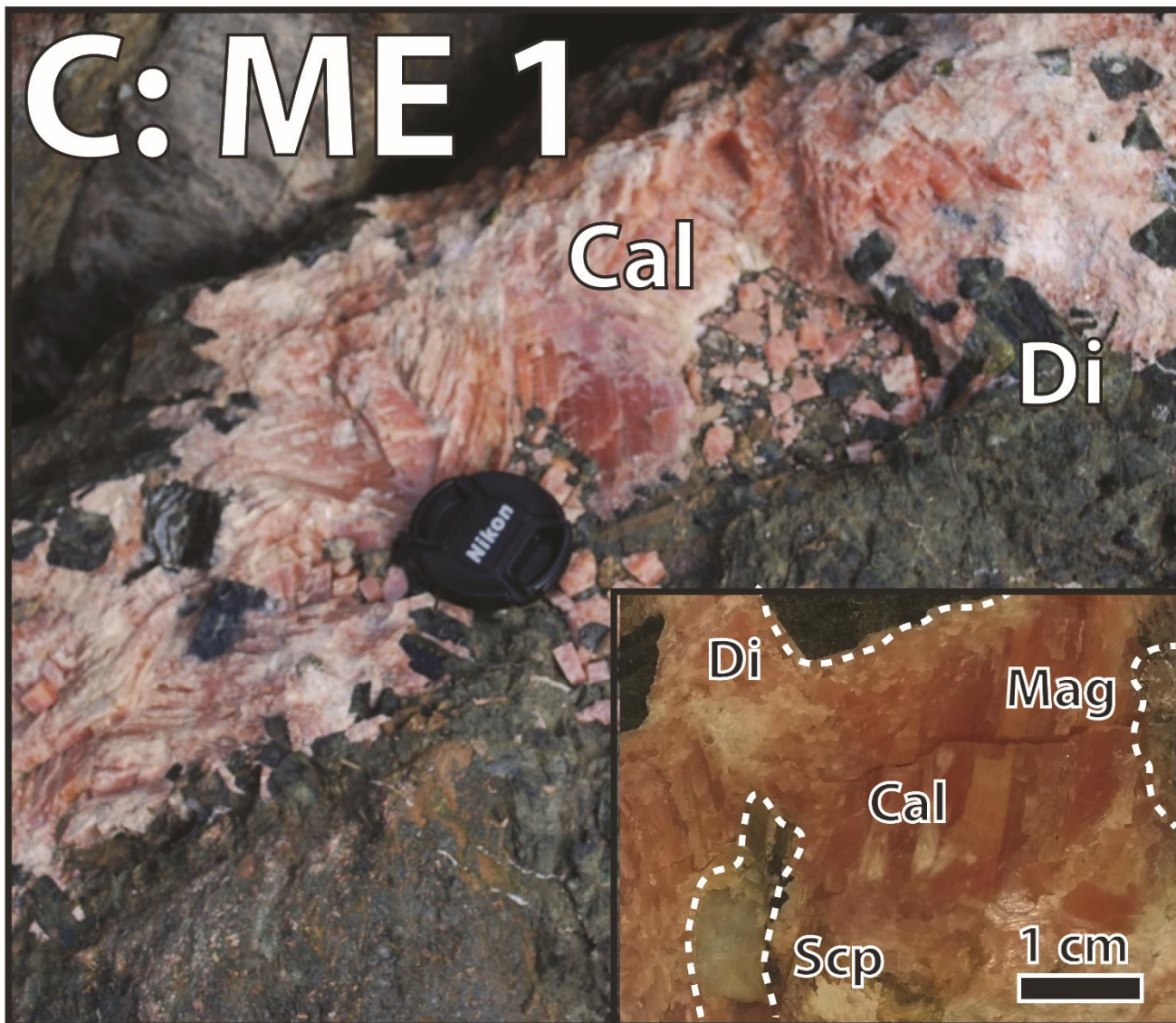


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670 **Figure C2: Large image of sample LC1 from Lime Creek, Mt Isa region, Australia. Image shows hand sample. Di: diopside, Cal:**

