Krypton-85 chronometry of spent nuclear fuel

Greg Balco¹, Andrew J. Conant², Dallas D. Reilly³, Dallin Barton³, Chelsea D. Willett¹, and Brett H. Isselhardt¹ ¹Lawrence Livermore National Laboratory, 7000 East Ave, Livermore CA USA ²Oak Ridge National Laboratory, Oak Ridge TN USA

³Pacific Northwest National Laboratory, Richland WA USA

Correspondence: Greg Balco (balco1@llnl.gov)

Abstract. We describe the use of the radionuclide ⁸⁵Kr, which is produced by nuclear fission and has a half-life of 10.76 years, to determine the age of spent nuclear fuel. The method is based on mass-spectrometric measurement of the relative abundance of fissiogenic Kr isotopes extracted from a fuel sample, and we show that it can be applied to micron-scale particles of spent fuel that are analogous to particles that have been released into the environment from various

5 nuclear facilities in the past. ⁸⁵Kr chronometry is potentially valuable for identifying and attributing nuclear materials, grouping samples into collections of common origin that can be used to reconstruct the origin and irradiation history of the material, and verifying declared nuclear activities in the context of international monitoring programs.

1 Introduction

In this paper we describe the use of the radionuclide krypton-85, which is produced by uranium and plutonium fission and

- 10 has a half-life of 10.76 years (ENDF/B-VIII.0; Brown et al., 2018), as a means of determining the age of spent nuclear fuel. The noble gases Kr and Xe are produced in large quantities by nuclear fission of U and Pu, comprising about 15% of all fission products. Because multiple fissiogenic isotopes of both elements provide a variety of diagnostic isotope ratios, and because their chemically inert nature after production facilitates transport and collection, noble gases produced and released by nuclear activities have been widely proposed and used as detection and monitoring tools. For example,
- 15 isotope ratios of Xe and Kr released when reprocessing nuclear fuel have been used for verifying that fuel history, usage, and plutonium production are consistent with declared nuclear activities (Hudson, 1993; Okano et al., 2006). In contrast, this paper focuses on Kr and Xe that have not been released to the atmosphere but are retained within irradiated nuclear fuels and therefore can be used to diagnose the origin, composition, and irradiation history of the material. In general, there exist a wide variety of isotope ratio signatures derived from actinides, decay products, and fission products that can
- 20 be used for 'nuclear forensics,' which is concerned with identifying and/or attributing nuclear materials of unknown origin (Moody et al., 2005; Fedchenko, 2015; Kristo et al., 2016). Some of these signatures are radiochronometers that provide information about the date that the material was produced, irradiated, and/or stored, and these are potentially valuable for investigative or forensic applications because they can be used to (i) identify or exclude potential sources of the material; (ii) provide information as to whether material found in the environment was derived from an active nuclear facility or

- from long-term storage; and (iii) provide a so-called 'position-independent signature' (e.g., Robel et al., 2018) that can be used to group samples that may have originated from different locations in the same fuel element or assembly. By grouping samples that were irradiated together, variations in irradiation conditions and history inferred from an array of coeval samples can be used to reconstruct reactor type and operating conditions (Dayman and Weber, 2018; Dayman et al., 2019; Savina et al., 2023). Furthermore, the ability to group samples into collections of common origin is particularly
- 30 important when working with fragments or particles, as might be available in a forensic investigation or an environmental contamination incident, rather than bulk quantities of fuel applicable in routine monitoring of reprocessing facilities.

Because its half-life is appropriate for events postdating the worldwide development of nuclear power in the 1950's, ⁸⁵Kr has been proposed as a useful chronometer for spent nuclear fuel (Okano et al., 2006; Park et al., 2010). It has several potential advantages for this purpose, including the relative simplicity of noble gas measurements, the routine availability of noble gas mass spectrometry systems, and the fact that noble gases can typically be extracted from solid samples by heating alone, potentially leaving the sample otherwise undamaged for further analysis. Here we (i) describe the theoretical basis of computing a ⁸⁵Kr age from measurements of Kr (and Xe) isotope ratios; (ii) apply the method to bulk and particulate samples from two example spent fuel rods; and (iii) discuss the relative merits of ⁸⁵Kr and other isotope ratio chronometers as dating methods and as means of grouping samples.

40 2 Theoretical basis of ⁸⁵Kr chronometry

Kr and Xe isotopes produced by nuclear fission include the radioisotope ⁸⁵Kr, three stable isotopes of Kr (83,84,86), four stable isotopes of Xe (131,132,134,136), and two short-lived Xe radioisotopes (133, with a half-life of 5 days, and 135, with a half-life of 9 hours). Fissiogenic Kr and Xe isotope ratios are variable among different fissioning nuclides (²³⁵U, ²³⁸U, ²³⁹Pu) and to a lesser extent with the energy spectrum of neutrons inducing fission (Table 1; Figure 1). Being inert

- 45 noble gases, Kr and Xe are somewhat mobile within fuel by thermally activated diffusion, and during fuel irradiation are known to both migrate from the fuel matrix into bubbles at grain boundaries and, to some extent, to escape entirely from fuel pellets. Commonly, 10-15% of total Kr and Xe produced during fuel irradiation is found to have been lost from fuel pellets into the plenum of spent fuel rods (Rest et al., 2019). Noble gas mobility, therefore, precludes computing a ⁸⁵Kr age simply by measuring the ratio of ⁸⁵Kr to its relatively immobile decay product ⁸⁵Rb. Instead, it is necessary to measure
- 50 the ratio of ⁸⁵Kr to one or more stable isotopes of Kr that are simultaneously produced and transported. Comparison of the measured isotope ratio to the initial ratio at the time of fission production provides an age.

There are two challenges in formulating a ratio of ⁸⁵Kr to other Kr isotopes for use in chronometry. First, after fissiogenic production, the abundances of some stable isotopes of Kr change due to neutron capture. In particular, conversion of ⁸³Kr to ⁸⁴Kr by neutron capture has a relatively high cross-section (Table 1), so the ⁸⁴Kr/⁸³Kr ratio increases significantly

55 with the total neutron fluence experienced during irradiation, and ⁸⁵Kr/⁸³Kr or ⁸⁵Kr/⁸⁴Kr ratios would therefore have a strong dependence on neutron fluence as well as material age and would not be useful for chronometry. Instead, we

use the ratio 85 Kr/(83 Kr+ 84 Kr), which is insensitive to neutron capture after production. Potentially, the ratios 85 Kr/ 86 Kr or 85 Kr/(83 Kr+ 84 Kr+ 86 Kr) could also be used in a similar way (Table 1), but for simplicity we discuss only 85 Kr/(83 Kr+ 84 Kr).



