

Short communication: Synchrotron-based elemental mapping of single grains to investigate variable infrared-radiofluorescence emissions for luminescence dating

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Abstract. During ionising irradiation, potassium (K)-rich feldspar grains emit infrared (IR) light, which is used for infrared-radiofluorescence (IR-RF) dating. The late-saturating IR-RF emission centred at ~880 nm represents a promising tool to date Quaternary sediments. ~~However, in~~ the present work, we report the presence of individual grains ~~of~~ in the K-feldspar ~~density fraction~~ displaying an aberrant IR-RF signal shape, whose combined intensity contaminates the sum signal of an aliquot composed of dozens of grains. ~~Our~~ experiments were carried out at the National Synchrotron Light Source (NSLS-II) ~~on~~ at the submicron resolution X-ray spectroscopy (SRX) beamline. We analysed coarse (> 90 µm) K-feldspar ~~bearing~~ grains of five samples of different ages, ~~nature~~ and origin in order to ~~characterise~~ characterize the composition of grains yielding the desired or contaminated IR-RF emission. Using micro-X-ray-fluorescence (~~µ~~XRF~~µ~~-XRF), we successfully acquired element distribution maps of fifteen elements (<1 µm resolution) of ~~the surface~~ sections of full grains previously used for luminescence dating. In keeping with current theories of IR-RF signal production, we observed a ~~correlation~~ trend between the relative proportions of Pb and Fe and the shape of the luminescence signal: most grains with the desired IR-RF signal shape had high Pb and low Fe contents. Interestingly, these grains were also defined by high Ba and low Ca contents. ~~Additionally, this~~ Our study ~~also~~ represents a proof-of-concept for mapping the oxidation states of Fe ~~ions~~ using micro-X-ray absorption near-edge structure spectroscopy (~~µ~~XANES~~µ~~-XANES) on individual grains. The high spatial resolution enabled by synchrotron ~~X-ray~~

33 spectroscopy makes it a powerful tool for future experiments to elucidate long-standing issues concerning the nature and type
34 of defect(s) associated with the main dosimetric trap in feldspar.

35 **1 Introduction**

36 Geochronologic data provide essential information for understanding the rates of Earth's surface processes, environmental
37 changes, and the evolution of life. Advances in dating techniques have fundamentally changed our capacity to piece together
38 our evolutionary past over millions of years, with luminescence dating proving to be a powerful tool in this field as it applies
39 to various types of sediments and contexts. The technique determines an age estimate for when mineral grains were last
40 exposed to daylight or heat. Luminescence dating methods rely on the capacity property of certain minerals to record the amount
41 of radiation to which they have been exposed during burial and release energy when exposed to sunlight or high temperature
42 (e.g., Aitken, 1985, 1998; Bateman, 2019). In the laboratory, the total amount of energy per unit mass stored in the mineral is
43 measured ~~as a~~ (dose ~~(\times)~~ Gy). The energy absorption rate per unit mass (dose rate, Gy a⁻¹) is derived from knowledge of the
44 natural radioactivity surrounding the sampled sediment. The quotient of these two values (dose/dose rate) gives the burial
45 timeage.

46 ~~Generally, Of the preferred mineral is two minerals routinely used for luminescence dating of sediments, quartz~~
47 ~~because of its high abundance and resistance to weathering. However potassium (K-) rich feldspar, the early saturation of the~~
48 ~~optically stimulated luminescence (OSL, Huntley et al., 1985) signal within quartz at ~200 Gy (Wintle and Adamiec, 2017)~~
49 ~~generally limits its application to the last 200, latter allows for the dating of older deposits of up to ~600 000 years (considering~~
50 ~~a low dose rate of 1 Gy ka⁻¹). By contrast, potassium (K-) rich feldspar minerals typically display significantly higher dose~~
51 ~~saturation levels) using infrared stimulated luminescence (IRSL, Hütt et al., 1988). The datable upper age limit is given by the~~
52 ~~IRSL signal saturation after exposure to radiation doses around 600 Gy (see summary in Sec. 8.1 in Murari et al., 2021a) using~~
53 ~~infrared stimulated luminescence (IRSL, Hütt et al., 1988), allowing the dating of older deposits up to 600,000 years~~
54 ~~(considering a low dose rate of 1 Gy ka⁻¹).~~

55). Over the past decades, different methods have been proposed to extend this upper age limit with varying degrees
56 of success. The infrared-radiofluorescence (IR-RF) signal of K-feldspar is a promising candidate for such an extension. ~~The~~
57 ~~IR RF emission at 880 nm (Kumar et al., 2018; Riedesel et al., 2021; Sontag-González et al., 2022) The RF signal~~ arises from
58 prompt radiative recombination of charge within crystalline materials during continuous exposure to ionizing radiation. The
59 IR-RF ~~signal~~ emission at 880 nm (e.g., Kumar et al., 2018; Riedesel et al., 2021; Sontag-González et al., 2022) decreases in
60 intensity with ~~the~~ dose accumulation as the electron traps fill until saturation (Trautmann et al., 1999a). This saturation level
61 constrains the time range over which IR-RF dating is applicable. ~~The IR RF dose is determined in three measurement steps:~~
62 ~~(i) an additive irradiation is given while the natural IR RF signal is recorded, (ii) the sample is optically bleached to empty the~~
63 ~~traps and after a pause, (iii) a second irradiation is performed during which the regenerated IR RF signal is measured. Finally,~~

~~the natural dose absorbed during burial is calculated from the horizontal distance over which the natural signal needs to be shifted to match the regenerative signal curve (see Murari et al., 2021a for a review).~~

Murari et al. (2018) demonstrated that an accurate dose recovery of a known dose of 3600 Gy is possible (a dose recovery test is a laboratory performance check of the measurement protocol, and successful dose recovery is a prerequisite for any protocol). If we assume typical environmental dose rates of between 3 Gy ka⁻¹ and 1 Gy ka⁻¹, then IR-RF dating could produce age estimates ranging from 1.2 Ma to 3.6 Ma, which is around four times greater than the upper dating limit of conventional luminescence dating methods. However, more recent studies (Murari et al., 2021b; Kreutzer et al., 2022) indicated a dose saturation ~~cap~~ at around 1500 Gy, reducing the previously assumed temporal limit of IR-RF dating. Hence, the uncertainty surrounding its upper age limit remains and further studies on known-age samples are required to assess whether the sample/grain geochemistry influences the age limit. There is undoubtedly a gap in our current understanding of the luminescence production processes in K-~~rich~~ feldspar, and a revised conceptual model might be needed.

~~To assess whether the grain geochemistry influences the IR-RF signal and perhaps~~The vast majority of IR-RF studies have been performed on multi-grain aliquots, so the possible effects of variability of the IR-RF signal from different grains has not received much attention in the literature so far, as detailed in section 2. Here, we investigate the IR-RF signal of five samples from different locations at single-grain resolution and discuss what effect the observed variability could have on multi-grain aliquots. To assess whether the grain geochemistry influences the IR-RF signal and potentially the age limit of IR-RF, we examined individual K-feldspar grains at the submicron resolution X-ray spectroscopy (SRX) beamline at the National Synchrotron Light Source II (NSLS-II) at Brookhaven National Laboratory. Measurements at such a high-resolution ~~will~~may lead to a better understanding of the luminescence kinetics in feldspars. ~~Here we~~We report on the feasibility and practicality of using μ -X-ray fluorescence (μ -XRF) and μ -X-ray absorption near-edge fine structure (μ -XANES) techniques in investigating the luminescence signal origin and ~~kinetic~~kinetics in K-feldspar.

2 Method and rationale

Identification of the defect type linked to the IR-RF signal and its concentration would enable us to better characterize the light emission (signal sensitivity) in different types of feldspar, while identification of the origin of possible contamination in the IR-RF signal could help us to gain a better understanding of the apparent early saturation or quenching of the IR-RF signal. μ -XRF and μ -XANES produce high-resolution maps of elements and their oxidation states and are well suited for the purposes of our study. μ -XRF elemental analyses are based on the characteristic fluorescence of atoms when stimulated with X-rays with a higher energy than their ionisation energy. In the case of μ -XANES, initial measurements of standards are run by varying the incident beam energy to determine the specific energy equal to the absorption edge (binding energy of inner shell electrons) of the element or ion of interest. This is apparent by an abrupt rise in the resulting fluorescence, which is different between oxidation states as they require different minimum stimulation energies before ionisation and subsequent fluorescence. μ -XRF maps using the obtained absorption edge energies allow for maps of the different oxidation states of the same element.