671 **calcite.**

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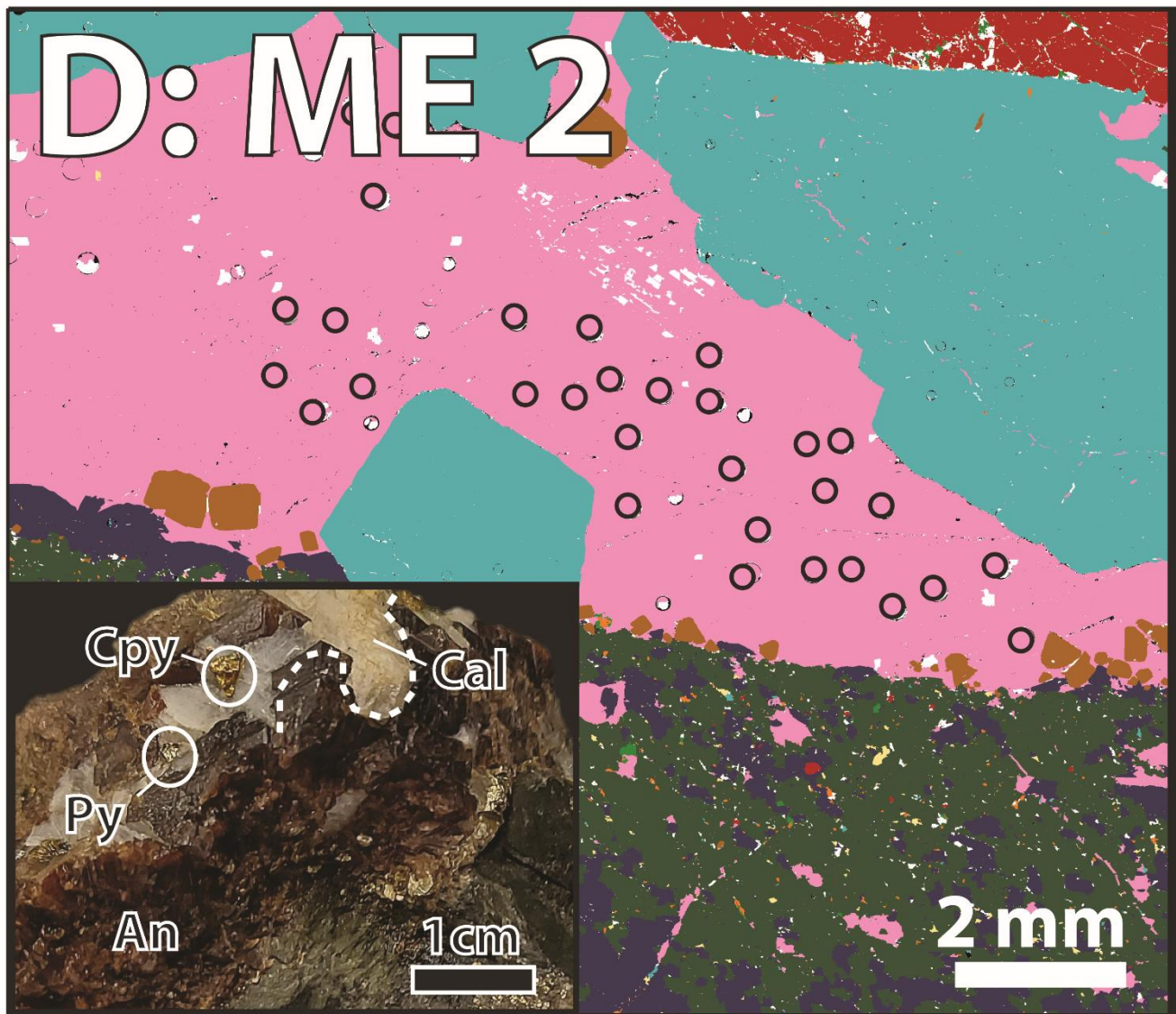