Figure 1. Variation in stable isotope ratios of fissiogenic Kr and Xe with fission source. The ratios ⁸⁶Kr/(⁸³Kr+⁸⁴Kr) and ¹³⁴Xe/(¹³¹Xe+¹³²Xe) are minimally influenced by neutron capture reactions after production and are therefore diagnostic of the fission reactions from which the gas was sourced. Panel (a) shows the expected isotope ratios calculated from cumulative fission yields for end-member fuel compositions and neutron spectra, as well as mixing lines between the end member compositions (see Cassata et al., 2023, for additional details on this diagram). The mixing lines are curved because fission yields are higher for Xe than for Kr. In other words, they represent mixtures of sources rather than mixtures of products: gas compositions produced over a long period of time from an evolving fuel composition could lie anywhere in the region spanned by the curves. Overall, possible fission gas compositions form an array between ²³⁵U and ²³⁹Pu end members, with secondary variability associated with neutron spectrum and contributions from fast neutron fission of ²³⁸U. Panel (b) shows the same predicted array with measured Xe and Kr isotope compositions in the samples analyzed in this study, which span a range of fission source from nearly pure ²³⁵U (for the 8%-enriched. low-burnup BR3 fuel) to a significant fraction of ²³⁹Pu (for the lower-enriched, high-burnup ATM-109 fuel). The results for the BR3 sample are from Cassata et al. (2023).

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Second, the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio at the time of Kr production by fission, as noted above, varies among different fission sources. The fuel in typical nuclear reactors is low-enriched uranium oxide (UO₂) with 5% or less 2^{35} U enrichment. When irradiation of fresh fuel begins, production is dominated by ²³⁵U fission. As irradiation proceeds and ²³⁵U is consumed, fission of ²³⁹Pu produced by neutron capture on ²³⁸U accounts for an increasing proportion of fission production of Kr. At high levels of fuel burnup ('burnup' is the cumulative fission energy production per unit mass of nuclear fuel). ²³⁵U is exhausted and Kr production is dominated by ²³⁹Pu fission, with a small contribution from fast neutron-induced fission of ²³⁸U. Thus, as irradiation proceeds, the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) value of newly produced Kr evolves from a value 65

characteristic of 235 U fission to a ~15% lower value characteristic of 239 Pu fission. However, it is possible to estimate the initial 85 Kr/(83 Kr+ 84 Kr) ratio applicable to Kr extracted from fuel by reference to another Kr isotope ratio, specifically 86 Kr/(83 Kr+ 84 Kr), which is also diagnostic of the fission source, also insensitive to neutron capture after production, and includes only stable isotopes so is not modified by radioactive decay after irradiation (Cassata et al., 2023, Figs 1,2).

- 70 This strategy is important because it allows a ⁸⁵Kr age to be computed from a single measurement of the Kr isotope composition of a gas sample: even though the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio varies with the relative contribution of different fissionable nuclides, it is possible to estimate it without independent knowledge of U and Pu isotope concentrations by exploiting its correlation with a stable isotope ratio that is diagnostic of the fission source.
- Thus, our overall procedure for computing a ⁸⁵Kr age is to (i) use the observed ⁸⁶Kr/(⁸³Kr+⁸⁴Kr) to estimate the composition of fissioning actinides responsible for Kr production and therefore the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio, and (ii) compare measured and initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) to obtain an apparent gas age. Alternatively, Xe and Kr isotopes are commonly measured simultaneously, so the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio can also be estimated from an equivalent sourcesensitive, non-neutron-capture-sensitive ratio of stable Xe isotopes, specifically ¹³⁴Xe/(¹³¹Xe+¹³²Xe) (Cassata et al., 2023, Figs. 1, 2). This is potentially advantageous because Xe is produced by fission in much greater abundance than Kr. so Xe isotope ratios are often measured more precisely.

Although it is evident from Figure 2 that knowing either the 86 Kr/(83 Kr+ 84 Kr) or 134 Xe/(131 Xe+ 132 Xe) ratio provides a fairly close estimate of the initial 85 Kr/(83 Kr+ 84 Kr), there is some variability associated with the neutron spectrum responsible for inducing fission. However, this variability is small for typical power reactors, in which fission is dominantly from thermal neutrons and therefore the initial ratios are fairly well approximated by the yields for thermal-neutron-induced fis-

85 sion shown in Figs 1 and 2. Thus, we use this scenario henceforth to estimate initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratios. Note that this would not be appropriate for unusual reactor types having an unmoderated fast neutron spectrum; initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratios from entirely fast-neutron-induced fission would be lower (Fig. 2). However, an unusual reactor type could easily be identified from other isotope ratio signatures.

Given the cumulative fission yields shown in Table 1 and production assumed to be entirely from thermal-neutroninduced fission, the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio *R*_{85,*i*} can be estimated from the measured ⁸⁶Kr/(⁸³Kr+⁸⁴Kr) ratio *R*_{86,*m*} by:

$$R_{85,i} = 0.08809R_{86,m} + 0.07145 \tag{1}$$

or from the measured 134 Xe/(131 Xe+ 132 Xe) ratio $R_{134,m}$ by a polynomial fit to the mixing curve shown in Figure 2:

$$R_{85,i} = 1.183 R_{134,m}^3 - 3.734 R_{134,m}^2 + 3.972 R_{134,m} - 1.243$$
⁽²⁾

95 The apparent ⁸⁵Kr age t_{85} is then:

$$t_{85} = \frac{-1}{\lambda_{85}} \ln\left(\frac{R_{85,m}}{R_{85,i}}\right)$$
(3)



where $R_{85,m}$ is the measured 85 Kr/(83 Kr+ 84 Kr) ratio and λ_{85} is the 85 Kr decay constant (yr⁻¹).

Figure 2. Panels (a) and (b) show the variation in the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio at production across a range of fission sources. As in Figure 1, these are calculated for cumulative fission yields for end member fuel compositions and neutron spectra; the symbols are the same in both figures. This ratio is correlated with the stable isotope ratios 86 Kr/(83 Kr+ 84 Kr) and 134 Xe/(131 Xe+ 132 Xe), so even without independent knowledge of the fuel composition, it is possible to estimate the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio from a measurement of one (or both) of the stable isotope ratios. Panel (c) compares the isotope ratio correlation predicted from the fission yields with the result of a reactor model simulation intended to simulate the irradiation of the BR3 fuel that includes all neutron capture reactions. We use the sum of ⁸⁵Kr and its decay product ⁸⁵Rb to allow model-data comparison without correcting for ⁸⁵Kr decay (see text for detailed discussion). The model results are within 1% of the prediction from the fission yields for enriched uranium with a typical neutron spectrum, showing that neutron fluence effects on the initial ratio are negligible. Note that the mixing lines for pure fast neutron production are not shown in the right panel.