96 The use of synchrotron μ -XRF allows us to improve the spatial resolution compared with a standard lab-bench μ -
97 XRF setup (e.g., Buylaert et al., 2018) by reducing the beam spot size from $\sim 25 \mu\text{m}$ to $1 \mu\text{m}$ or $0.5 \mu\text{m}$. Though both the grain
98 geochemistry and crystallography should be investigated to characterize the defect type and its environment, in the present
99 study, we focus only on geochemistry.

100 The defect(s) responsible for the IR-RF emissions are still subject to debate. It has been suggested that IR-RF occurs
101 as a result of the change in the oxidation state of the participating lead (Pb) defect via the transition: $\text{Pb}^{2+} \rightarrow (\text{Pb}^+)^* \rightarrow \text{Pb}^+$
102 (Nagli and Dyachenko, 1986; Erfurt, 2003). A similar transition has been suggested for amazonite (see Ostrooumov, 2016),
103 but the direct connection between the Pb-centre and IR-RF has not yet been evidenced. Other reactions involving Pb^{4+} would
104 also be possible but haven't yet been formally proposed. Additionally, the IR-RF signal is composed of at least two separate
105 emissions. Previous publications placed the main IR emission at 1.43 eV (865 nm) based on Trautmann et al. (1999a, b) and
106 Erfurt and Krbetschek (2003), but more recent work including corrections for the spectrometer efficiency places the IR
107 emission closer to 880 nm (Kumar et al., 2018; Riedesel et al., 2021; Sontag-González et al., 2022). A second IR emission
108 centred at 955 nm (1.30 eV) at lower intensity has also been identified (Kumar et al., 2018), which partly overlaps with the
109 880 nm peak.

110 The presence of iron (Fe) in feldspar is known to lead to red RF (e.g., Telfer and Walker, 1978; Brooks et al., 2002;
111 Visocekas et al., 2014), with the maximum peak wavelength varying between 700 nm and 770 nm depending on feldspar
112 composition (Dütsch and Krbetschek, 1997; Krbetschek et al., 2002). Such observations are in line with the suggestion of more
113 than one component in the red photoluminescence of K-feldspar (Prasad and Jain, 2018). Despite the occurrence of the red RF
114 emission in Fe^{3+} state, its initial state remains a subject of many debates, with conflicting opinions suggesting either Fe^{2+} (Fe^{2+}
115 + h \rightarrow Fe^{3+} ; here h stands for hole) or Fe^{4+} (Fe^{4+} + $e^- \rightarrow \text{Fe}^{3+}$; here e^- stands for electron) (Kirsh and Townsend, 1988; Jain et
116 al., 2015). Recently, Kumar et al. (2020) argued that the initial state must be Fe^{4+} based on their findings using
117 cathodoluminescence microscopy. Spectral analyses showed that, with dose exposure, the red RF emission (~ 710 nm emission
118 in K-feldspar) increases, while the 880 nm emission decreases (Krbetschek et al., 2000; Erfurt and Krbetschek, 2003; Kumar
119 et al., 2018; Frouin et al., 2019). The thermal stability of the ~ 710 nm emission has been, however, questioned (e.g., Krbetschek
120 et al., 2000). Such a reduced thermal stability might be an issue for IR-RF dating, as it has been suggested that the tail of the
121 ~ 710 nm emission overlaps with the 880 nm emission, thus potentially playing a role in the shape of the measured IR-RF.
122 Such a contribution can be reduced to less than 5% of the IR-RF signal by using a bandpass filter centred at 850 nm (FWHM
123 40 nm), but can still affect the equivalent dose (D_e) value (see Sontag-González and Fuchs, 2022). D_e values are determined
124 by sliding the IR-RF dose-response curve of grains containing the natural signal onto that obtained after a full bleach of the
125 same aliquot. In summary, although previous studies have identified factors that may influence the IR-RF signal in several
126 ways, e.g., whether the IR-RF signal originates from Pb, and is affected by the presence of Fe^{2+} or Fe^{4+} , a conclusive
127 confirmation or comprehensive linkage between these factors is yet to be established.

128 A possible variability of the several RF emissions in individual grains has received scant attention, so far. Trautmann
129 et al. (2000) were the first to analyse the IR-RF signal of individual K-feldspar grains. Using spectral measurements on ~~twenty-~~

130 ~~one of 21~~ to ~~forty two~~42 grains from three samples, they observed up to four emissions (IR, red, yellow, blue) with variable
131 intensities (a fourth sample appears in their figure 3 but is not mentioned in the main text). An IR-RF dose-response curve was
132 only reported for one grain, which had a similar shape, albeit a later onset of saturation, when compared to the response from
133 the multi-grain aliquot of the same sample. More recently, Mittelstraß and Kreutzer (2021) analysed ~~sixty~~60 grains from two
134 samples, of which 55% and 80% emitted a detectable signal. In that study, between one and three grains per sample (~9% of
135 signal-emitting grains for both samples) were rejected due to a bad match between the natural and regenerative curves, which
136 might have been caused by equipment issues, but also due to sensitivity changes (Varma et al., 2013). However, all grains that
137 emitted a detectable signal displayed the expected decay shape: for IR-RF (decreasing signal with increasing dose). Likewise,
138 ~~our own~~ laboratory observations ~~indicate~~indicated that the signal varies in sensitivity across feldspar minerals and can be
139 contaminated for various reasons, leading to spectral interference or quenching, ultimately influencing the saturation level
140 and/or the shape of the IR-RF signal (Frouin et al., 2017, 2019; Kumar et al., 2020).

141 ~~Previous publications placed the~~To investigate these issues, ~~first~~IR emission at 1.43 eV (865 nm) based on Trautmann
142 et al. (1999a, b) and Erfurt and Krbetschek (2003), but more recent work including corrections for the spectrometer efficiency
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144 IR emission centred at 955 nm (1.30 eV) at lower intensity has also been identified (Kumar et al., 2018). ~~With dose exposure,~~
145 ~~the 955 nm emission increases and overlaps with the 880 nm peak.~~

146 ~~Previous spectral analyses of K-rich feldspar indicated that the IR-RF emission occurs as a result of the change in the~~
147 ~~oxidation state of the participating defect via the transition: $Pb^{2+} \rightarrow (Pb^+)^* \rightarrow Pb^+$ (Nagli and Dyachenko, 1986; Erfurt, 2003).~~
148 ~~A similar transition has been suggested for Amazonite (see Ostrooumov, 2016), but the direct connection between the Pb-~~
149 ~~centre and IR-RF is not yet evidenced. Similarly, the presence of Fe in feldspar is also known to lead to a red RF emission~~
150 ~~(e.g., Telfer and Walker, 1978; Brooks et al., 2002; Visocekas et al., 2014), with the maximum peak wavelength varying~~
151 ~~between 700 nm and 770 nm depending on feldspar composition (Krbetschek et al., 2002). Despite the occurrence of the red~~
152 ~~RF emission in Fe^{3+} state, its precise origin remains a subject of debate, with conflicting opinions suggesting Fe^{2+} ($Fe^{2+} + h \rightarrow$~~
153 ~~Fe^{3+} ; here h stands for hole) or Fe^{4+} ($Fe^{4+} + e^- \rightarrow Fe^{3+}$; here e^- stands for electron) as the potential sources (Kumar et al., 2020).~~
154 ~~Furthermore, spectral analyses showed that, with dose exposure, the red RF emission (~710 nm emission in K-feldspar)~~
155 ~~increases, while the 880 nm emission decreases (Krbetschek et al., 2000; Erfurt and Krbetschek, 2003; Kumar et al., 2018;~~
156 ~~Frouin et al., 2019). The thermal stability of the ~710 nm emission has been, however, questioned (Krbetschek et al., 2000).~~
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160 ~~nm) (see Sontag-González and Fuchs, 2022). In summary, although previous studies have identified factors that may influence~~
161 ~~the IR-RF signal in several ways, e.g., whether the IR-RF signal originates from Pb, and is affected by the presence of Fe^{2+} or~~
162 ~~Fe^{4+} , a conclusive confirmation or comprehensive linkage between these factors is yet to be established.~~

163 Identification of the defect type linked to the IR-RF signal and its concentration will enable us to better characterise
164 the light emission (signal sensitivity) in different types of feldspar, while identification of the origin of possible contamination
165 in the IR-RF signal will help us to gain a better understanding of the apparent early saturation or quenching of the IR-RF
166 signal. The μ -XRF and μ -XANES techniques are best suited for this purpose by producing high resolution maps of elements
167 and their oxidation states. The use of synchrotron μ -XRF allows us to improve the spatial resolution compared with previous
168 uses of μ -XRF (e.g., Buylaert et al., 2018) by reducing the beam spot size from ~~25 μ m to 1~~ or 0.5 μ m. Though both the grain
169 geochemistry and crystallography should be investigated to characterise the defect type and its environment, in the present
170 study, we decided to focus only on geochemistry.