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674 **Figure C3: Large image of sample ME1 from Mt Elliott, Mt Isa region, Australia. Large images shows sample location and inset**

675 **shows hand sample. Di: diopside, Scp: scapolite, Cal: calcite, Mag: magnetite**

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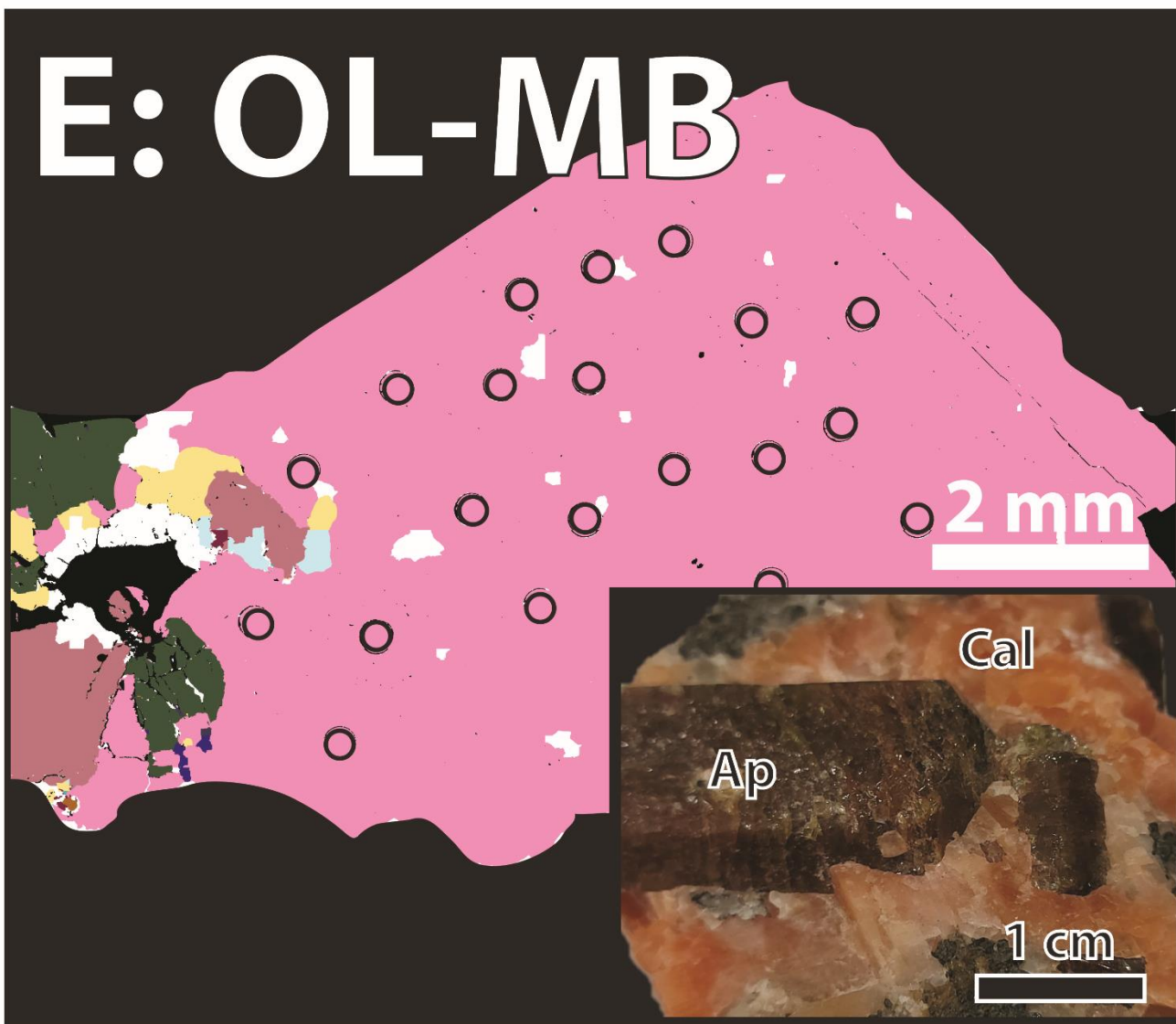
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678 **Figure C4: Large image of sample ME2 from Mt Elliott, Mt Isa region, Australia. Large images shows SEM mineral map with black**