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The form of Equation 3 implicitly assumes that all ⁸⁵Kr is produced instantaneously and decays for the same length of time. Of course this is not the case in practice, because reactor fuel elements are typically irradiated for several years. This is a significant fraction of the ⁸⁵Kr half-life, so substantial decay of ⁸⁵Kr occurs simultaneously with production. Thus, the ⁸⁵Kr age defined in Equation 3 is strictly an 'apparent age' that does not correspond to a single point-like event (such as the introduction or removal of fuel from the reactor) but instead is only constrained to lie somewhere within the period of fuel irradiation. As discussed below, given the additional assumptions that the Kr production rate is constant, or varies in a known way, during the irradiation, and that all Kr produced is guantitatively retained, one can calculate the expected ⁸⁵Kr age corresponding to a particular irradiation history. In general, however, it is not possible to determine either the 105 start or end of an irradiation from the ⁸⁵Kr age alone. In practice, this limitation may not be a significant problem for many

applications, because typical irradiation durations for commercial reactor fuel are in the 3-5 year range, which is relatively short compared to the many decades since the initial development of nuclear power. Likewise, the application in which radiochronometry is used to group samples into collections of common origin and irradiation history does not rely on

110 associating a measured age with a specific event.

A final point is that, although ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) is relatively insensitive to neutron capture reactions during irradiation because the summation of ⁸³Kr and ⁸⁴Kr in the denominator removes fluence dependence related to the dominant neutron capture reaction on ⁸³Kr, the other Kr isotopes have neutron capture cross-sections that are nonzero (although two orders of magnitude less than ⁸³Kr; see Table 1), so there is likewise the possibility of a nonzero fluence dependence

- 115 for ⁸⁵Kr/(⁸³Kr+⁸⁴Kr). We investigated this using the results of the pin cell model for a pressurized water reactor described in Cassata et al. (2023), which tracks all fission, neutron capture, and decay reactions and therefore should expose any deviation in the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio from the expected fission yield ratios due to neutron interactions. We show the results of this comparison in Figure 2 (right panel), although we include ⁸⁵Rb, the decay product of ⁸⁵Kr, in the numerator as well so that decay of ⁸⁵Kr during the model simulation does not obscure any fluence effects. The
- 120 cross-section for neutron capture on ⁸⁵Rb is greater than that for ⁸⁵Kr, so this comparison likely yields an upper limit on the effect of fluence on the ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio. Regardless, deviations between the model simulation and a simple calculation of the expected initial ratio from the fission yields are less than 1% for the entire model simulation. In practice, the uncertainty in estimating ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) from ⁸⁶Kr/(⁸³Kr+⁸⁴Kr) or ¹³⁴Xe/(¹³¹Xe+¹³²Xe) will therefore most likely be dominated by measurement uncertainty on the isotope ratios, with uncertainty in the assumed neutron spectrum 125 providing a minor contribution. In the presumably unusual case where the reactor type from which a sample originated
- is completely unknown, uncertainty in the neutron energy spectrum could yield an additional \sim 4-5% uncertainty in the initial 85 Kr/(83 Kr+ 84 Kr) ratio estimated from one of the stable isotope ratios.

3 Sample acquisition and analysis

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We describe noble gas measurements on portions of fuel rods from two reactors. Typical commercial power reactor fuel
 consists of cylindrical UO₂ pellets with diameter ~1 cm that are stacked and encapsulated in airtight metallic cladding to form 0.5-4 meter-long rods. The neutron spectrum and flux within a reactor are spatially variable, so different portions of a fuel rod will evolve differently and therefore have somewhat different fission gas isotope compositions. As discussed in the introduction, we are interested in establishing whether the ⁸⁵Kr age of a fuel element is or is not position-independent, so the purpose of sample selection was to capture both axial and radial variability within fuel rods. In addition, the sample
 set includes both bulk analysis of complete slices of a fuel rod and analysis of individual microparticles excised from bulk fuel.

The first set of samples are slices taken at different axial positions along a fuel rod from Belgian Reactor No. 3 (henceforth, 'BR3'), a pressurized water reactor in Mol, Belgium. This rod had an initial ²³⁵U enrichment of 8.26% and underwent two periods of irradiation, between July 1976-April 1978 and June 1979-September 1980. This fuel sample has been characterized in detail for a variety of research purposes (e.g., Hanson and Pollington, 2021; Savina et al., 2021, 2023). Four

slices from different axial positions spanning the 1-meter length of the rod were completely dissolved in a sealed system,

and the gas released during dissolution was collected for noble gas separation and analysis. Analytical methods for these samples as well as stable Xe and Kr isotope compositions are described in detail in Cassata et al. (2023) and summarized in Table 2. Average burnup inferred from dissolution and actinide analysis of these samples was also reported by Cassata et al. and ranged from 19-48 GWd/ton. Most aspects of the fission gas isotope composition of these samples

were discussed in the Cassata reference: here we add discussion of the ⁸⁵Kr results.

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The second set of samples are microparticles isolated from a slice of a fuel rod that was irradiated in the Quad Cities Unit 1 reactor, a boiling water reactor located in Cordova, IL, USA. This fuel rod had an initial ²³⁵U enrichment of 3%, included 2 wt. % Gd₂O₃, and was irradiated between February 1979-September 1987 and then again from November 1989-September 1992. This irradiation is substantially longer than typical in commercial reactor operations. Portions of this fuel were subsequently made available for study as 'Approved Testing Material 109' (henceforth, 'ATM-109'), which has also been characterized for various research purposes (e.g., Wolf et al., 2005; Buck et al., 2015; Pellegrini et al., 2019; Clark et al., 2020). Analyses of various subsamples of this material indicated radially averaged burnup in the range 61-78 GWd/ton (Wolf et al., 2005). We obtained a cross-sectional slice of the ATM-109 material from near the axial center of the

- 155 fuel rod and further sectioned it into a matchstick-shaped segment (Figure 3). A focused ion beam (FIB) system was used to cut and separate roughly cube-shaped samples of fuel 1-4 microns in size ("cubes"; see Figs. 3,4) from various radial positions, which were then attached to molybdenum carrier plates for further handling. As spent nuclear fuel is highly radioactive, the purpose of this procedure is mainly to obtain samples that are small enough that they can be worked with in non-radiological, low-background analytical facilities; in addition, these samples are analogous to microparticles
- 160 that have been released to the environment from various nuclear facilities in the past (e.g., Steinhauser, 2018). More information on this sampling procedure is provided in Reilly et al. (2020). We cut and analyzed multiple replicate cubes at each radial position.