171 First, we recorded IR-RF curves from individual grains in our luminescence dating laboratory at the Research
172 Laboratory for Archaeology and the History of Art (RLAHA) at the University of Oxford (UK). Then, during our
173 ~~beamtime~~beam time (96 h), and as a proof of concept, we ~~optimised~~optimized the measurement conditions and obtained
174 compositional maps of the individual K-feldspar grains. We paid particular attention to K, Ca, Fe and Pb. After analysing the
175 μ -XRF maps, μ -XANES measurements were done at selective spots where Fe and Pb ~~occur~~occurred in greater concentrations.
176 Note: The atomic number of sodium (Na; another end member of the feldspar ternary system) is too low to be measured at the
177 current SRX beamline.

178 3 Material and instrumentation

179 A total of five samples were selected to represent a diversity of i) geological context, ii) geochemistry, iii) shape of the IR-RF
180 signal, and iv) age. Sample Gi326 from a Triassic sandstone from Bayreuth, Germany, is composed of 89% of K-rich feldspar
181 (Sontag-González and Fuchs, 2022) and has previously been used as a reference sample in a laboratory comparison of IR-RF
182 dating (Murari et al., 2021b). X7343 was collected from a Pliocene sediment at the Nyayanga site in the Homa peninsula,
183 Kenya (Plummer et al., 2023). X7363 was taken from the Gele Tuff in the Turkana Basin, Kenya, and dated by ~~Argon-~~
184 ~~Argon~~Ar/Ar at 1.32 ± 0.2 Ma (Phillips et al., 2023). X7368 is a sediment sample collected above the Silbo Tuff (0.751 ± 0.022
185 Ma, McDougall and Brown, 2006) and below the Kale Tuff (~~undated~~younger than the Silbo Tuff but not directly dated) in the
186 Turkana Basin, Kenya. Sample H22550 is a coastal marine sample from Sula, Russia, dated by quartz single-aliquot-
187 regenerative optically stimulated luminescence (OSL) at 103 ± 8 ka (Murray et al., 2007) and was used in the past as a reference
188 sample to test the accuracy of IR-RF dating (Buylaert et al., 2012). All samples were prepared following conventional
189 treatments (e.g., Preusser et al., 2008), including wet-sieving to isolate the desired grain size fraction, chemical treatment with
190 HCl at 10% to remove carbonates and H_2O_2 at 30% for a few hours to a few days to remove organic matter, and density
191 separation at 2.58 g cm^{-3} using a heavy liquid solution to enrich K-rich feldspar grains. Sample H22550 was then etched with
192 diluted HF (10%, 40 min). All ~~the~~ grains were exposed under a solar simulator SOL ~~Heinle~~Hönle 2 for a few days to reset their
193 signal.

194 ~~Grains~~Unmeasured grains of sample X7343 were placed on a stub mount on a piece of carbon tape, then imaged with
195 a scanning electron microscope (SEM) equipped for energy-dispersive X-ray spectroscopy (EDS) at Archéosciences Bordeaux,
196 France (FR) (JEOL JSM-6460LV; detector: Oxford Instruments X-Max (51-XXM0002); software: Oxford Instruments INCA
197 version 4.11). The SEM was operated at 20 kV ~~voltage~~ and 55 μ A beam current. Sample X7343 is referred to as BDX22338
198 in the Archéosciences Bordeaux system.

199 IR-RF measurements were recorded with a Lexsyg/lexsyg research luminescence reader fitted with an annular $^{90}\text{Sr}/^{90}\text{Y}$
200 beta source (Richter et al., 2013) using a bandpass filter centred at 850 nm (FWHM 40 nm) mounted in front of a Hamamatsu
201 H7421-50 photo-multiplier tube. Measurements were performed at 70°C, following Frouin et al. (2017). Multi-~~grain~~grains and
202 single-grains were measured on stainless steel cups. High-resolution compositional analysis of the grains was undertaken at
203 the ~~5-ID-SRX~~ SRX beamline at ~~the National Synchrotron Light Source II (NSLS-II) at Brookhaven National Laboratory~~ (Chen-
204 Wiegart et al., 2016). ~~Grains-After IR-RF measurements, the grains~~ were removed from the stainless steel cups and fixed on a
205 polymer microscope slide (UVT acrylic; Agar Scientific) with a small piece of carbon tape to avoid misplacement during
206 measurement (supplementary ~~figure 1~~-Fig. S1). μ -XRF maps were obtained by scanning across pre-selected regions on the
207 grains with low topographic changes (90 x 90 μm maps, with a step size of 0.67 μm and an integration time ~~of~~ 0.1 s). ~~A~~
208 ~~resolution of 0.67 μm -The incident X-ray beam~~ was achieved/focussed by ~~focusing the beam with a a~~ pair of Kirkpatrick-Baez
209 mirrors. An incident beam energy of 13.5 keV was used for the μ -XRF measurements. The excited elements' characteristic
210 fluorescence was ~~detected through~~obtained from the sum of ~~4the four elements of a~~ silicon drift ~~detectors~~detector. All μ -XRF
211 measurements were ~~normalised~~normalized to the corresponding incident X-ray flux (I_0) (supplementary ~~figure 2~~-Fig. S2). ~~sa~~

212 The μ -XANES maps ~~had a resolution~~cover 60 x 60 μm in steps of 0.5 μm (~~60 x 60 μm), thus creating a grid with~~
213 120 x 120 data points (i.e., 14 400 μ -XRF spectra). To ~~do μ -XANES mapping~~obtain μ -XANES maps of Fe-states in our
214 samples, we varied the incident beam energy according to the absorption edge values obtained from the ~~μ -XANES~~ μ -XANES
215 spectral measurements of Fe standards (Fe foil, pyrite, hematite). The ~~μ -XANES~~ μ -XANES maps were measured three times
216 to obtain ~~μ -XRF~~ μ -XRF emission spectra restricting the Fe species to either (i) the total Fe (at 7.275 keV), (ii) the sum of Fe³⁺
217 and Fe²⁺ (at 7.134 keV), and (iii) only from Fe²⁺ (at 7.122 keV). The difference between the intensity levels of the latter two
218 measurements can qualitatively give the intensity levels of Fe³⁺, i.e., $I_{\text{Fe}^{2+} \& \text{Fe}^{3+}} - I_{\text{Fe}^{2+}} = \text{Fe}^3 I_{\text{Fe}^{3+}}$ where I (a.u.) refers to intensity,
219 thus, the ~~μ -XANES~~ μ -XANES map of Fe³⁺. We also attempted to record Pb states, however, the Pb standard available at ~~5-~~
220 ~~ID~~the SRX beamline was fully ~~oxidised~~oxidized, which hindered establishing the correct beam energy for mapping. Therefore,
221 no Pb oxidation state maps were possible. μ -XRF and μ -XANES data were analysed using the open-source software PyXRF
222 ~~v.1.0.23~~1.0.23 (Li et al., 2017) and ATHENA ~~v.0.9.26~~0.9.26 (Ravel and Newville, 2005), respectively. Maps and plots were created
223 using R (R Core Team, 2022).