679 **circles showing laser spot locations. Inset shows hand sample photo. Cpy: Chalcopyrite, Py: pyrite, Cal: calcite, An: andradite.**

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E: OL-MB

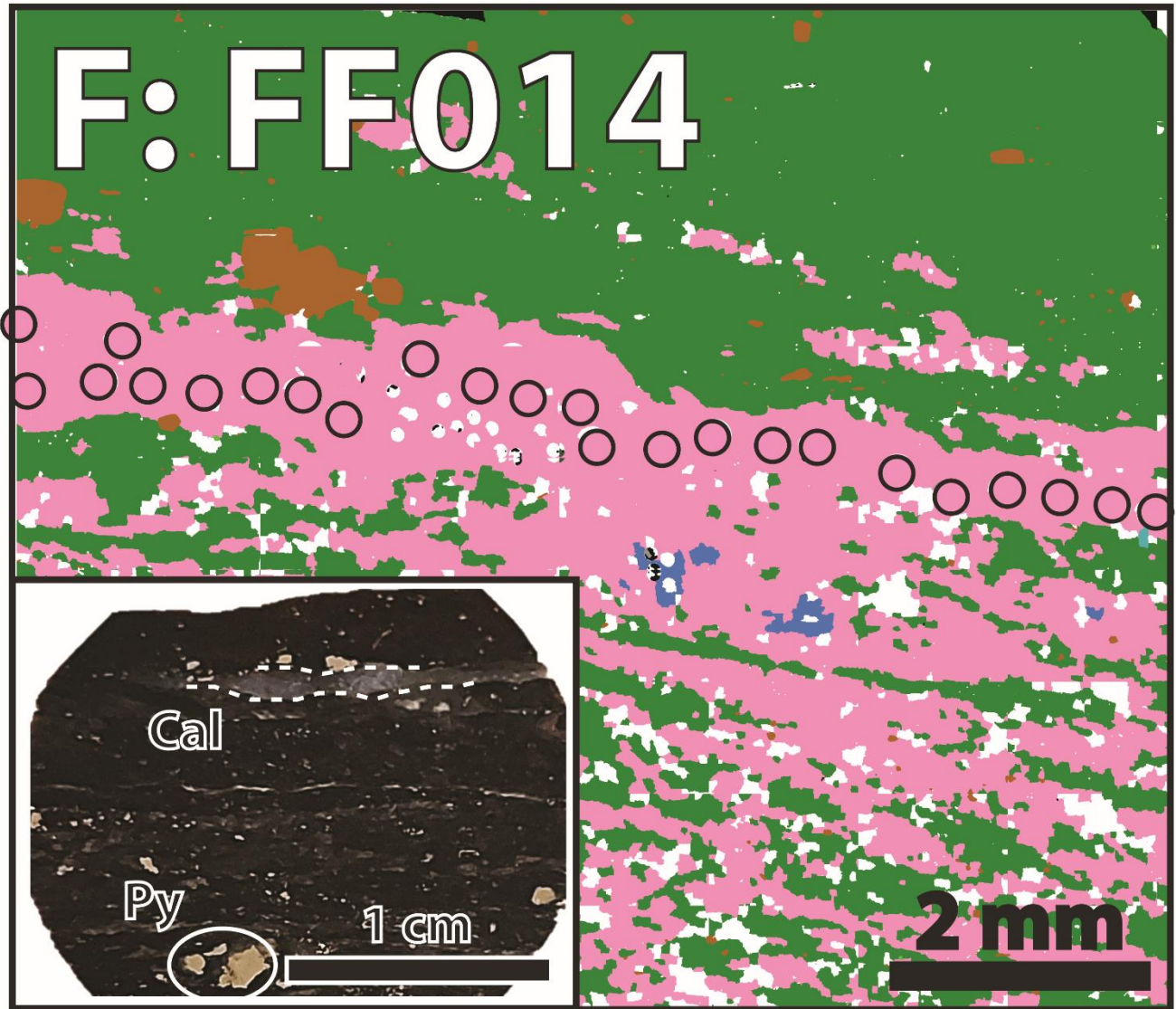


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682 **Figure C5: Large image of sample OL-MB from the Yates mine, Canada. Large images shows SEM mineral map with black circles**

683 **showing laser spot locations. Inset shows hand sample photo. Ap: apatite, Cal: calcite.**

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686 **Figure C6: Large image of sample FF014 from Flin Flon, Manitoba and Saskatchewan, Canada. Large images shows SEM mineral**

687 **map with black circles showing laser spot locations. Inset shows hand sample photo. Py: pyrite, Cal: calcite.**

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695 **APPENDIX TABLE 1: ANALYSIS AND LA-ICP-MS/MS TUNING PARAMETERS***plasma parameters*

<i>RF power</i>	1350 W
<i>Sample Depth</i>	4 mm
<i>Ar carrier gas</i>	0.94 L/min
<i>He carrier gas</i>	0.38 L/min
<i>N₂ addition</i>	3.5 mL/min
<i>Lens Parameters</i>	
<i>Extract 1</i>	-1.5 V
<i>Extract 2</i>	-140 V
<i>Omega Bias</i>	-70 V