Xe and Kr were extracted from FIB "cubes" by heating under vacuum. We placed each molybdenum carrier inside a flattened segment ("packet") of Ta tubing that had previously been annealed under vacuum. The packets were placed

- 165 under vacuum and heated with a 150 W, 970 nm diode laser. The laser system is equipped with a coaxial optical pyrometer that permits measurement of the packet temperature during heating, and for a subset of the samples we calibrated the pyrometer to the emissivity of the packets, yielding accurate measurements of true temperatures during heating. After initial experiments on some samples to determine the approximate temperature necessary to release noble gases from the fuel matrix, for most samples we applied a heating schedule consisting of a 90-second preheating step at 600°C to
- 170 desorb any contaminant Xe and/or Kr derived from the atmosphere, followed by a series of higher-temperature heating steps starting at 1000°-1200° C and increasing in temperature until the amount of Xe released in each step began to decrease significantly. With the exception of one sample, we did not continue to increase the temperature to the point where zero gas was observed in the final step (one of our objectives was to determine whether or not the samples could be recovered for additional analyses, so we did not heat to temperatures higher than necessary). Thus, we can not verify
- 175 complete gas extraction for all samples. Based on the results for the one sample that was heated to exhaustion and the fact that we observed near-zero gas release in the final step for many samples, however, we estimate that >95% of



Figure 3. Scanning electron microscope image mosaic of slice of ATM-109 fuel pellet sampled for this work. The image shows a matchstick-shaped segment of a disc cut from the middle of a fuel rod, including both fuel and cladding. The surrounding material with a cracked texture visible at the edges of the image is epoxy used for mounting the sample, and part of the sample at upper left has been covered by copper tape. Colored dots show the locations from which samples were removed using a focused ion beam; 2-3 replicate cubes were cut from each location. The same colors are used to indicate radial position in subsequent figures. Light-colored grid lines have 1-millimeter spacing. The diameter of the fuel pellet is 1.06 cm. Images were collected using a FEI Helios NanoLab 660 FIB/SEM at 5 kV and 1.6 nA, and stitched using FEI MAPS 2.5 software.

fission gas present was most likely released from all samples. With the exception of the one sample that was heated to exhaustion at \sim 1600 °C, we did not exceed 1440 °C in any heating step. These temperatures are well below the melting temperature of uranium oxide (\sim 2800 °C), and microscopic examination of some samples after heating showed that they were intact and could potentially be used for further chemical analysis.

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Gas extracted by heating was exposed to hot and cold getters to remove any reactive gases, and Xe and Kr were separated from other noble gases by exposure to a helium-cooled stainless steel cold trap held at 60° K. The trap was then warmed to 220°K and Kr and Xe were let into a Nu Noblesse multicollector noble gas mass spectrometer. All stable isotopes of Xe and Kr, plus ⁸⁵Kr, were measured. Detector intercalibration and correction for mass discrimination were ac-185 complished by analysis of an atmospheric Xe-Kr standard that was measured at least twice daily. Background corrections to sample and standard analyses were based on full system "cold" blanks, measured in the same way as the samples except that the packets were not heated. After background correction, further corrections for any atmospheric Xe and/or Kr present were made by assuming that all ¹²⁹Xe and ⁸⁰Kr observed were atmospheric, and correcting other isotopes accordingly. Thus, data presented in Table 2 are isotope ratios for fissiogenic Xe and Kr after atmospheric correction. "Hot" blanks measured by heating empty Ta packets in the same apparatus released near-background amounts of Xe 190 and Kr with isotopic composition indistinguishable from atmosphere, so correction for any blank contribution from direct or collateral heating of the sample chamber is included in the atmospheric correction and is not considered separately. With one exception, we did not observe any significant variation in Xe or Kr isotope ratios among sequential heating steps for any sample. The exception has to do with a minority of samples that released anomalously large amounts of

195 gas in the 600° preheating step; the isotopic composition of this gas was distinguishable from that of the gas released in subsequent higher-temperature steps. Potentially, this is the result of early release of gas trapped in bubbles that are



Figure 4. Scanning electron microscope image of a sample of ATM-109 fuel prepared for noble gas analysis, collected using a FEI Inspect 4 SEM with the operating parameters shown in the image footer. The sample is a 4-micron-long rectangular solid ("cube") cut from the fuel using a focused ion beam (FIB) system (see Reilly et al., 2020). It is mounted to a molybdenum substrate for handling.

easily decrepitated by heating, but we have not investigated this in detail. Regardless, the observation that there is no significant isotope ratio variation during gas extraction after the preheat step indicates that diffusive fractionation of Xe and Kr isotopes is negligible at experimental heating temperatures, which in turn implies that the completeness of gas
extraction is not significant to the interpretation of isotope ratios. Thus, total amounts of all isotopes released in heating steps after the initial preheat step were summed to calculate summary isotope ratios for gas released from each sample. These results are shown in Table 2 and the complete step-degassing data are in the supplement.

4 Results and discussion

Figure 1 shows stable Kr and Xe isotope signatures that are diagnostic of fission source. As discussed in Cassata et al.
(2023), Kr-Xe data for the BR3 bulk samples indicate that fission gas production is dominated by ²³⁵U fission, as expected for a moderately-enriched fuel that experienced only moderate burnup. Stable isotope ratios for the ATM-109 fuel, on the other hand, indicate a much larger contribution from ²³⁹Pu fission, as expected for a lower-enriched fuel at high burnup that is likely nearly entirely depleted in ²³⁵U. In addition, the ATM-109 data display a strong edge effect such that samples from closer to the edge have a fission gas composition closer to the ²³⁹Pu end member. Again, this is expected from the fact that neutron capture on ²³⁸U and therefore ²³⁹Pu production is highest at the edge of fuel pellets. Overall, the

important observation from these data is that the particulate samples span a much larger range of fission sources and therefore fission gas compositions than the bulk fuel samples.



Figure 5. Graphical representation of the apparent ⁸⁵Kr age of spent fuel samples. The gray line at the top is the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio at production for thermal-neutron-induced fission of mixtures of ²³⁵U and ²³⁹Pu, as shown in Figure 2. As discussed above, the initial ratio is weakly variable among fission source but can be inferred from its correlation with the stable isotope ratios ⁸⁶Kr/(⁸³Kr+⁸⁴Kr) (panel a) or 134 Xe/(131 Xe+ 132 Xe) (panel b). Decay of 85 Kr after production then causes the 85 Kr/(83 Kr+ 84 Kr) ratio to decrease over time, as indicated by the colored isochrons labeled in years. This allows an apparent age of the fission gas to be calculated. Error bars on data are 1σ and, where not visible, are smaller than the size of the plotting symbols. Comparison of these plots shows that, for the particulate samples from ATM-109, relatively large uncertainties in the ⁸⁶Kr/(⁸³Kr+⁸⁴Kr) ratio measured on small amounts of gas mean that the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) can be estimated more accurately from the Xe data. Note that there is a 3-year gap between ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) measurements for the two sample sets (Table 3), but in this figure they are plotted as measured at the time of analysis and have not been normalized to a common analysis date.