224 4 Results

225 4.1 Multi-grain IR-RF signal

226 The IR-RF signal of a multi-grain aliquot of 8 mm diameter of sample X7343 was first measured. The aliquot contained
227 hundreds of grains. The expected IR-RF signal of K-~~rich~~-feldspar grains is a decaying function, e.g., a stretched single-
228 exponential (Erfurt et al., 2003). For sample X7343, however, we observed an unexpected shape of the IR-RF-~~measurements~~,
229 consisting of a signal decrease until 500 Gy succeeded by an increase, roughly following a saturating exponential shape that
230 keeps increasing beyond ~3800 Gy. The regenerative signal for one representative aliquot is shown in figure 1 (top right). ~~To~~
231 ~~remove this signal contamination, We hypothesized that the unexpected signal increase originates from a different source,~~
232 potentially ~~coming from a~~ coating around the grains, ~~we used due to the observation of a pinkish/reddish hue on some grains.~~
233 ~~Clay, Fe-oxide or other grain coatings are a common occurrence and additional preparation steps are sometimes undertaken to~~
234 ~~remove them prior to luminescence measurements (e.g., Jayangondaperumal et al., 2012; Lomax et al., 2007; Rasmussen et~~
235 ~~al., 2023). We attempted to remove this signal contamination, using~~ different chemical treatments such as HF, regal water, and
236 heated regal ~~water~~, however, without success. Therefore, we decided to investigate the mineral composition of sample X7343,
237 using SEM-EDS on ~~a hundred~~118 randomly selected grains. Despite using density separation to isolate K-~~rich~~-feldspar grains
238 during chemical pre-treatment, we found that this sample was mainly composed of low-K grains (Fig. 1; top left). Indeed, over
239 half of the grains had K-contents less than 2% and less than 5% of the grains had K-contents above 11%. The remainder
240 ~~had exhibited~~ K-contents between 2% and 10%.- Note that a K-feldspar end member is 14% K (e.g., Gupta, 2015). The low-K
241 grains, which correspond to the majority of grains, also had high Fe-contents of ~10%.

242 We then tested whether it was possible to isolate the desired decreasing IR-RF signal by handpicking grains based on
243 their visual appearance under a microscope. Between ~~ten~~10 to ~~thirty~~30 grains were placed onto two aliquots, one for transparent
244 shiny angular grains and one for white-pinkish rounded grains. The regenerated IR-RF signals showed a clear distinction
245 between the two aliquots (Fig. 1), proving it is possible to separate the two observed IR-RF shapes.

246 By manually selecting the grains based on their shape and colour, we made three important observations: i) The
247 decreasing IR-RF signal originates from a small number of grains (less than 5%), presumed to be K-~~rich~~-feldspar. ii) The IR-
248 RF signal of these grains ~~decreases~~decreased beyond 3800 Gy without reaching a plateau, indicating that a dose could be
249 estimated beyond that value. iii) The increasing IR-RF signal originates from a different subset of grains, presumed to be the
250 low-K, Fe-rich minerals identified via SEM-EDS.

251 4.2 Single-grain IR-RF characterisation

252 To further investigate this phenomenon, we measured the IR-RF signal of ~~twenty-two~~22 individual grains ~~coming~~ from five
253 samples ~~from of~~ different origins (between one to eight grains per sample, Table 1). Each grain was manually placed on a
254 sample holder (cup) and their signal was recorded over a ~~32653,265~~ Gy beta irradiation. For each grain, their IR-RF signal
255 shape falls into three categories (Fig. 2): Category #1 for grains with a decreasing IR-RF signal, category #2 for grains with

an increasing IR-RF signal, and category #3 for grains with a flat signal. Within categories #1 and #2, the saturation level of the individual grains varies (Fig. 2b, d).

Among the five samples, one is a tuff, and two are originated from nearby volcanic environments and might, thus, be expected to yield abnormal ~~behaviour. However, we also observed the unwanted decreasing~~ luminescence behaviour. Common issues with volcanic samples are dim signals, different proportions of emissions, high fading rates and complex grain mineralogy (e.g., Krbetschek et al., 1997; Guérin and Visocekas, 2015; Joordens et al., 2015; Sontag-González et al., 2021; O’Gorman et al., 2021). However, we also observed the unwanted increasing IR-RF signal for one of the four grains for sample H22550, which is from a coastal sand deposit. The significance of this find is illustrated in figure 3, where curves representing the signals of individual grains from categories #1 and #2 were added together to simulate a multi-grain aliquot. We used the curves obtained from fitting a single stretched exponential decay function to the ~~normalised~~normalized data of one grain of sample X7343 (category #2) and one of Gi326 (category #1), since no category #1 grain was measured for sample X7343. When the total signal of the theoretical aliquot was composed of more than 50% of signal from the category #2 grain, we observed the same decay shape as in figure 1 for a multi-grain aliquot of sample X7343. Importantly, a synthetic mixture containing 20% of grains from category #2 still displayed the decaying shape characteristic of category #1 grains. However, the curvature of its dose-response curve was altered, i.e., saturating earlier than the ‘pure’ grain. Possible differences in long-term signal stability between the two grain categories could cause differences between the ~~total~~summed curves of natural and regenerated IR-RF signals and thus lead to inaccurate equivalent doses for these mixtures.

Further, our results demonstrate that a satisfying IR-RF signal can be measured for all our samples, but only by selecting grains with the appropriate luminescence characteristics (presumably K-rich feldspar grains, while other (presumed)). We hypothesize that the low-K grains ~~can contaminate the~~ with high Fe-content are the source of a contaminant IR-RF emission ~~and, which if not removed~~ might result in a wrong equivalent dose estimation (i.e., a wrong age estimate).

4.3 ~~Subgrain μ -XRF~~Sub-single grain μ -XRF elemental maps

~~We~~Utilising μ -XRF, we identified up to fifteen elements in the grains (see supplementary figure 3 for the total μ -XRF spectra). We then fitted each of the spectra in the 135 by 135 pixel grid (i.e., 18 225 spectra) for each grain to obtain maps describing the XRF intensity of each identified element. These maps only serve as qualitative indicators for the presence of elements and ~~do~~are not ~~consider~~corrected for the element-specific emission intensity or the energy-dependent efficiency of the detectors. For two grains, we recorded additional μ -XRF ~~maps to~~ characterisecharacterize visible inclusions (see ~~Table~~table 1).

First, we consider only the presence/absence of each element with the IR-RF signals previously obtained. Among the grains displaying a decreasing IR-RF signal (category #1), all contain K, Pb, Fe and Ba, among other elements (Fig. 4, middle). Most of the grains from categories #2 and #3 also contain K, Pb and Fe, but few contain Ba. Further, most grains from categories #2 and #3 contain Ca, Ti and Mn, which are rare in the grains from category #1. The μ -XRF intensities also allow for a qualitative comparison of elemental composition. As shown in the boxplots in figure 4 (right-hand side), category #1

288 grains differ from those in category #2 and #3 primarily by a higher μ -XRF signal contribution from K and a lower contribution
289 from Fe.

290 ~~The XRF intensities allow for a qualitative comparison of elemental composition.~~ If we compare the relative
291 intensities of Pb, Fe and K, we can identify a pattern in the composition of grains from each category (Fig. 5). Grains from
292 category #1 tend to have high proportions of K and Pb and medium-to-low proportions of Fe. All grains from categories #2
293 and #3 have medium-to-high proportions of Fe, and most have low levels of K and Pb. No grains from categories #2 or #3
294 have high levels of both K and Pb. The element that distinguishes grains from categories #2 and #3 is Ca, which is only present
295 in category #2 grains to a high proportion (see [also](#) supplementary figure 3).

296 Interestingly, the grains from category #3 cluster relatively close to those from category #2, suggesting that the
297 elemental composition alone is not responsible for the lack of IR-RF signal but that the spatial configuration of the elements
298 is another important factor to consider. This is exemplified in figure 6, which contains elemental maps of K, Pb and Fe for one
299 grain of each category. The category #1 grain (top row) tends to have K and Pb ~~co-localised~~localized (overlap shown in green),
300 whereas the category #3 grain (bottom row) also contains both elements, but they appear in separate locations (shown in cyan
301 and yellow). In ~~fact, in~~ this grain, K appears ~~co-localised~~localized with Fe (overlap shown in dark blue). Our preliminary
302 observations ~~must still be confirmed for a larger number of grains~~require broader confirmation, but they are in line with the
303 current hypotheses of emission origins for the K-feldspar IR-RF signal decreasing with dose (due to Pb) and a contaminating
304 red RF signal increasing with dose (due to Fe). Furthermore, the lack of an IR-RF signal in category #3 grains appears to stem
305 from low levels of K co-existing with high proportions of Fe, supporting the observations made by Kumar et al. (2020).

306 **4.4 Mapping oxidation states with μ XANES/ μ -XANES**

307 By analysing absorption of X-rays near the absorption edge, μ XANES/ μ -XANES spectra can provide information on the
308 presence of potential oxidation states of an element, as shown in figure 7 through measurements of different standards of Fe;
309 an increase in the oxidation state is generally accompanied by a shift in the absorption edge to higher energy (Fig. 7, inset).