<i>Omega Lens</i>	8.0 V
<i>Q1 entrance</i>	-45 V
<i>Q1 exit</i>	1.0 V
<i>Cell focus</i>	1.0 V
<i>Cell Entrance</i>	-120 V
<i>Cell Exit</i>	-100 V
<i>Deflect</i>	10.0 V
<i>Plate Bias</i>	-60 V
<i>Q1 parameters</i>	

<i>Q1 bias</i>	-1.0 V
<i>Q1 Prefilter Bias</i>	-10.0 V
<i>Q1 Postfilter Bias</i>	-10.0 V
<i>Cell Parameters</i>	
<i>He flow</i>	1.0 mL/min
<i>10% HN₃ + 90% He gas flow</i>	3 mL/min
<i>Octopole bias</i>	-2.0 V
<i>Axial Acceleration</i>	2.0 V
<i>Octopole RF</i>	180 V
<i>Energy Discrimination</i>	-13.0 V
<i>Q2 parameters</i>	
<i>Q2 bias</i>	-15 V
<i>Wait time offset</i>	5ms
<i>Analysis Parameters</i>	
<i>Laser wavelength</i>	193 nm
<i>Laser fluence</i>	10 J/cm ²
<i>Laser spot diameter</i>	257 μm (43 μm; NIST610 glass)
<i>Laser repetition rate</i>	10 Hz
<i>washout</i>	30 s (post cleaning pulse) + 20 s (post analysis)
<i>Background</i>	30 s
<i>Analysis time</i>	40 s

<i>Isotopes</i>	<i>measured/dwell</i>	^{27}Al (2), ^{43}Ca (2), ^{47}Ti (2), ^{89}Y (2), ^{90}Zr (2), ^{140}Ce (2), ^{172}Yb
<i>times (ms)</i>		(10), ^{175}Lu (10), $^{175+82}\text{Lu}$ (100), $^{176+82}\text{Hf}$ (150), $^{178+82}\text{Hf}$ (150)

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