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Figure 5 shows the relationship of the ⁸⁵Kr data to the stable Kr and Xe isotope ratios indicative of fission source. Across the full range of stable isotope compositions, the differences between measured ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratios and production ratios expected from fission yields are consistent with 34-43 years of decay following irradiation. Apparent ages of the BR3 fuel show minimal axial variation among samples. As observed in the stable isotope ratios, apparent ages of the ATM-109 particle samples show significant radial variability, with samples near the edge having younger apparent ages. It is evident from both Figures 1 and 5 that for the ATM-109 particle samples, Kr isotope ratios have significantly larger uncertainties and scatter than Xe isotope ratios; this is primarily due to low abundance of Kr. For mass-spectrometric measurement of 220 these samples, signals for Kr isotopes on ion counters were in the range 10-100 cps, whereas Xe isotope signals were an order of magnitude higher. For the bulk dissolution samples, a large quantity of gas was available and Kr measurement precision was not signal-limited.



Figure 6. Apparent ⁸⁵Kr dates for bulk fuel samples from BR3 (a) and microparticle samples from ATM-109 (b). The prominent colored symbols are the dates calculated using the ¹³⁴Xe/(¹³¹Xe+¹³²Xe) ratio to estimate the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio, and the lighter unshaded symbols behind them are calculated using the ⁸⁶Kr/(⁸³Kr+⁸⁴Kr) ratio as the estimator. The two approaches yield indistinguishable results. The shaded regions are the periods of time during which the respective fuels were irradiated, and the dashed lines are the apparent ⁸⁵Kr dates expected given simple assumptions that ⁸⁵Kr production is constant throughout the irradiation and all gas is retained. The color-coding of the symbols is the same as in previous figures.

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Table 3 shows apparent ⁸⁵Kr dates calculated from Kr and Xe isotope ratios, and Figure 6 compares them with the known irradiation dates of the two fuels. Apparent ⁸⁵Kr ages in this figure and in Table 3 have been calculated using Equations 1-3, in which we assume that initial ⁸⁵Kr(⁸³Kr+⁸⁴Kr) ratios are characteristic of thermal-neutron-induced fission as shown in Figs. 2 and 5, and use this mixing curve with the measured ⁸⁶Kr/(⁸³Kr+⁸⁴Kr) or ¹³⁴Xe/(¹³¹Xe+¹³²Xe) ratios to estimate the initial ratio applicable to each sample. We now discuss several aspects of these results.

First, we discuss internal variation in apparent ⁸⁵Kr dates for both fuels. Dates for the BR3 bulk fuel samples are indistinguishable across a range of axial positions. On the other hand, ⁸⁵Kr dates on the ATM-109 particle samples show

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both (i) variance in excess of measurement uncertainty among replicate samples taken from the same radial position, and (ii) an edge effect in which ⁸⁵Kr dates from samples within \sim 100 μ m of the fuel pellet edge are significantly younger than those from the pellet center. We hypothesize that the scatter among replicates is likely due to partitioning of fission gas between fuel matrix and bubbles due to gas mobility during irradiation. Xe and Kr are known to migrate into micronor sub-micron-scale bubbles during fuel irradiation (see, e.g., Rest et al., 2019) and therefore, on average, gas in bubbles 235 must have an older production age than gas in the matrix. This would imply micron-scale inhomogeneity in gas age, which appears to have been captured by our sampling procedure.

The edge effect is nearly certainly due to gas loss associated with so-called "high-burnup structure" (e.g., Rondinella and Wiss, 2010) at fuel pellet edges. High-burnup structure forms by recrystallization of fuel subjected to high local concentrations of fissions, primarily at the outer surface of fuel pellets. Various microbeam studies of spent fuel (Walker

et al., 2012) have observed lower concentrations of fission gases within regions of high-burnup structure, implying that recrystallization of the fuel matrix either directly rejects noble gases or temporarily or permanently enhances their diffusivity. If the development of high-burnup structure late in the irradiation results in the loss of gas produced early in the irradiation, then only gas produced late in the irradiation remains in the sample for measurement, and this gas will have a younger ⁸⁵Kr age than the gas that was lost. Thus, a sample that experienced formation of high-burnup structure late in the irradiation. The observation of younger apparent ⁸⁵Kr ages in edge samples where high-burnup structure is expected is

exactly consistent with this mechanism.



Figure 7. Radial variation in fluence-sensitive ⁸⁴Kr/⁸³Kr (a) and ¹³²Xe/¹³¹Xe (b) ratios in ATM-109 fuel. Color-coding of symbols is the same as in previous figures.

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The hypothesis that the edge effect in apparent ⁸⁵Kr ages in the ATM-109 fuel is due to gas loss is consistent with other stable isotope evidence. Specifically, again, by 'gas loss' we refer to preferential loss of gas produced early in the irradiation, either by a continuous process active throughout the irradiation (e.g., thermally activated diffusion) or a distinct event (e.g., formation of high-burnup structure). As diffusive fractionation of Xe and Kr isotopes in these samples appears to be negligible (see discussion in section 3 above), we have no reason to believe that the process of gas loss itself affects the isotope composition of the gas that is being lost; instead, isotope ratios are different between samples that have and have not experienced gas loss because the gas that is still present and available for measurement was produced during

- 255 different parts of the irradiation. For example, Figure 7 shows radial variation in the ratios ⁸⁴Kr/⁸³Kr and ¹³²Xe/¹³¹Xe. Because of the high neutron capture cross-sections of ⁸³Kr and ¹³¹Xe, these ratios are diagnostic of total neutron fluence experienced by the fuel and should therefore increase monotonically with irradiation time (Cassata et al., 2023). However, measured ratios are much lower near the fuel pellet edge. As the edge and center samples themselves have experienced a the same irradiation time, the only explanation for this is that the fission gas present in edge samples has experienced a shorter irradiation time than that in the pellet center. This is only possible if a significant fraction of gas produced early in
- 260 shorter irradiation time than that in the pellet cen the irradiation has been lost from edge samples.

Second. we compare observed to expected ⁸⁵Kr ages. Because both ⁸⁵Kr production and decay are taking place throughout the period of irradiation, an apparent ⁸⁵Kr gas age is expected to date some time in the middle of the irradiation. In Figure 6 we calculate an expected age for each fuel from the known irradiation dates by making the simplifying assumptions that the Kr production rate is constant throughout the irradiation and all gas is retained. The results differ 265 from this expectation in several ways. With the exception of the edge samples that are biased young by gas loss as discussed above, the ATM-109 particles have apparent ages older than expected. This is relatively easy to explain by the observation that Kr fission yields from ²³⁹Pu are 40-60% lower than from ²³⁵U. For this unusually-high-burnup fuel. the Kr production rate has therefore nearly certainly decreased over time as ²³⁵U is depleted and Kr production late in the irradiation becomes dominated by ²³⁹Pu fission. If the Kr production rate has decreased throughout the course of the 270 irradiation and all gas is retained, the apparent ⁸⁵Kr age will be older than the age expected from a constant-production assumption. This appears to be the case for particle samples from the center of the ATM-109 fuel pellet. Overall, for the ATM-109 particle data. offsets from the ⁸⁵Kr age expected from simple assumptions are significant with respect to measurement uncertainties and spatially variable, but can be satisfactorily explained by steadily decreasing Kr production 275 during a long irradiation as well as gas loss during formation of high-burnup structure at the pellet edge.