310 We targeted the Fe-rich region of a category #1 grain (X7368) for mapping- (location shown by the blue square in
311 figure 6). Figure 8 shows the μ XANES/ μ -XANES maps of three oxidation states (Fe³⁺ and Fe²⁺ combined, Fe²⁺ and Fe³⁺), -all
312 ~~normalised~~normalized to the maximum intensity level of the total Fe map. These maps suggest that Fe exists ~~on~~in the
313 surface/top section of this feldspar grain in its Fe³⁺ and Fe²⁺ states. Note how Fe²⁺ is mainly clustered in one region, possibly
314 within a mineral inclusion with a rim of Fe³⁺.

315 **5 Conclusions and future work**

316 We demonstrated that individual K-feldspar grains of the same five samples display different IR-RF behaviour, illustrated by
317 different signal decay shapes (i.e., increasing or decreasing with dose and different saturation levels). These behaviours are
318 cumulative (see figure 1), and therefore, the IR-RF signal of a multi-grain aliquot can lead to inaccurate equivalent doses.

319 Despite the use of chemical preparations to remove contaminants, manually picking individual grains was necessary to isolate
320 K-feldspar grains, which is unrealistic for routine dating applications in a low-light laboratory. A more realistic way to remove
321 such contamination is by selecting K-rich feldspar grain populations by isolating the emission signal of individual grains with
322 an imaging system. Here, we wanted to gain a further understanding of the production and origin of the emission signal, which
323 ultimately will help us design a more appropriate imaging system for IR-RF dating. For sample X7343, we show through
324 SEM-EDS analyses that the different emissions can be linked to different grain mineralogy. Since K-feldspar grains are known
325 to be heterogeneous on a subgrain level, we propose synchrotron-based X-ray spectroscopy to ~~characterise~~characterize the
326 grains on a submicron scale and investigate the origin of the IR-RF and other linked emissions. Information on the oxidation
327 states of, e.g., Fe ~~and Pb~~, possibly allow~~s~~ for the characterisation of the reactions behind the electronic changes leading
328 to radiofluorescence.

329 In the preliminary work presented here, we successfully applied ~~μ -XRF~~ μ -XRF and ~~μ -XANES~~ μ -XANES at the SRX
330 beamline (NSLS-II) to obtain ~~mineralogical~~element and oxidation state maps of regions of interest within individual K-feldspar
331 coarse grains previously used for IR-RF measurements. We were able to correlate the desired IR-RF signal shape (category
332 #1) with compositions of high proportions of K, Pb, and Ba and low proportions of Fe. High proportions of Fe in the ~~μ -XRF~~ μ -
333 XRF spectra were found in grains of categories #2 and #3, but the possible role of Fe as a contaminant remains unclear. During
334 our next ~~beamtime~~beam time, we will polish the grains down to a uniform surface prior to ~~μ -XRF~~ μ -XRF and ~~μ -XANES~~ μ -
335 XANES measurements to ~~avoid~~limit surface effects. Such a setup will also allow us to test the hypothesis that the
336 contaminating IR-RF signal is coming from an element present at the surface of the grain (e.g., iron coating possibly due to
337 weathering), but not within the grain.

338 The relation between the chemical composition, crystal structure, and the shape of the IR-RF signal in individual K-
339 feldspar grains is still poorly understood, and efforts should be made to identify and quantify at high resolution the element
340 responsible for producing the IR-RF signal with the highest dynamic range (i.e., saturation at high dose). Our future work will
341 include implementing a second detector to simultaneously measure ~~μ -XRF/ μ -XANES~~ μ -XRF/ μ -XANES and the IR-RF signal
342 induced by the X-rays. Though not widely used, X-rays are a suitable alternative to radioactive sources for luminescence
343 dosimetry including ~~radiofluorescence~~RF. The dual detection will allow us to isolate emissions from different mineral
344 inclusions and directly correlate them to the elemental composition, thereby assessing the extent of overlap of the desired IR-
345 RF emission centred at 880 nm and contaminating ones such as the possible unstable red emission associated with Fe³⁺.

346 **Data availability**

347 The SEM dataset and the original data used to produce ~~μ -XRF~~ μ -XRF maps are available online (Sontag-González et al., 2023).

348 **Author contribution**

349 MF, JT, RK and JLS designed the experiments and prepared the samples. RK, [JLS](#), and MF carried out the IR-RF
350 measurements. SK ~~organised~~[organized](#) and analysed the SEM EDS measurements. MF and JT carried out the [μXRF](#)-XRF
351 and [μXANES](#)-XANES measurements. RK, MSG and MF analysed the results. MSG and RK prepared the manuscript with
352 contributions from all authors. MF, JT and JLS obtained funding.

353 **Competing interests**

354 The authors declare that they have no conflict of interest.

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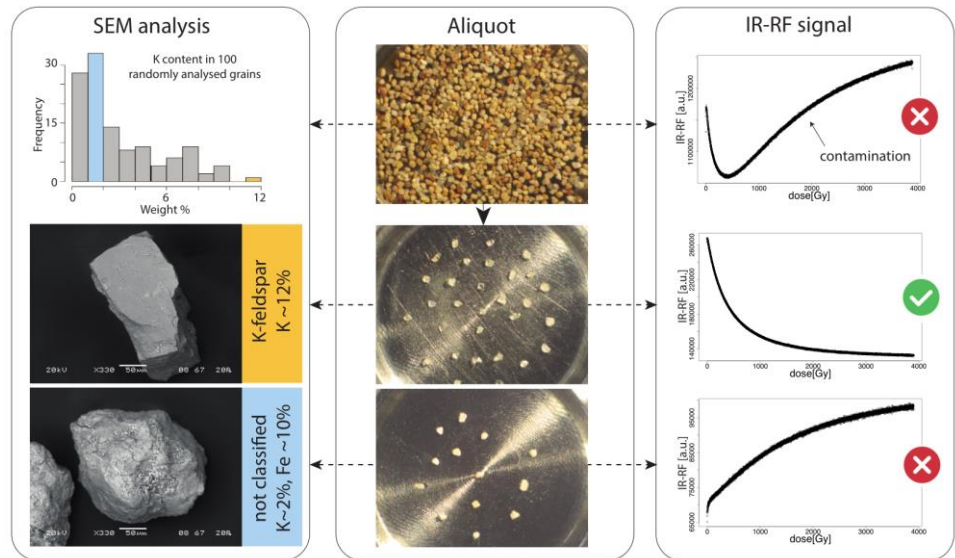
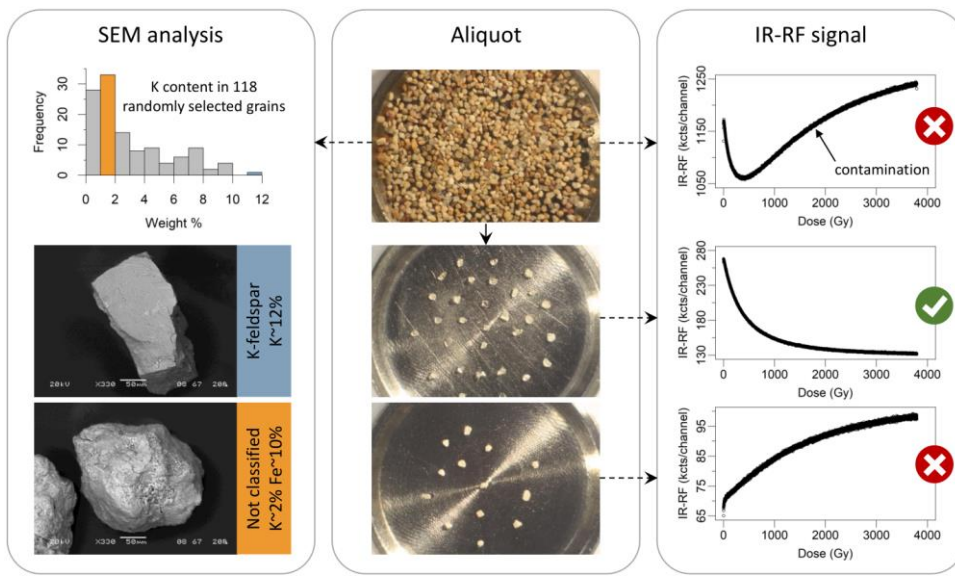
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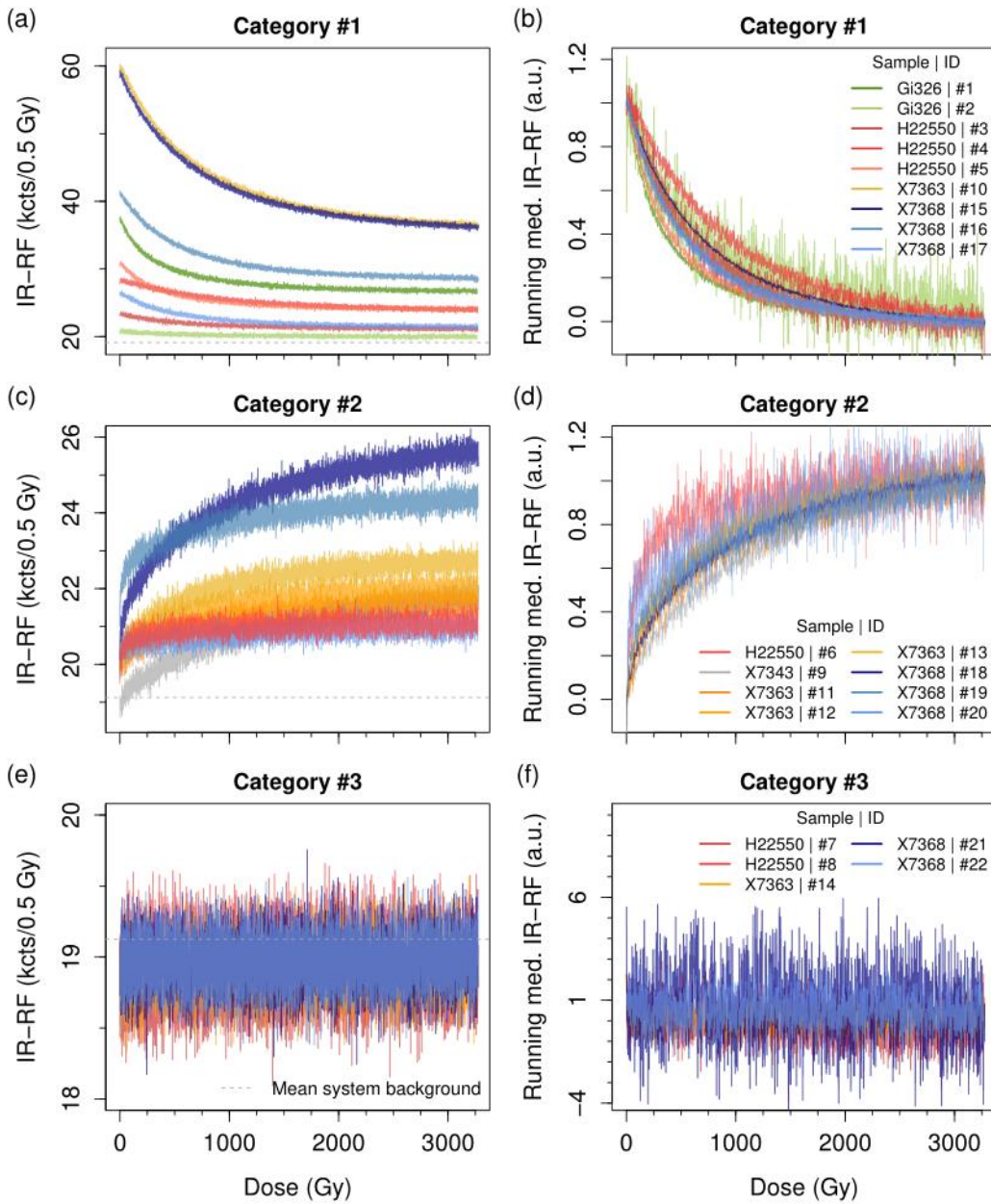
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519 **Figure 1:** Illustration showing how contamination of the IR-RF signal can be removed by selecting only K-rich feldspar grains from sample
520 X7343. The regenerative IR-RF curves were obtained from aliquots containing hundreds of unsorted grains (top) or 10–30 grains manually
521 sorted into transparent shiny angular grains (middle) or white-pinkish rounded grains (bottom). The histogram shows the K-content
522 determined by SEM-EDS for ~~100~~118 grains (not measured for IR-RF). Representative examples of grains classified as K-rich and
523 contaminating Fe-rich grains are shown.



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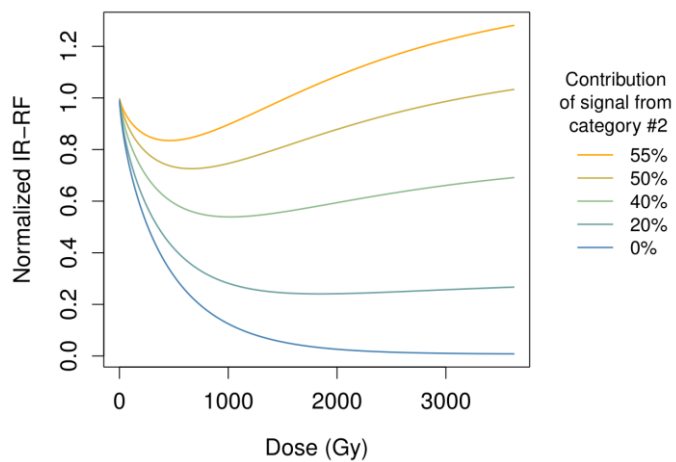
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Figure 2: IR-RF dose-response curves of individual grains obtained after bleaching. Categories #1–#3 refer to grains with decreasing, increasing or no detectable signal, respectively (one category per row). The curves are shown (a, c, e) unnormalised and without background correction and (b, d, f) with intensities normalised to the signal maxima (defined as the median value of (b, f) the initial and (d) the final 20 channels) after subtracting as background the minimum signal of each grain (defined as the median value of (b, f) the final and (d) the initial 20 channels). For better visualisation, the normalised plots show the running median IR-RF with a window of 7 values. The system background was determined as the mean value obtained from measuring five empty cups under the same conditions as the grains.

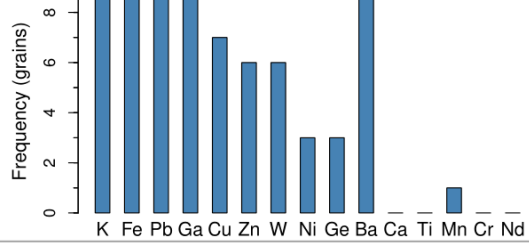
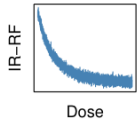
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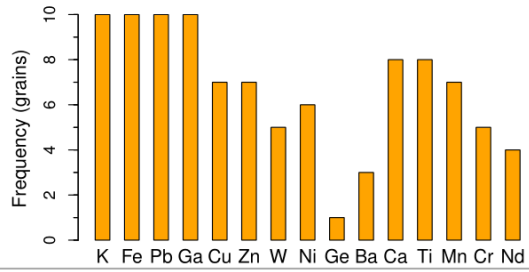
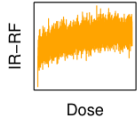
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535 **Figure 3:** Simulated dose-response curves of theoretical aliquots varying the proportion of grains from categories #1 (desired decreasing
536 signal) and #2 (increasing signal). The curves are the sum of two stretched exponentials using parameters obtained from fits of grains from
537 samples Gi326 (category #1) and X3743 (category #2). The higher the signal contribution from category #2 grains, the more aberrant the
538 sum curve becomes.