On the other hand, apparent ⁸⁵Kr dates for the BR3 fuel are younger than expected for constant production, and close to the end of the irradiation. This is unexpected for this 8%-enriched, moderate-burnup fuel, for which the assumption of constant production during the irradiation should be much closer to correct than for the 3%-enriched, high-burnup ATM-109 fuel. Model simulations of the BR3 fuel (Cassata et al., 2023, see also Fig. 2) predict that its apparent age at discharge should be within 0.05-0.1 yr of the expected age computed from a constant-production assumption. It is likely that a nonzero amount of fission gas was lost from the fuel to the rod plenum during irradiation, and this gas was not captured in the analysis of these samples, so a young bias to the age due to gas loss is possible. However, typical losses of 10-15% are insufficient to explain the magnitude of the observed young bias in apparent ages. In addition, gas loss would be expected to be highest at the axial center of the rod where fuel temperatures are the highest, so if gas loss was a good explanation for the young bias, we would see younger apparent ages in the axial center of the rod. In fact, we see a slightly older (although indistinguishable at uncertainty) age from the axial center sample. Overall, although the apparent ages from the BR3 samples fall within the known period of irradiation, we do not have a satisfying explanation

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for why they are systematically 1-2 years younger than expected.

Third, we discuss the implications of age variations for grouping of samples. Although it is evident from the BR3 results that the apparent ⁸⁵Kr age is position-independent at the bulk sample level, the ATM-109 microparticle results show that this is not strictly the case at the level of individual particles. All apparent ages from ATM-109 samples fall within the period during which the fuel was irradiated, but variations in apparent age in excess of measurement uncertainty are evident both across the radial array of samples and in replicate data from the same radial position. Although the long irradiation and high burnup of the ATM-109 fuel likely give rise to more internal variability than would be expected for a fuel with a more typical history, strictly, these samples could not be associated with each other purely on the basis of the apparent ⁸⁵Kr age alone. However, the apparent ⁸⁵Kr age is not determined in isolation, but as part of a larger set of Xe and Kr isotope ratio measurements that provides additional leverage in grouping samples. It is clear from, for example,

³⁰⁰ Finally, we highlight several aspects of uncertainty analysis. Given the assumption (discussed above) that the neutron spectrum is typical of a normal commercial power reactor, nominal measurement uncertainty in the apparent ⁸⁵Kr age is dominated by measurement uncertainty on ⁸⁵Kr and to a lesser extent on other Kr isotopes, especially for signal-limited measurements from small particles. These nominal uncertainties (Table 2) are 0.2-0.6 yr for non-signal-limited measurements from the BR3 dissolution gas and 0.5-3 yr for lower-abundance measurements on the ATM-109 microparticles. As

Figure 5, that even though the ATM-109 samples do not all have exactly the same ⁸⁵Kr age, when multiple isotope ratios

- 305 noted above, an additional uncertainty on the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio would arise if the reactor type and therefore the neutron spectrum responsible for fission were completely unknown. However, it is evident from Figure 6 that the most important uncertainty in interpreting apparent ⁸⁵Kr ages arises not from measurement uncertainties but from deviations between the observed ages and the ages that we would expect to observe given the known irradiation histories of the samples. These are larger than measurement uncertainty and are related to irradiation history and duration, fuel enrich-
- 310 ment and burnup, and, at least in the high-burnup ATM-109 samples, enhanced fission gas loss near fuel pellet edges. These complications could be important if, for example, seeking to use ⁸⁵Kr age data to associate fuel samples with one of several candidate irradiation periods, or to infer the irradiation duration of fuel samples with a known discharge date. Overall, systematic variations in apparent ⁸⁵Kr age that depend on irradiation conditions are a more important limit to the interpretability of these data than the measurement uncertainties.

315 5 ⁸⁵Kr compared to other spent fuel chronometers

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are considered they belong to a coherent array.

Other radiochronometers have been proposed and/or used for spent fuel samples. Their relative applicability and usefulness compared to ⁸⁵Kr chronometry are likely to vary with fuel enrichment and burnup, type and amount of sample, and available measurement technology. However, here we highlight some general differences between ⁸⁵Kr and two other proposed chronometers that have been applied to small particle samples of spent fuel: ⁹⁰Sr (Savina et al., 2023) and ²⁴¹Pu (Hanson and Pollington, 2021).

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⁹⁰Sr is a fission product with a half-life of 28.2 years that can be used as a chronometer by comparison with the stable fission product ⁸⁸Sr. As for the ratio ⁸⁵Kr/(⁸³Kr+⁸⁴Kr), the initial ⁹⁰Sr/⁸⁸Sr ratio depends on the proportion of fissions derived from the various actinide isotope and neutron energies. With ⁸⁵Kr, we exploit variations in stable Kr and Xe isotopes that are produced simultaneously to infer the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio, and model simulations show that the initial ratio estimate has percent-level uncertainty. For ⁹⁰Sr, the ⁹⁰Sr/⁸⁸Sr ratio only varies at the percent level across the full range of fission sources, so can be estimated with comparable precision. Savina et al. (2023) used a resonance ionization mass spectrometer (RIMS) to measure ⁹⁰Sr/⁸⁸Sr ages on 10-micron microparticles derived from a sample of the BR3 fuel, and obtained apparent ages very close to the age expected from constant production (1-2 years older than our apparent ⁸⁵Kr ages for bulk samples of this fuel). Important differences between the ⁸⁵Kr and ⁹⁰Sr chronometers are as follows:

- ⁸⁵Kr has a lower fission yield than ⁹⁰Sr and also can be lost by diffusion, so the ⁹⁰Sr concentration in spent fuel is expected to be higher. Thus, measurement precision on ⁹⁰Sr may be better for small samples.
- The production rate of both ⁹⁰Sr and ⁸⁵Kr is expected to be relatively constant throughout an irradiation in most cases. Thus, apparent ages for both should fall near the middle of the irradiation period.
- Sr is expected to be far less diffusively mobile than Kr, which implies that variability in apparent ages at the microparticle scale that are caused by Kr mobility should not be evident in ⁹⁰Sr ages.
 - Measurement of Kr isotopes utilizes the relatively simple and widely used analytical method of heating under vacuum followed by noble gas mass spectrometry, whereas measurement of Sr isotopes requires either sample dissolution and chemical separation, which is complex and time-consuming, or measurement by RIMS, which is not widely available.

²⁴¹Pu is produced by neutron capture during irradiation and decays to ²⁴¹Am with a half-life of 14.35 years. Thus, the ²⁴¹Am/²⁴¹Pu ratio in spent fuel provides an age estimate for the fuel. This chronometer was, for example, applied to 10to 20-micron size samples of the BR3 fuel by Hanson and Pollington (2021). By comparison with ⁸⁵Kr chronometry,

- Due to neutron capture on ²⁴¹Am, the ²⁴¹Am/²⁴¹Pu ratio at discharge is lower than expected from ²⁴¹Pu production alone, and varies with fluence. Thus, to achieve comparable accuracy to the ⁸⁵Kr or ⁹⁰Sr chronometers, it would most likely be necessary to apply an independent fluence signature to account for ²⁴¹Am loss during irradiation.
 - Production of ²⁴¹Pu requires three successive neutron capture reactions, so is weighted toward the end of an irradiation. Thus, a ²⁴¹Pu age is expected to date a later time in an irradiation than an ⁸⁵Kr or ⁹⁰Sr age.
- Measurement of the ²⁴¹Am/²⁴¹Pu ratio requires sample dissolution and chemical separation of Am from Pu, followed by two separate isotope dilution measurements. Thus, it is much more complex and time-consuming than measurement of Kr isotope ratios by vacuum heating/NGMS or measurement of Sr isotope ratios by RIMS.

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- Am and Pu are not expected to be diffusively mobile, so in principle ²⁴¹Am/²⁴¹Pu ages should display less internal variability than ⁸⁵Kr ages. However, in the one application of this method to small particle samples, Hanson and Pollington (2021) observed relatively high and apparently nonsystematic internal variability, well in excess of nominal measurement uncertainty.

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6 Conclusions

The date of irradiation of spent reactor fuel can be determined by measuring the ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio of gas extracted from spent fuel samples, and using simultaneous measurements of other Kr and Xe isotope ratios to estimate the initial ⁸⁵Kr/(⁸³Kr+⁸⁴Kr) ratio at the time of fission gas production. This method has several possible advantages in relation to other radiochronometric methods, in particular that it can be applied to micron-size samples of spent fuel and only 360 requires a single simultaneous measurement of Kr and Xe isotopes, which is relatively rapid and utilizes fairly simple and widely available vacuum heating and noble gas mass spectrometry systems. When applied to microparticles of spent U oxide fuel, the method is nondestructive in the sense that the samples are not melted or otherwise damaged and could be subjected to further chemical analysis, although it is possible that some other volatile elements besides the noble gases are depleted during heating.

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⁸⁵Kr chronometry applied to two sets of samples of spent reactor fuel yields apparent ages that, as expected, fall entirely within known periods of irradiation. From this perspective, the method is successful. When considered in more detail, we observed both systematic (e.g., from edge to center of a fuel pellet) and random (e.g., variability among replicates) variation in apparent age among microparticle samples, as well as differences between measured ⁸⁵Kr ages and expected ages near irradiation midpoints for both bulk and particulate samples. We attribute these observations to aspects of the fuel history (e.g., irradiation duration and burnup) as well as mobility and loss of noble gases within uranium oxide fuel.

From the perspective of using the ⁸⁵Kr chronometer to group samples of unknown origin, indistinguishable ⁸⁵Kr ages from bulk dissolutions of fuel slices at different axial positions show that the ⁸⁵Kr age is a position-independent signature

- at the bulk sample level. On the other hand, observed variations in apparent ⁸⁵Kr ages among microparticles from the 375 same fuel sample show that this is not strictly the case at the microparticle scale. Especially for high-burnup fuel, internal variation in ⁸⁵Kr ages at the particle scale is likely. However, both our observations and our explanation of observed internal variability indicate that the magnitude of internal variability (i) cannot exceed the duration of fuel irradiation, and (ii) is most likely maximized in high-burnup samples. Thus, for typical fuel irradiation durations of 1-2 years, internal
- variability in ⁸⁵Kr ages on particle samples will likely not exceed measurement uncertainty, and therefore will not preclude 380 accurate grouping of samples or association of samples with candidate irradiation periods. In addition, when both the ⁸⁵Kr age and other simultaneously measured stable Kr and/or Xe isotope ratios are considered, data from particles of common origin will form coherent arrays. Even in the presence of some internal variability in ⁸⁵Kr ages, sample grouping can be improved by considering the 85 Kr age as part of a set of multiple noble gas isotope ratios.

Author contributions. Conceptualization: GB, AJC, DDR, DB, BHI; Methodology: GB, AJC, DDR, DB; Software: GB, AJC; Validation: GB, AJC; Formal analysis: GB, AJC, CDW; Investigation: GB, CDW; Resources: GB, AJC, CDW; Writing (original draft): GB, CDW; Writing (review and editing): all authors; Visualization: GB, DDR, DB; Project administration: GB, BHI; Funding acquisition: AJC, BHI.

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						Neutron capture x-section s(n,g) (barns)				
	$^{235}U_{therm}$	²³⁵ U _{fast}	²³⁸ U _{fast}	²³⁹ Pu _{therm}	²³⁹ Pu _{fast}	Thermal	Resonance	Fast		
	Cumulative fis	ssion yields (%)							
⁸² Se						<0.1	0.2	<0.1		
⁸³ Kr	0.536	0.577	0.396	0.297	0.315	198	188	<0.1		
⁸⁴ Kr	1.002	1.031	0.826	0.480	0.453	0.1	2.1	<0.1		
⁸⁵ Kr	0.283	0.275	0.149	0.123	0.110	1.7	2.7	<0.1		
⁸⁶ Kr	1.965	1.947	1.296	0.766	0.787	<0.1	<0.1	<0.1		
	Cumulative fis	ssion yield ratio	os							
⁸⁵ Kr/ ⁸⁶ Kr	0.144	0.141	0.115	0.160	0.139					
85Kr/(83Kr+84Kr)	0.184	0.171	0.122	0.158	0.143					
⁸⁵ Kr/(⁸³ Kr+ ⁸⁴ Kr+ ⁸⁶ Kr)	0.081	0.077	0.059	0.080	0.071					

Table 1: Cumulative fission yields, cumulative fission yield ratios, and neutron capture cross-sections for relevant Kr isotopes. Data are from ENDF/B-VIII.

Table 2. Isotopic composition of fissiogenic Xe and Kr in spent fuel samples.