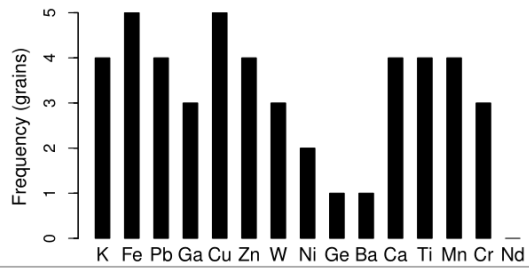
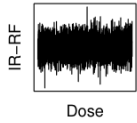
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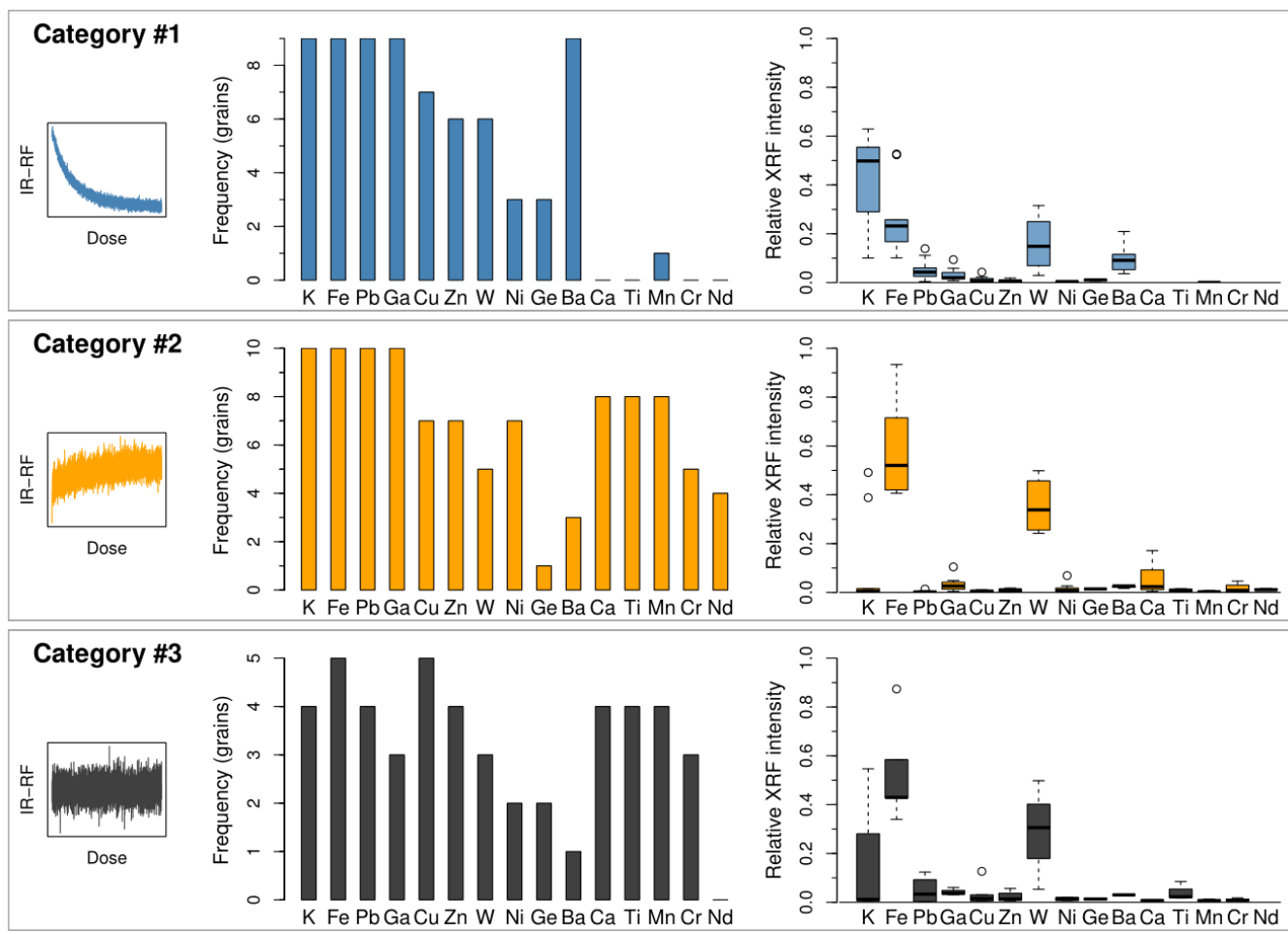


Category #2



Category #3





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542 **Figure 4:** Bar charts of elements identified in μ -XRF spectra and boxplots of the relative μ -XRF intensities for grains in three
 543 categories, as exemplified in the insets: decreasing IR-RF signal (category #1), increasing IR-RF signal (category #2) or flat IR-RF signal
 544 indistinguishable from the background (category #3) during beta irradiation.

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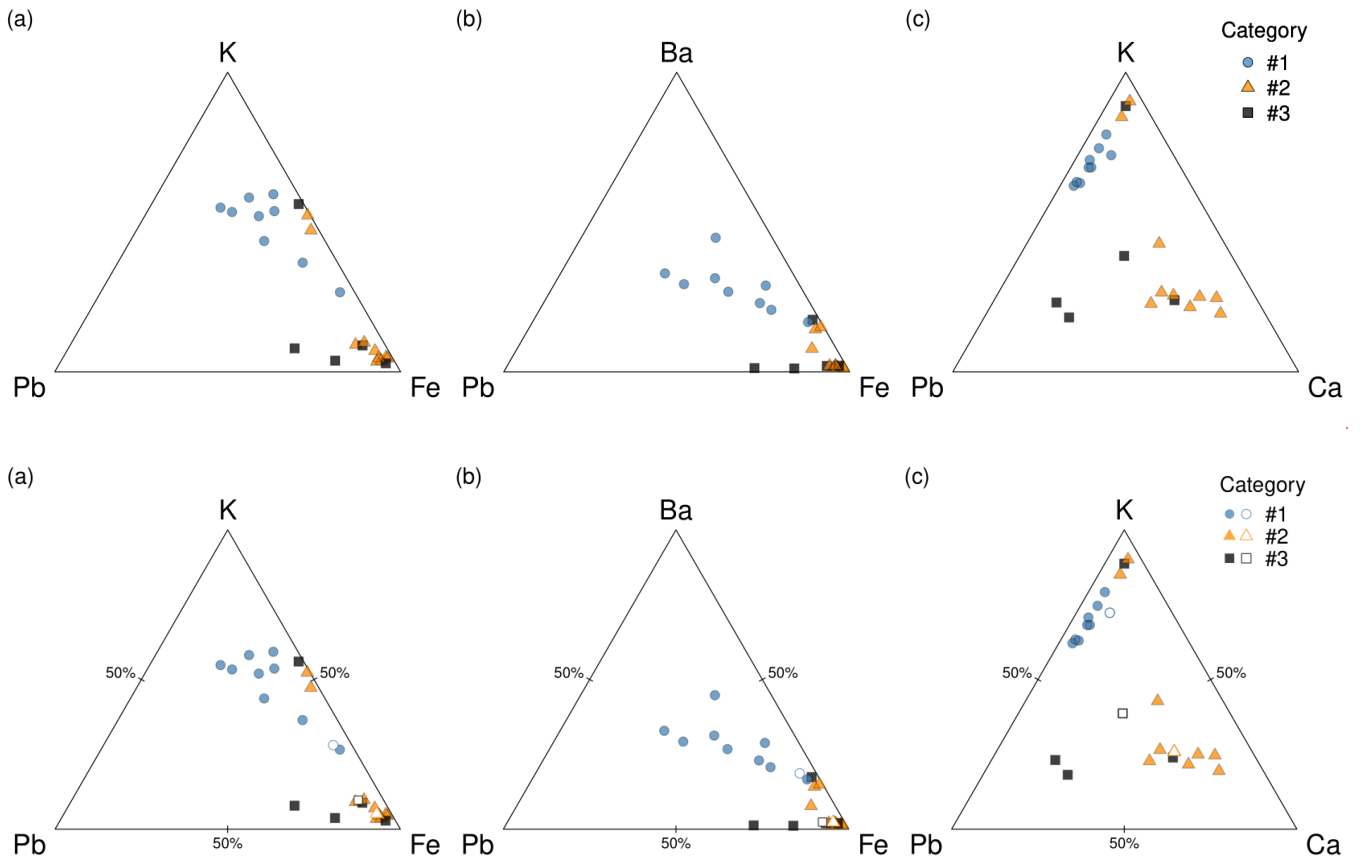
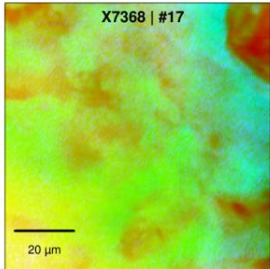
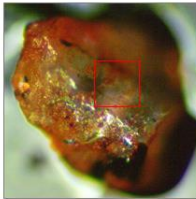
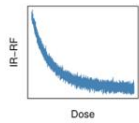
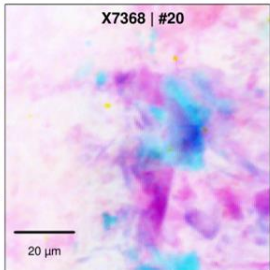
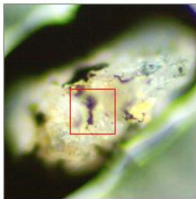
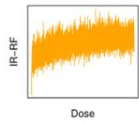


Figure 5: Ternary diagrams of relative μ -XRF intensities attributed to (a) K, Fe and Pb, (b) Ba, Fe and Pb, and (c) K, Ca and Pb for grains of the three categories. Note that the contributions are not calibrated to mass or stoichiometry. The relative K contribution is, thus, not directly comparable to the K-feldspar K-content. The three grains shown in figure 6 are marked as open symbols in each ternary diagram.