		Relative Kr abundances (Kr-86 = 100)						Relative Xe abundances (Xe-134 = 100)					
Sample name		Kr-83	+/-	Kr-84	+/-	Kr-85	+/-	Xe-131	+/-	Xe-132	+/-	Xe-136	+/-
ATM-109 microparticle	e samples												
	Radial distance from pellet edge (microns)												
ATM-109-20-54	22	26.8	1.3	65.1	5.8	1.694	0.166	35.87	0.20	76.95	0.35	157.05	0.61
ATM-109-20-55	22	27.7	2.9	65.4	12.5	1.656	0.285	35.20	0.33	77.81	0.51	156.76	0.81
ATM-109-60-54	70	21.0	1.2	68.3	5.3	1.408	0.155	31.02	0.18	80.57	0.34	156.99	0.56
ATM-109-60-55	70	19.8	1.4	67.7	6.4	1.264	0.168	29.62	0.20	81.65	0.38	156.13	0.64
ATM-109-120-01	122	17.6	1.0	64.6	4.0	1.189	0.063	25.69	0.17	83.54	0.36	155.55	0.81
ATM-109-120-02	122	18.93	0.55	73.4	2.2	1.107	0.050	25.63	0.14	83.38	0.32	154.96	0.88
ATM-109-400-01	397	17.07	0.51	70.1	2.0	1.058	0.043	21.83	0.12	84.47	0.31	152.95	0.80
ATM-109-400-02	397	17.33	0.29	73.2	1.0	1.004	0.043	22.71	0.11	83.74	0.31	152.56	0.86
ATM-109-800-01	790	15.02	0.20	71.0	0.8	1.102	0.025	20.100	0.076	84.38	0.23	151.13	0.67
ATM-109-800-02	790	17.32	0.27	74.1	0.6	1.103	0.037	21.55	0.12	83.59	0.35	151.91	0.99
ATM-109-1600-01	1562	16.78	0.40	71.4	1.5	1.240	0.052	21.75	0.11	83.09	0.30	152.01	0.81
ATM-109-1600-02	1572	15.42	0.21	71.90	0.61	1.083	0.032	20.59	0.10	83.32	0.31	151.73	0.94
ATM-109-3200-01	3115	15.65	0.52	67.7	2.2	1.132	0.043	21.77	0.12	81.80	0.29	151.67	0.80
ATM-109_3-1	3905	15.38	0.26	66.41	0.90	1.195	0.055	21.321	0.069	81.65	0.19	151.30	0.43
ATM-109_3-2	3910	15.50	0.38	65.5	1.4	1.428	0.069	21.344	0.094	81.74	0.26	151.25	0.65
ATM-109_3-3	3915	16.34	0.37	66.8	1.5	1.083	0.035	21.686	0.082	81.21	0.20	151.58	0.53
BR-3 total dissolution	s (reproduced from Cass	ata et al., 202	23)										
	Axial distance from rod bottom (cm)												
G14	95	25.98	0.39	52.34	0.90	1.023	0.012	34.359	0.019	58.799	0.057	115.492	0.082
G16	5	25.62	0.31	53.38	0.40	1.026	0.041	34.84	0.12	60.50	0.12	118.33	0.31
G18	50	23.32	0.16	56.17	0.45	1.001	0.017	29.774	0.092	64.64	0.16	138.95	0.23
G21	21	24.36	0.11	54.79	0.16	1.057	0.011	30.378	0.042	62.486	0.071	134.88	0.26

Notes: Relative abundances shown for ATM-109 particles reflect the sum of gas released in all heating steps after the initial preheat.

Table 3. ⁸⁵Kr age calculations.

		Measured		Measured		Estimated initial		⁸⁵ Kr age		⁸⁵ Kr age	
Sample name	Analysis date	85Kr/(83Kr+84Kr)	+/-	134Xe/(131Xe+132Xe)	+/-	⁸⁵ Kr/(⁸³ Kr+ ⁸⁴ Kr)	+/-	(yr before analysis)	+/- (yr)	(calendar date)	+/- (days)
ATM-109 microparticle sa	imples										
ATM-109-20-54 ATM-109-60-54 ATM-109-60-54 ATM-109-60-54 ATM-109-120-01 ATM-109-120-02 ATM-109-400-02 ATM-109-400-01 ATM-109-400-02 ATM-109-400-01 ATM-109-600-01 ATM-109-1600-01 ATM-109-1600-02 ATM-109-320-01 ATM-109-321	4/19/23 4/19/23 4/19/23 8/11/23 8/11/23 8/11/23 8/13/23 8/14/23 8/15/23 8/16/23 8/16/23 8/18/23 11/6/21	0.0184 0.0178 0.0158 0.0145 0.0145 0.01214 0.01214 0.01201 0.01221 0.01206 0.01207 0.01240 0.01388 0.01462	0.0021 0.0038 0.0019 0.0022 0.0010 0.00060 0.00055 0.00048 0.00039 0.00039 0.00039 0.00062 0.00039 0.00062	0.8864 0.8845 0.8862 0.9155 0.9175 0.9407 0.9394 0.9571 0.9511 0.9539 0.9624 0.9656 0.9711	0.0036 0.0053 0.0034 0.0039 0.0038 0.0035 0.0034 0.0034 0.0034 0.0039 0.0033 0.0035 0.0035 0.0035	0.1689 0.1687 0.1703 0.1703 0.1727 0.1729 0.1752 0.1751 0.1766 0.1761 0.1764 0.1770 0.1773 0.1777	0.0018 0.0019 0.0018 0.0018 0.0018 0.0018 0.0018 0.0018 0.0018 0.0018 0.0018 0.0018 0.0018 0.0018 0.0018	34.3 36.9 38.3 38.4 41.34 41.37 42.74 40.65 41.54 39.18 41.19 39.81 39.70	1.8 3.3 1.9 2.3 1.1 0.79 0.73 0.69 0.53 0.70 0.47 0.71 0.74	12/23/88 6/11/88 6/4/86 1/15/85 3/10/85 4/9/82 3/31/82 11/13/80 12/18/82 6/10/84 6/7/82 10/27/83 2/23/83	650 1200 700 850 400 290 250 140 190 260 170 260 270
ATM-109_3-2 ATM-109_3-3	12/22/21 1/21/22	0.01763 0.01303	0.00090 0.00046	0.9701 0.9718	0.0030	0.1776 0.1777	0.0018 0.0018	35.79 40.48	0.80 0.57	3/9/86 7/27/80	290 210
BR3 total dissolutions											
G14 G16 G18 G21	8/12/20 7/27/20 8/4/20 8/6/20	0.01306 0.01299 0.01260 0.01335	0.00022 0.00053 0.00022 0.00014	1.07343 1.0489 1.0591 1.0768	0.00069 0.0019 0.0020 0.0010	0.1832 0.1820 0.1825 0.1834	0.0018 0.0018 0.0018 0.0018	40.92 40.90 41.42 40.59	0.31 0.65 0.31 0.23	9/12/79 9/2/79 3/6/79 1/4/80	110 240 110 83

Notes: Only initial ⁸⁵Kr/(⁶⁵Kr+⁸⁵Kr) ratios and ages calculated from stable Xe isotope ratios are shown. Corresponding calculations based on stable Kr isotopes are not included.

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