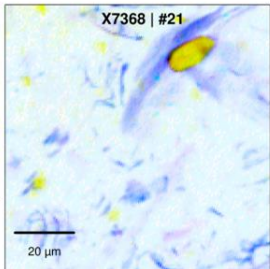
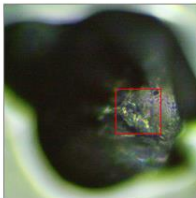
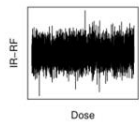
Category #1

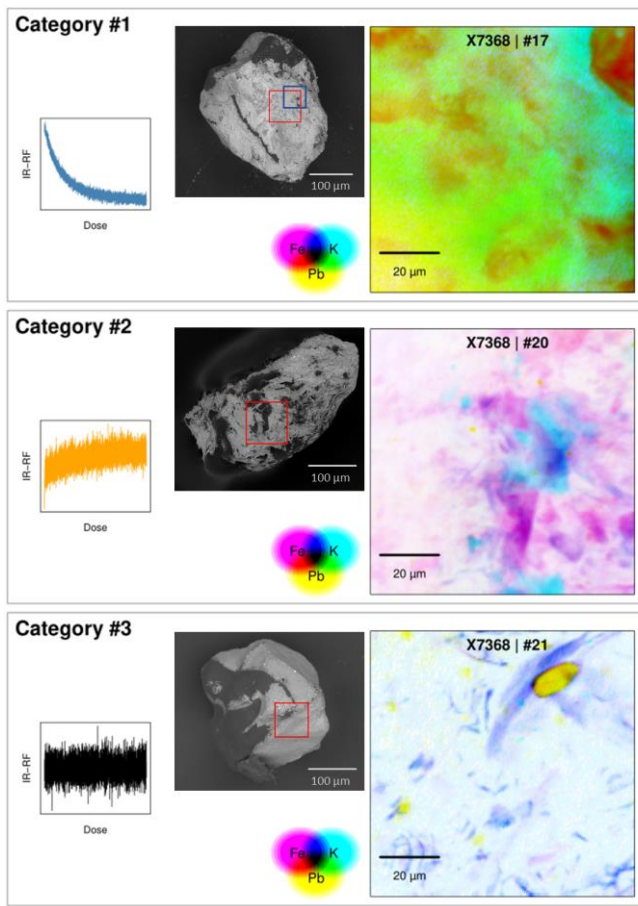


Category #2



Category #3





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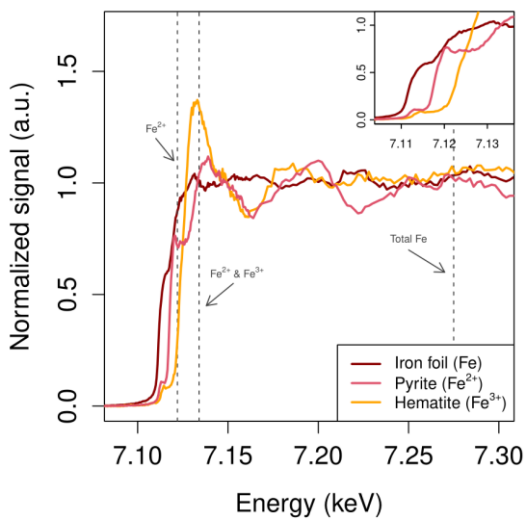
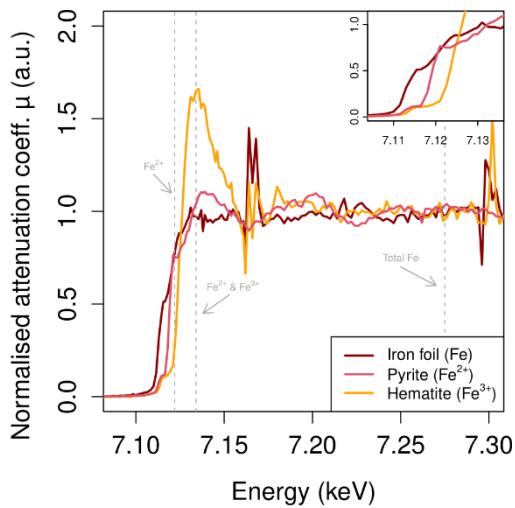
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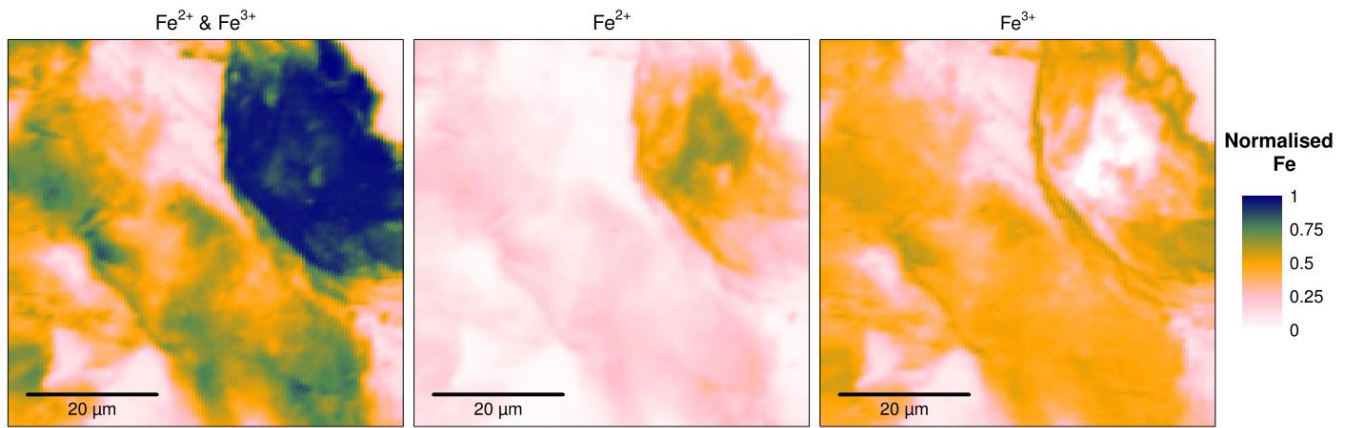
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Figure 6: Illustration showing three IR-RF curves obtained from three grains of sample X7368, classified as follows: decreasing IR-RF signal (category #1), increasing IR-RF signal (category #2) or flat IR-RF signal indistinguishable from the background (category #3) during beta irradiation. μ -XRF μ -XRF spectra were measured from the area bordered by red squares on the μ -XRF images of the grains. The maps show the presence of K, Fe, and Pb on the same grains as the IR-RF curves. The elemental compositions are shown overlaid, with the colour scales ~~normalised~~normalized to the maximum contribution of each element for each grain. The area bordered by a blue square in the category #1 grain corresponds to the map shown in figure 8.



561 **Figure 7:** μ XANES/ μ -XANES spectra of Fe standards. The dashed vertical lines indicate the incident beam energies necessary to isolate
 562 emissions from specific oxidation states. The inset shows a magnification of the energy region relevant to determine the incident beam
 563 energies.



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Figure 8: Maps of Fe oxidation states for a grain of sample X7368 (category #1; ID #17). Intensities are normalised to the maximum intensity of total Fe.

568 **Table1:** Overview of measured grains. Categories #1–#3 refer to grains with decreasing, increasing or no detectable signal, respectively.
569 For two grains, two regions each were mapped, so we measured a total of 24 μ -XRF μ -XRF maps.

Sample	Grain size (μm)	Number of measured grains		
		Category #1	Category #2	Category #3
Gi326	90–200	2	0	0
H22550	180–250	3	1*	2
X7343	180–255	0	1*	0
X7363	180–255	1	3	1
X7368	180–255	3	3	2
Total		9	8	5

570 *For these grains, two regions were mapped by μ -XRF μ -XRF: the grain ‘matrix’ and an inclusion.